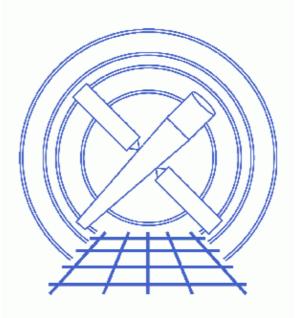
Examining Grating Spectra and Regions: PHA2 files



CIAO 3.4 Science Threads

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Examining Grating Spectra and Regions: PHA2 files

CIAO 3.4 Science Threads

Overview

Last Update: 1 Dec 2006 - updated for CIAO 3.4: ChIPS and Sherpa versions

Synopsis:

An overview of displaying grating data Type II PHA files and the source and background extraction regions.

Related Links:

• <u>Analysis Guide for Chandra High Resolution Spectroscopy</u>: an in-depth discussion of grating analysis.

Proceed to the <u>HTML</u> or hardcopy (PDF: <u>A4 / letter</u>) version of the thread.

What is a PHA2 File?

A Type II PHA file is a standard FITS format in which each row contains array columns. The PHA file is a product of standard data processing and is identified by the pha2.fits extension; note that the "2" in the filename refers to the fact that it is a level=2 data product, *not* that it is a Type II file. In the case that the user has to manually reprocess an event file (e.g. when applying an updated order sorting table), the PHA2 spectrum file is obtained from the level 2 event file by tgextract; see the Obtain Grating Spectra from HETG/ACIS-S Data for an example of this.

The SPECTRUM block of a PHA2 file has 13 columns of data:

ColNo	Name	Description
1	SPEC_NUM	Spectrum Number
2	TG_M	Diffraction order (m)
3	TG_PART	Spectral component (HEG, MEG, LEG, HESF parts)
4	TG_SRCID	Source ID, output by tgdetect
5	Х	X sky coord of source
б	Y	Y sky coord of source
7	CHANNEL[8192]	Vector of spectral bin numbers.
8	COUNTS[8192]	Counts array (a spectrum)
9	STAT_ERR[8192]	Statistical uncertainty (error) on counts colum
10	BACKGROUND_UP[8192]	Upper Background count vector.
11	BACKGROUND_DOWN[8192]	Lower Background count vector.
12	BIN_LO[8192]	Bin boundary, left edge
13	BIN_HI[8192]	Bin boundary, right edge

There are two columns that are especially relevant when doing analysis:

- TG_M indicates the order of the spectrum (+/-1, +/-2, +/-3)
- TG_PART indicates the spectral component / grating arm (1 = HEG, 2 = MEG, 3 = LEG)

Examining the Files with Prism

Sample ObsIDs used: 459 (HETG/ACIS–S 3C 273); 460 (LETG/HRC–S, 3C 273); 1198 (LETG/ACIS–S, 3C 273); 1800 (LETG/HRC–I, PKS2155–304)

File types needed: evt2; pha2

ACIS-S HETG/LETG Observations

We can use prism to examine the PHA2 file for ObsID 459:

unix% prism acisf00459N002_pha2.fits &

which will give you something like <u>Figure 1</u> a. In this example, there are twelve rows – all the +/– orders for both HEG and MEG – for the observation. The columns CHANNEL, COUNTS, BIN_LO, etc. are all so–called "vector columns"; each contains a vector of elements which, in this example, is 8192 elements long.

An ACIS–S/LETG observation (ObsID 1198) looks similar in *Prism* to but only contains 6 rows (+/- orders for the LEG).

Each vector column can be viewed as follows:

unix% prism acisf00459N002_pha2.fits &

left-click on column of interest (to select column) Navigate menu -> Expand Column (to expand column)

Figure 3 the shows the (partial) result of the expansion of column BIN_LO.

HRC-S/HRC-I LETG Observations

Examining an HRC–S/LETG observation (ObsID 460) is done in the same way as an ACIS grating observation:

unix% prism hrcf00460N002_pha2.fits &

but there is an important difference in the results. As seen in the <u>Prism display</u>, there are only two rows for the LEG observation. HRC–S cannot resolve orders and the COUNTS in the +/– 1 order are in fact the *total counts* of *all orders combined*. Also, the BIN_LO and BIN_HI columns should be considered for reference only; they actually represent the boundary wavelength of the +/– 1 order alone, while photons from all orders are included in the spectra.

The same holds true for HRC–I/LETG observations, as seen Figure 5 to in the example of ObsID 1800.

Displaying the Spectrum

With Prism

The quickest way to display one of the spectra of a PHA2 file is with *Prism*. For example a user who wants to take a quick look at BIN_LO vs. COUNTS for the order +1 (TG_M=1) of the HEG spectrum (TG_PART=1) should do the following (ObsId 459):

```
unix% prism acisf00459N002_pha2.fits &
middle-click on row #4 (to select spectrum)
left-click on BIN_LO column (to select X-axis column)
left-click on COUNTS column (to select Y-axis column)
Vizualization menu -> Interactive plot (to view the plot)
Vizualization menu -> Print plot (to print the plot)
```

As clearly visible from the <u>output in Figure 6</u>, the plot is good for a general overview only (are there enough counts, are there striking features, etc.). At the same time as the plot was created, a *ChIPS* session was launched by *Prism* in an xterm window. This window can be used to customize the plot; see the <u>Introduction</u> to *ChIPS* thread for more information on plotting with *ChIPS*.

With ChIPS

In order to display a spectrum with <u>*ChIPS*</u>, we need to separate it out from the PHA2 file (keep in mind that each row of the PHA2 file corresponds to one spectrum):

unix% <u>dmtype2split</u> "acisf00459N002_pha2.fits[#row=4]" acisf00459N002_heg_p1.pha

The newly created file, which is the +1 order HEG spectrum for this observation, can be used as input to *ChIPS*:

```
unix% chips
Welcome to ChIPS, version CIAO 3.4
Copyright (C) 1999-2003, Smithsonian Astrophysical Observatory
chips> plot "acisf00459N002_heg_p1.pha[cols bin_lo,counts]"
chips> symbol none
chips> step
chips> step
chips> xlabel "Wavelength (\AA)"
chips> ylabel Counts
chips> title "ACIS+HEG order=+1"
```

These commands produce the plot shown in Figure 7

An alternative way to read and display a PHA2 spectrum is through <u>S-Lang variables</u>. In this case the PHA2 does not need to be split first. To produce the plot above using S-Lang syntax:

```
chips> clear
chips> pha2 = readbintab("acisf00459N002_pha2.fits")
chips> print(pha2)
_filename = acisf00459N002_pha2.fits
_path
              = /data/ciao/threads/
_filter
             = NULL
_filetype
              = 4
_header
              = String_Type[190]
_ncols
              =
                 13
               = 12
nrows
SPEC_NUM
              = Short_Type[12]
```

```
TG_M
                   Short_Type[12]
                =
TG_PART
                   Short_Type[12]
                =
TG_SRCID
                =
                   Short_Type[12]
                = Float_Type[12]
Х
                = Float_Type[12]
Y
CHANNEL
               = Short_Type[12,8192]
COUNTS
               = Short_Type[12,8192]
               = Float_Type[12,8192]
STAT ERR
BACKGROUND_UP = Short_Type[12,8192]
BACKGROUND_DOWN = Short_Type[12,8192]
       = Double_Type[12,8192]
BIN LO
BIN HI
                = Double_Type[12,8192]
chips> bin_lo_heg_p1 = pha2.BIN_LO[3,*]
chips> counts_heg_p1 = pha2.COUNTS[3,*]
chips> curve(bin_lo_heg_p1,counts_heg_p1)
0
chips> symbol none
chips> step
chips> xlabel "Wavelength (\AA)"
chips> ylabel Counts
chips> title "ACIS+HEG order=+1"
```

Note that S-Lang follows C conventions by numbering array indexes from 0, rather than 1. Therefore the 4th row (HEG, +1 order) in the PHA2 file is the 3rd row within the pha2 data arrays (hence the syntax "bin_lo_heg_p1 = pha2.BIN_LO[3,*]" and "counts_heg_p1 = pha2.COUNTS[3,*])".

With Sherpa

<u>Sherpa</u> can also be used to plot a PHA2 spectrum. In this case, the spectra do not need to be split first; *Sherpa* reads all the rows and allows you to specify individual ones for plotting or fitting purposes:

```
unix% sherpa
Welcome to Sherpa: CXC's Modeling and Fitting Program
Version: CIAO 3.4
Type AHELP SHERPA for overview.
Type EXIT, QUIT, or BYE to leave the program.
Notes:
    Temporary files for visualization will be written to the directory:
    /tmp
    To change this so that these files are not deleted when you exit Sherpa,
    edit $ASCDS_WORK_PATH in your 'ciao' setup script.
    Abundances set to Anders & Grevesse
sherpa> data acisf00459N002_pha2.fits
The inferred file type is PHA Type II. If this is not what you want, please
specify the type explicitly in the data command.
Warning: could not find SYS_ERR column
WARNING: statistical errors specified in the PHA file.
         These are currently IGNORED. To use them, type:
         READ ERRORS "<filename>[cols CHANNEL,STAT_ERR]" fitsbin
WARNING: backgrounds UP and DOWN are being read from this file,
        and are being combined into a single background dataset.
WARNING: multiple datasets have been input.
         The next available dataset number is 13.
sherpa> sherpa.dataplot.curvestyle="histo"
sherpa> sherpa.dataplot.symbolstyle="none"
```

sherpa> <u>analysis</u> channel sherpa> lp 2 data 3 data 4

Figure 8 to shows the plot of the HEG -1 order (row 3, upper drawing area) and +1 order (row 4, lower drawing area) that is created. To change the axes to CHANNEL vs. COUNTS:

sherpa> ploty 3 counts sherpa> ploty 4 counts sherpa> lp 2 data 3 data 4

These commands alter the plots to look like Figure 9 a. *ChIPS* commands (within *Sherpa*) can be used at this point to customize the plot.

```
sherpa> exit
Goodbye.
```

Displaying the Extraction Regions

Each pha2 file has a second block, named REGION, which stores the regions used by <u>tgextract</u> to extract the source and background spectra. There are three regions associated with each order: source, upper background, and lower background. For an ACIS/HETG observation, this gives 36 regions: 12 spectral components (+/-3, +/-2, and +/-1 for HEG and MEG) times 3 regions apiece (source and two backgrounds).

To look at the columns of a REGION block:

unix%	dmlist "acisf00459N0	02_pha2.fits	[REGION]"	cols		
 Column 	s for Table Block RE	GION				
ColNo	Name	Unit	Туре	Range		
1	SPEC_NUM		Int2	1:32767		Spectrum number
2	ROWID		String[64]		Source or a backgro
3	SHAPE		String[]	16]		Shape of region
4	TG_LAM	angstrom	Real4	0:	400.0	Dispersion coordina
5	TG_D	degrees	Real4	-2.0:	2.0	Cross-dispersion co
б	R[2]	(angstrom	, degrees)	Real4(2)	-Inf:+Inf	Raduis vector for S
7	ROTANG	degrees	Real4	-360.0:	360.0	Rotation angle for
8	TG_PART		Int2	0:9		Grating part index
9	TG_SRCID		Int2	1:32767		Source identificati
10	TG_M		Int2	-62:62		Diffraction order
11	COMPONENT		Int2	-		Component number

ds9 cannot display these regions as they are written in the pha2 file. In order to view them, we need to rename the (TG_LAM, TG_D) columns to (X, Y) so that ds9 knows how to interpret them. We will also need to create images in (TG_LAM, TG_D) coordinates, on which we can display the regions.

The following <u>dmcopy</u> commands create image and region files for the 1st and 3rd orders of the MEG arm:

```
unix% dmcopy \
    "acisf00459N002_evt2.fits[bin tg_lam=0:30:0.08,tg_d=-0.01:0.01:0.00008][tg_m=-1,1,tg_part=
    459_order1.fits
unix% dmcopy \
    "acisf00459N002_evt2.fits[bin tg_lam=0:15:0.08,tg_d=-0.01:0.01:0.00008][tg_m=-3,3,tg_part=
    459_order3.fits
unix% dmcopy \
```

"acisf00459N002_pha2.fits[region][tg_m=1,tg_part=2][cols x=tg_lam,y=tg_d,*]" \
region_order1.fits
unix% dmcopy \

```
"acisf00459N002_pha2.fits[region][tg_m=3,tg_part=2][cols x=tg_lam,y=tg_d,*]" \
region_order3.fits
```

In creating the images, the filter includes + and - orders to obtain more events in the image. Since the regions are the same for +/- orders, it is only necessary to copy one (the + orders were used here). The image limits are typical for ACIS/HETG observations, but will need to be adjusted for other configurations.

To display the event files with the regions ovelaid:

```
unix% ds9 -tile 459_order1.fits -region region_order1.fits -cmap a\
459_order3.fits -region region_order3.fits -cmap a
```

which produces Figure 10 0. We can see that all events are contained within at least one extraction region.

History

- 01 Jun 2004 reviewed for CIAO 3.2: no changes
- 06 Dec 2005 updated for CIAO 3.3: version numbers
- 01 Dec 2006 updated for CIAO 3.4: ChIPS and Sherpa versions

URL: http://cxc.harvard.edu/ciao/threads/examinepha2/

Last modified: 1 Dec 2006

Image 1: Prism view of ACIS-S HETG dataset

IMAGE TABLE TABLE	SPECTRUM 1	ULL 13 cols, 12 r cols, 36 rov		T continued o T character a T on the next 2000-10-31T BS 2000-01-10T ND 2000-01-10T S TT 50814.0	ver multip t the end keyword w 13:18:03 06:47:15	le keyword of each su hich has f / Date d / Date d / Date d / Time s / MJD zu	g string keywor ds. The HEASAR ubstring which the name CONTIN and time of fil and time of obs and time of obs system ero point for t correction	& conven is then UE. e creati ervation ervation
I	SPEC_NUM	TG_M	TG_PART	TG_SRCID	x	Y	CHANNEL	COUNTS
Units					pixel	pixel		count
Types	short	short	short	short	float	float	short	short
1	1	-3	1	1	4124.92	4068.34	short[8192]	short[81
2	2	-2	1	1	4124.92	4068.34	short[8192]	short[81
3	3	-1	1	1	4124.92	4068.34	short[8192]	short[81
4	4	1	1	1	4124.92	4068.34	short[8192]	short[81
5	5	2	1	1	4124.92	4068.34	short[8192]	short[81
6	6	3	1	1	4124.92	4068.34	short[8192]	short[81
7	7	-3	2	1	4124.92	4068.34	short[8192]	short[81
8	8	-2	2	1	4124.92	4068.34	short[8192]	short[81
9	9	-1	2	1	4124.92	4068.34	short[8192]	short[81
10	10	1	2	1	4124.92	4068.34	short[8192]	short[81
11	11	2	2	1	4124.92	4068.34	short[8192]	short[81
12	12	3	2	1	4124.92	4068.34	short[8192]	short[81
	≤ ode: Read/Write Sep 14:00:28 Load		ssing: 11 of f00459N002_pha					

 \leq

Image 2: Prism view of ACIS-S LETG dataset

	SPECTRUM REGION	13 cols, 6 n 11 cols, 18 n		IT continued o IT character a IT on the next 2000-10-311 0BS 2000-01-091 2ND 2000-01-101 2S TT 5 50814.0	over multip) at the end o keyword wh 18:39:28 19:18:45	le keyword of each su nich has t / Date a / Date a / Date a / Time s / MJD ze	s. The HEASAR bstring which he name CONTIN nd time of fil nd time of obs nd time of obs	e creation ervation start ervation stop	ise iue
	SPEC_NUM	TG_M	TG_PART	TG_SRCID	х	Ŷ	CHANNEL	COUNTS	
its					pixel	pixel		count	
pes	short	short	short	short	float	float	short	short	
1	1	-3	3	1	4162.24	4084.68	short[8192]	short[8192]	
2	2	-2	3	1	4162.24	4084.68	short[8192]	short[8192]	
3	3	-1	3	1	4162.24	4084.68	short[8192]	short[8192]	
4	4	1	3	1	4162.24	4084.68	short[8192]	short[8192]	
5	5	2	3	1	4162.24	4084.68	short[8192]	short[8192]	
6	6	3	3	1	4162.24	4084.68	short[8192]	short[8192]	

<u>ा</u>

Element 1	Element 2	Element 3	Element 4	Element 5	Element 6	Element :
6.992867	6.992034	6.991200	6.990367	6.989533	6.988700	6.987867
10.489300	10.488050	10.486800	10.485550	10.484300	10.483050	10.481800
20.978601	20.976101	20.973601	20.971101	20.968600	20.966100	20.963600
20.978601	20.976101	20.973601	20.971101	20.968600	20.966100	20.963600
10.489300	10.488050	10.486800	10.485550	10.484300	10.483050	10.481800
6.992867	6.992034	6.991200	6.990367	6.989533	6.988700	6.987867
13.985000	13.983333	13.981667	13.980000	13.978333	13.976667	13.975000
20.977500	20.975000	20.972500	20.970000	20.967500	20.965000	20.962500
41.955000	41.950000	41.945000	41.940000	41.935000	41.930000	41.925000
41.955000	41.950000	41.945000	41.940000	41.935000	41.930000	41.925000
20.977500	20.975000	20.972500	20.970000	20.967500	20.965000	20.962500
13.985000	13.983333	13.981667	13.980000	13.978333	13.976667	13.975000
			•			
Processing :	11 of 12				Goto	orward 8a
	ок				Help	

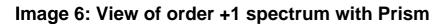
Image 3: Expanded view of BIN_LO column

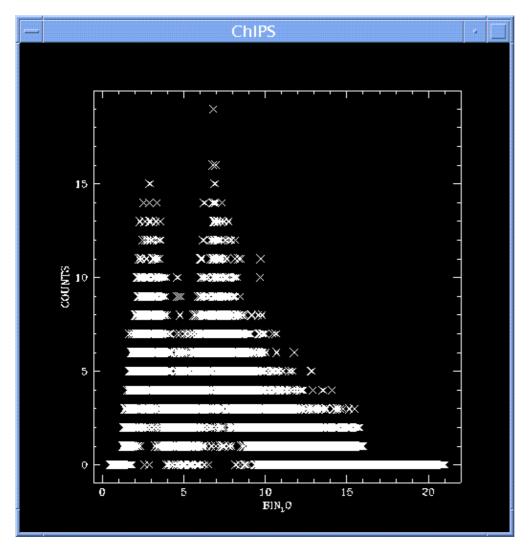
Image 4: Prism view of HRC-S LETG dataset

ABLE	SPECTRUM 1	JLL 3 cols, 2 ro cols, 6 row:		continued ove character at on the next w 2000-11-01T02 5 2000-01-09T03 0 2000-01-09T15 TT 50814.0	er multipl the end o ceyword wh 2:28:25 7:47:42	e keywords f each sub ich has th / Date an / Date an / Date an / Time sy / MJD zer	string keyword of The HEASARC of string which is a name CONTINUE d time of file of d time of observe d time of observe stem o point for time orrection	convention use then continue creation vation start vation stop
	SPEC_NUM	TG_M	TG_PART	TG_SRCID	x	Y	CHANNEL	COUNTS
Inits					pixel	pixel		count
Types	short	short	short	short	float	float	short	short
1	1	-1	3	1	32831.1	32640	short[16384]	short[16384]
2	2	1	3	1	32831.1	32640	short[16384]	short[16384]
	<u></u>							

Image 5: Prism view of HRC-I LETG dataset

File E IMAGE TABLE TABLE	PRIMARY SPECTRUM	Visualization NULL 13 cols, 2 ro 1 cols, 6 row	COMMENT	continued over character at 2001-03-01T11 2000-08-10T12 2000-08-10T23 TT 50814.0	er multipl the end o keyword wh 5:34:00 7:38:15	e keywords f each sub ich has th / Date ar / Date ar / Date ar / Time su / MJD zer	string keyword s. The HEASARC pstring which is ne name CONTINUE nd time of file nd time of obser nd time of obser pstem ro point for tim correction	conventi then co creation vation s vation s
	SPEC_NUM	TG_M	TG_PART	TG_SRC1D	x	Y	CHANNEL	COUNTS
Units					pixel	pixel		count
Types	short	short	short	short	float	float	short	short
1	1	-1	3	1	11783.5	19085	short[16384]	short[18
2	2	1	3	1	11783.5	19085	short[16384]	short[18
	lode: Read/Write Sep 15:06:06 Los		ssing : 1 of 2 01800N002_pha					





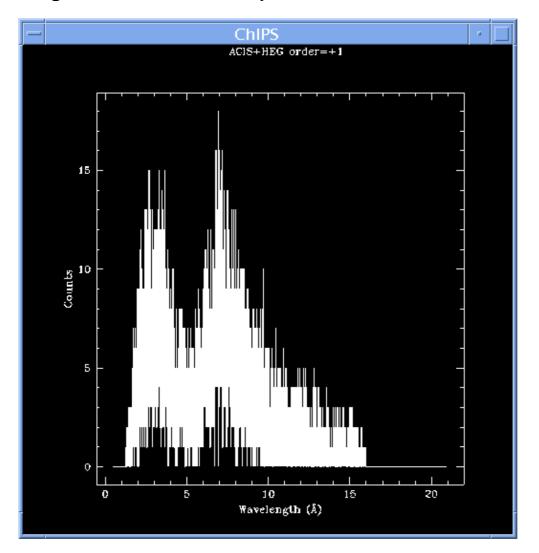


Image 7: View of order +1 spectrum with ChIPS

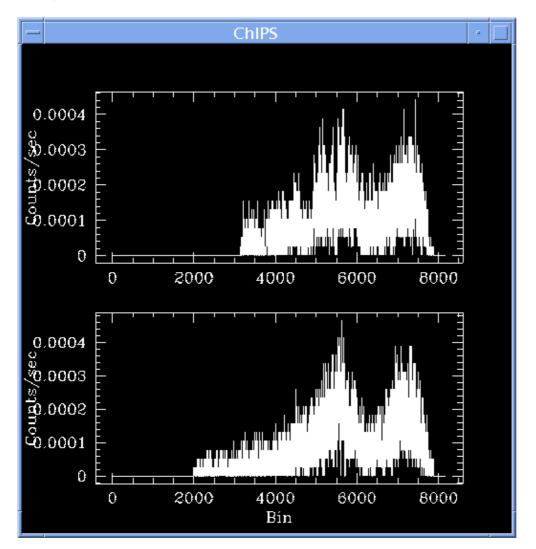


Image 8: View of order +/-1 spectrum with Sherpa: count rate

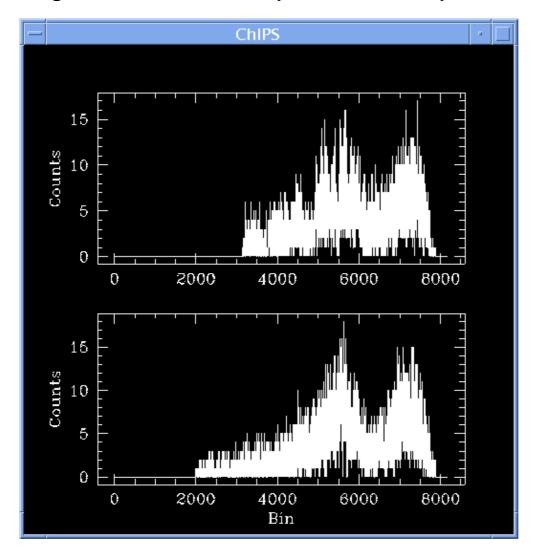


Image 9: View of order +/-1 spectrum with Sherpa: counts

Image 10: Data with source and background regions overlaid

The first order (left) and third (right) order images with extraction regions overlaid.

