Hunting misaligned black hole accretion disks



K.Chatterjee¹, D. Yoon¹, S.B. Markoff¹, Z. Younsi², M.T.P. Liska³, A. Tchekhovskoy⁴, D. van Eijnatten¹, C. Hesp¹, A. Ingram⁵, M.B.M. van der Klis¹

- 1: Anton Pannekoek Institute, UvA,
- 2: University College London,
- 3: Harvard University,
- 4: CIERA, Northwestern University,
- 5: University of Oxford;

Contact: k.chatterjee@uva.nl

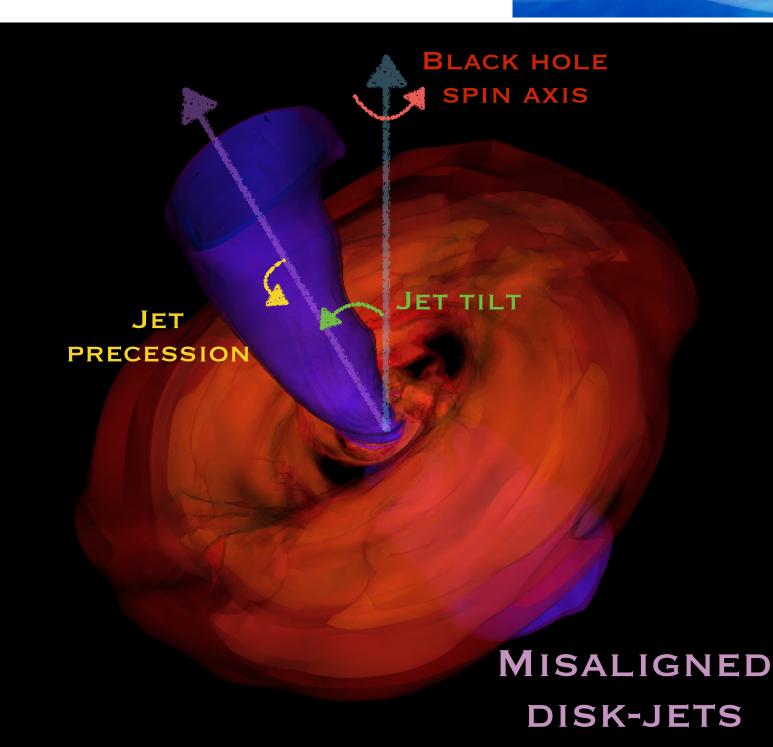
Black holes and their surroundings

Black holes are perhaps the most diverse laboratories of physics due to the interplay between the strong gravity, matter and electromagnetic fields. Since the in-falling gas remains blissfully unaware of the black hole spin axis till very small distances, misalignment between the accretion disk and the black hole spin axis is thought to be common. However, is it possible to tell observationally if a black hole disk is tilted?

H-AMR: the world's first GPU accelerated GRMHD code: some prominent results

- First simulation to show precessing jets from misaligned disks [1].
- Largest 2D simulation to date explains spine-sheath jet structure [2].

Imaging misaligned disks



Event Horizon Telescope

GRMHD simulation parameters: resolution 448x144x240 (in spherical

