





#### Jonathan McDowell (CXC Science Data Systems)





### Chandra data and software

I will report on CIAO (user software) and the standard processing pipeline software, which come into being thanks to:

CXC Data Systems team

software design, development, operation, archive, etc.

CXC Science Data Systems team requirements, documentation, testing, helpdesk, interface with the science community





# **Community Support:**

# Downloads, Documentation, Helpdesk











### **Download by OS**



2019 CUC Meeting





### Downloads (lifetime)

OS	CIAO 4.1	0	CIAO 4	1.11
Linux CentOS	268	700	201	700
Linux Ubuntu	514	102	499	700
OSX Yosemite	23	532	0	463
OSX ElCapitan	76		33	
macOS Sierra	433		56	
macOS HighSierra	0		374	
Source	40		52	
Total	1354		1215	





# Documentation





### Documentation

- CIAO 4.12 routine roll-out and updates
- Improved python documentation
  - Matplotlib, jupyter notebooks, installing external python code
- Improved documentation to pip install astropy
  - Restrictions due to difference in numpy versions.
- Improved dmextract GTI usage description when making light curves.
  - Only the 1st GTI is used. Typically this is correct since the GTIs are usually identical; however, for some situations such as telemetry saturation users need to process each CCD separately.
- Improved links to MARX documentation
- New TSTART/TSTOP dictionary entry:

http://cxc.cfa.harvard.edu/ciao/dictionary/tstart\_tstop.html







## Internal technical documentation

Will review for migration to public docs

- CC mode spatial mask vs. GTIs
- Definition of TSTART and TSTOP
- Discussion of ACIS events not processed beyond L0
- Analysis challenges using on-board energy filters
- Analysis challenges using on-board event sub-sampling
- WCS issues in dmregrid tools
- Computational underflows when computing probabilities
- Improvements to ACA products' metadata.







#### X-ray astronomy primer (cxc.harvard.edu/cdo/xray\_primer.pdf)

Documentation for the wider astronomical community

- Inspired by community interactions at the CXC booth over several AAS meetings and helpdesk.
- For the most part, not *Chandra*-specific, intended as introductions for those new to X-ray astronomy.
  - introduction for astronomers new to X-ray astronomy.
  - overview of data products and highlights analysis considerations.
- Made available as handout for workshop and at this year's winter and summer AAS meetings.
- Intended to be updated with more information over time and with community input.









### Website statistics: 2018-09-01 through 2019-08-31

	CIAO	Sherpa	Chips
Sessions	64,253	16,606	4,001
Users	21,804	7,332	3,201
Page Views	203,779	43,271	6,825
Duration	5:55	4:16	1:19





### Long Term Website Trends





#### Page Views



#### 2019 CUC Meeting





# Helpdesk





### Helpdesk Stats

	2018	2019
Number of Tickets	298	317
Median time to 1st contact [hrs]	1.01	1.67
Median time to close [hrs]	7.07	7.78
Maximum time to close [days]	50	46
% handled by techs	84.9	83.3

Maximum time was for a question about limiting sensitivity, blank sky backgrounds, units, and exposure maps; extended over end-of-year & AAS.

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### Long Term Helpdesk Trends





#### Time To Close



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- 2018 - 2019



Time from Submit to Close (hours)





#### **Iterations per Ticket**













### Common helpdesk topics

#### • Examples of bugs

- Problem with find\_chandra\_obsid on macOS with incompatible libssl
- Issue with sherpa optical models (voigt)
- ciao-install issue with proxy servers
- user contributed software patch (parameter interface)

#### • Example of documentation updates

- Definition of TSTART & TSTOP
- Multiple GTI handling when creating lightcurves
- Instruction of installing numpy-compatible version of astropy

#### • Examples of calibration questions passed on to science staff

- Fe line shifts between observations
- Accuracy of wavelengths, x-cal between HETG & LETG
- Differences between nH from CXC's 'colden' and HEASARCs 'nH'
- Creating HRC RMF files on arbitrary grids





# Community





### Community Support

- January 2019 AAS/Seattle First Chandra/CIAO AAS Workshop
- September 2019 Chandra/CIAO regional workshop in Bologna





### CIAO Workshop #15 at AAS233/Seattle Jan 2019



- 2 days workshop held on Saturday and Sunday before conference.
- 29 Students registered out of 30 available seats. 24 students attended including several walk-ons.
- Mostly grad students, some Professors and even High School students.
- Morning talks (incl 1 remote), afternoon hands on.





## CIAO Workshop #15 User Feedback

#### • YES! 100%

- Overall, did you find the CIAO workshop beneficial?
- Did you get all the support you needed during the hands-on session?
- Would you recommend to your students, colleagues, and|or adviser that they attend a similar workshop?

#### Are there specific topics you wish we had addressed?

- I didn't come in with any specific topics, I wanted support getting started on CIAO and doing the workbook. I came out with a solid understanding of the functions in the workbook, and how to problem solve and look up other functions. Though the program was meant for graduate students and postdocs, everyone was very supportive and worked with me until I understood how to do it myself.
- **Gratings** data -- I felt like there was a large emphasis on imaging spectroscopy
- Grating, more spectral analysis, HETG/LETG
- The nitty gritty of spectroscopic analysis; running a point source from observation to Lx, Gamma, etc.
- Additional information on **grating** spectroscopy would be beneficial.
- More specific demo of DS9 using.





## CIAO Workshop #15 laudatory quotes

- Everyone's positivity was appreciated!
- Thank you very much for an excellent workshop!
- The duration and the content are perfect. Thanks the team so much!
- Keep up the great work!
- One of the best I have seen
- The documentation appears to be very well laid out, and the workbook is very helpful.







# CIAO Workshop #16 after X-Ray Meeting, Bologna, Italy 14-15 September 2019

- first time in Europe
- 24 participants from institutions in 11 countries







## CIAO Workshop #16

- Followed feedback from workshop #15 and added grating + advanced stats talks
- Comments still trickling in
- For now YES! **100% to the questions**:
  - Overall, did you find the CIAO workshop beneficial?
  - Did you get all the support you needed during the hands-on session?
  - Would you recommend to your students, colleagues, and|or adviser that they attend a similar workshop?

## CIAO Workshop #17 AAS235/Honolulu - Jan 2020

• Registration currently open





2019

### Long Term Social Media Trends





2018

2019 CUC Meeting





# CIAO 4.11 and Scripts Overview

## CIAO 4.11 Highlights

- Python 3 only (previous release had both Python 2 and 3)
- Improved python compatibility
- Ships with Jupyter notebook and matplotlib

## CIAO 4.12 plan

- Remove ChIPS and rely on matplotlib
- Scripts with plotting rewritten to use matplotlib
- DS9/DAX analysis scripts which used ChIPS now using DS9 internal plotting
- Move UI tools from Gtk+ version 2 to version 3
- SSL libraries in package, simplifying multiple platform support
- focus support on tcsh and bash shells
- ciao-install:
  - Single Linux distribution
  - Single macOS distribution
- Experimental support for installing CIAO using the conda package/environment management system
  - Linux and macOS
  - currently finalising OTS versions (primarily Python)

## Repro 5 (2020)

- First complete archive reprocessing in 6 years
- Update header keywords etc (like running chandra\_repro on all archive data)
- Improved off-axis angles (see next slide)
- Other algorithm improvements
- Pre-compute grating responses
- Uniformly processed archive

# Off-axis angle algorithm

- New off axis angle calculation to be introduced in 2020
- Will not change RA/Dec values
- Will change DETX/DETY values and off-axis angles (up to 20" change)



Background:

Fiducial lights on ACIS/HRC projected onto aspect camera field

- gives us shifts of ACA frame wrt instrument frame every 2 seconds
- Up to few arcseconds over an observation
- Trend of 20" over life of mission BUT
- degeneracy: cannot separate instrument/mirror shift from mirror/ACA shift
- 20 years of PSF observations finally allow us to disentangle this: mirror/ACA is shifting, not instrument/mirror
- Oops! We guessed wrong in 1998
- Minimal science impact so far, but want to fix, just in case the change trends larger





### **Contributed Scripts**

- CIAO 4.11 highlights (four releases: December, April, May, and August)
  - bugfix: srcflux was overestimating fluxes (net\_photflux, net\_flux), typically by 5% to 15%
  - Major DAX update:
    - replace ChIPS plotting by DS9 (BLT toolkit)
    - new and updated tasks
    - provide access to the output products of the tasks
  - Create psfmaps needed by wavdetect/celldetect in "imaging" scripts: eg fluximage, merge\_obs
  - Convert scripts and modules to use Matplotlib rather than ChIPS (eg deflare, calc\_ecf)
  - Retain all columns in HRC L2 data products
  - Address user tickets (mainly corner cases such as event files with 0 exposure)
  - Update CSC scripts in preparation for the CSC2 release
  - Continue our pithy naming convention: convert\_ds9\_region\_to\_ciao\_stack

#### 2020 Planning

- DAX: continue plotting improvements (thanks to recent DS9 updates)
- Removal of ChIPS support
- Support distribution via conda





### Tracking Testing with Trello

- Seeded with information from weekly CIAO builds (ciaox)
- Used to track testing status: backlog, working, test submitted, added to automated regression tests, etc.
- Identify doc update, team assignments, etc.







# Catalog Support





SDS supports the catalog at 2.0 FTE level (Primini, Burke, Lee, with some additional relevant work by others)

Details of progress in I Evans presentation

Primini:

- Statistical Characterization of the catalog
- Quality assurance and reviews on data, detect list and data products

#### Burke

- convex hull support,
- quality assurance/reviews of source properties pipeline
- visualization (catalog display in WWT)

Siemiginowska

- fitting

Burke, McDowell, Lee, Primini, Siemiginowska - Documentation.









# Sherpa





### Sherpa Development

- Standalone Releases: 4.10.1 (Oct.16), 4.10.2 (Dec 14), 4.11.0 (Feb.20, 2019), 4.11.1 (Aug.1, 2019)
- Sherpa 4.11 for CIAO was released on Dec. 13, 2018; It contains the code introduced in the Sherpa standalone during 2018 development year:
  - Sherpa runs under Python 2.7 and 3 (3.5 in CIAO; 2.7 will be deprecated)
  - The main updates include:
    - the XSPEC models in the version 12.10.0e and modifications to model parameters names and limits to reflect changes introduced by XSPEC
    - •matplotlib for plotting in CIAO
    - Details on https://github.com/sherpa/sherpa/releases
- Development in 2019 focused on a few specific issues:
  - Model evaluation on arbitrary grid allowing the size of the PSF bins to be independent of the data bins - partial functionality released in 4.11
  - Function to define a diagonal matrix (e.g. RMFs) 4.11
  - Function to calculate uncertainties on line equivalent width using sampling 4.11
  - Performance updates to simultaneous fitting of multiple data sets. in testing for 4.12
  - Documentation changes use *sphinx* for building web documents 4.11
    - https://sherpa.readthedocs.io/en/4.11.1/





#### $\leftarrow \rightarrow \ C$ $\ \ \,$ https://sherpa.readthedocs.io/en/latest/

latest	Docs » welcome to Sherpa's documentation			
Search docs				
INTRODUCTION				
Installation	Sharpa			
A quick guide to modeling and fitting in Sherpa	oncipa			
Sherpa and CIAO	Welcome to the Sherpa documentation. Sherpa is a Python package for modeling and fitting data. It			
USER DOCUMENTATION	was originally developed by the Chandra X-ray Center for use in analysing X-ray data (both spectral and imaging) from the Chandra X-ray telescope, but it is designed to be a general-purpose package,			
What data is to be fit?				
Creating model instances	which can be enhanced with domain-specific tasks (such as X-ray Astronomy). Sherpa contains an			
Evaluating a model	expressive and powerful modeling language, coupled with a range or statistics and robust optimisers.			
Available Models				
What statistic is to be used?	Sherpa is released under the GNU General Public License v3.0, and is compatible with Python			
Optimisers: How to improve the current parameter values	versions 3.5, 3.6, and 3.7. Information on recent releases and citation information for Sherpa is available using the Digital Object Identifier (DOI) 10.5281/zenodo.593753.			
Fitting the data	The last sensitive of Channel and a Ville with Dath and Discuss the 4.4.4.4 miles of			
Visualisation	The last version of Sherpa compatible with Python 2.7 was the 4.11.1 release.			
Markov Chain Monte Carlo and Poisson data	Introduction			
Utility routines	Installation			
	Quick overview			
WORKED EXAMPLES	Requirements			
Simple Interpolation	Keleases and version numbers     Installing a pre-compiled version of Sherpa			
Simple user model	<ul> <li>Building from source</li> </ul>			
AN INTERACTIVE APPLICATION	Testing the Sherpa installation			
Using Sessions to manage models and	A quick guide to modeling and fitting in Sherpa			
data	Getting started     Fitting a grap dimensional data gat			
GETTING HELP	Including errors			
Bug Reports	Fitting two-dimensional data			
Contributing to Sherpa development	Simultaneous fits			
Indices and tables	Sherpa and CIAO			
function countforcer(str) (	User Documentation			

- Reference/API
   Creating model instances
- Beat Triplebyte's online coding quiz. Get offers from top companies. Skip resumes
- Read the Docs

Combining modelsChanging a parameter

• Creating a model instance





## Using Sherpa in Astronomy Research

1178 publications in ApJ, AJ, MNRAS and A&A use Sherpa (since 2001 and including astro-ph abstracts)

252 citations to Freeman et al 2001 SPIE paper

65 research papers in Jan-Sep 2019, 68 papers in 2018

7 PhD theses listed in ADS that used Sherpa





# Instruments Gratings





#### acis\_process\_events updates for CIAO 4.12:

The relationship between Charge Transfer Inefficiency (CTI) and ACIS focal plane temperature is changing as ACIS operates at warmer temperatures. To account for this change, a modified CTI routine has been implemented in acis process events in CIAO 4.12

Additional error checking has been implemented in ACIS process events and descriptive warning/error messages will be reported to users that provide unusual input parameters.

An explanation regarding the setting of event status bits has been added to the acis process events specification version 4.17

Events that land in the region affected by the ACIS frame store shadow will now be set as bad by default (event STATUS[4] = 1). This will affect only a small number of rows near the edge of each ccd.







#### TGCAT updates continue

1969 extractions for 490 objects (increase of 76 extractions, 7 objects since last report) Total volume 444 GB Downloads during period: 199 packages, 65 GB

In progress: upgrade server, port to CentOS 7





Example of data recently added to TGCAT: 13 observations of SN1987a (310 ks)









Working on improved analysis for light curves and spectra for multiple obsids of one source.

Planning cal products and updates to support rare grating/instrument combos (HRC-I/HETG, HRC-I/LETG)

Working on script to handle crowded regions in HETG

Evaluating possibility of adding grating responses to archive

Working on improved grating proposal cookbook

#### A Method for Handling Confusion in Crowded Regions Verified with Marx Simulations

#### **Purpose**

To identify regions in your extracted HETG spectra with contaminating photons from a different source.

#### Input

- RA and DEC of sources if field
- Estimate of source counts
- Observation event file
- User tweak-able parameters

#### Output

For each input object, the python routine will identify where confusion from another source can occur and creates Sherpa/ISIS scripts to allow the user to over plot this location in the spectrum and to ignore it for model fitting. The routine will output several ds9 region files used for visual verification of confusion for very crowded fields.













PSF





### ChaRT



Volume spike in 2019-03 is real. Mostly a single user doing large simulation of Pictor A observations. 2019 CUC Meeting





Obs: 1000 - 2000 eV

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# MARX

- Continue to update calibration files in MARX with CALDB updates
- prepare conda package for easier installation in the future
- user questions show that MARX is mostly use to simulate the PSF, and users commonly hit the limitations
  - for < 1-2 pix (limited calibration for sub-pixel repositioning)
  - for "hook feature" in PSF (not reproducible with current mirror models)
  - for PSF wings > 20 pix (not reproduced well by mirror models)
- Current work is on improving simulated grade distribution for sub-pixel event positioning



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## Visualization







SAOImageDS9 is an imaging program widely used in the astronomical community, well beyond X-ray astronomy.

DS9 is a key part of the Chandra data analysis infrastructure – it directly supports X-ray event files.





#### Releases

- Version 8.0 release in Dec 2018 with CIAO 4.11
- Version 8.1b1 release in May 2019
- Version 8.1b2 release in Aug 2019

#### New Features 8.1

- Improved Mask support, based on WCS
- Support for Animated GIF movies
- Enhanced Plot Tool to replace ChIPS line plot functionality
- Simplified internal code to support a wide variety of operating systems
- Better command line parser, error messages

#### **Ongoing Development**

- Add Chandra Footprint Server support
- Support Catalog FITS Tables
- Position Angle convention
- Convert regions into pixel masks





#### **GitHub Activity**

- 308 Commits
- 39 Release Note Entries

#### **Help Desk**

- 38 From CXC HelpDesk
- 43 From DS9 HelpDesk (Non-SAO)
- 20 From SAO Help Requests

#### Downloads

- 3215 April
- 3304 May
- 2997 June
- 3000 July
- 3014 August





Sample Image Gallery







Crab Nebula Animated GIF courtesy of Kenny Glotfelty

