Chandra Users Committee September 2018



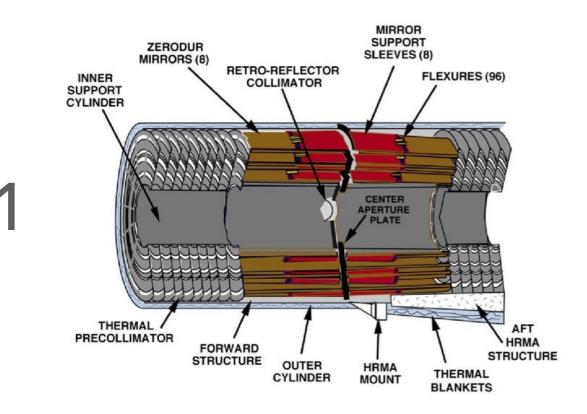
JEREMY DRAKE AND THE CXC CALIBRATION GROUP

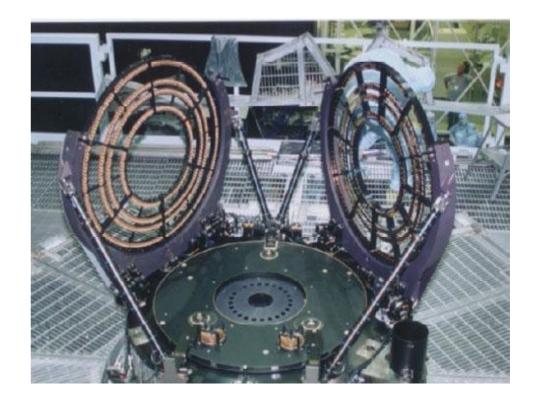
CHANDRA CALIBRATION STATUS

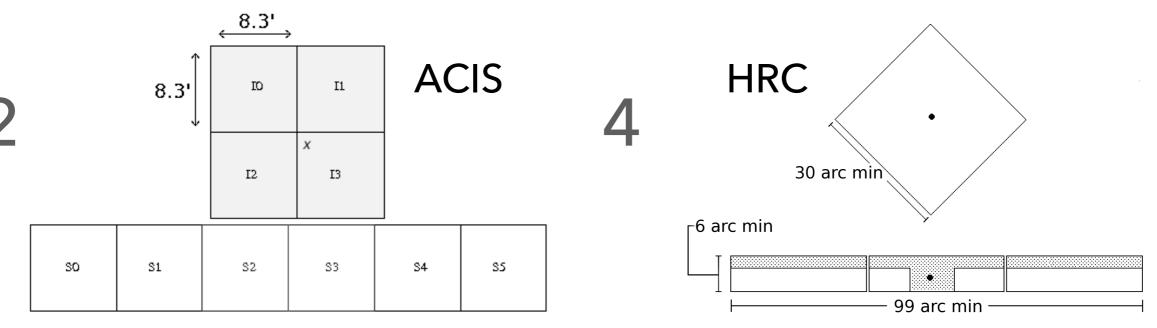
LAST CUC MEETING RECOMMENDATIONS

- Continue; continue to provide regular updates (this presentation)
- Continue to play an active role in IACHEC
 - The calibration group appreciates the CUC's support for these activities
- Bakeout: report on possible outcomes and risks (Paul's presentation next)

CHANDRA HARDWARE COMPONENTS: ORDER OF PRESENTATION







CURRENT IN-FLIGHT CALIBRATION TARGET SUMMARY (~0.9 Ms)

TARGET	INSTRUMENT	PURPOSE
E0102-72 (SNR)	ACIS	CONTAM, GAIN, CROSS-CAL
A1795 (GAL CLUSTER)	ACIS	CONTAM, GAIN, CROSS-CAL
RXJ1856.5-3754 (ISOLATED NS)	LETG+ACIS	CONTAM
MKN 421 (BLAZAR)	HETGS, LETGS	CONTAM, EFF. AREA, CROSS-CAL
3C 273 (QSO)	HETGS	CROSS-CAL
AR LAC (ACTIVE BINARY)	HRC	PSF, GAIN
HZ43 (HOT WD)	LETG,HRC	EFFECTIVE AREA, QE, GAIN
CAPELLA (ACTIVE BINARY)	HETGS, LETGS	DISPERSION, LINE RESPONSE
VEGA (A0 V)	HRC	UVIS UV LEAK

OUTLINE

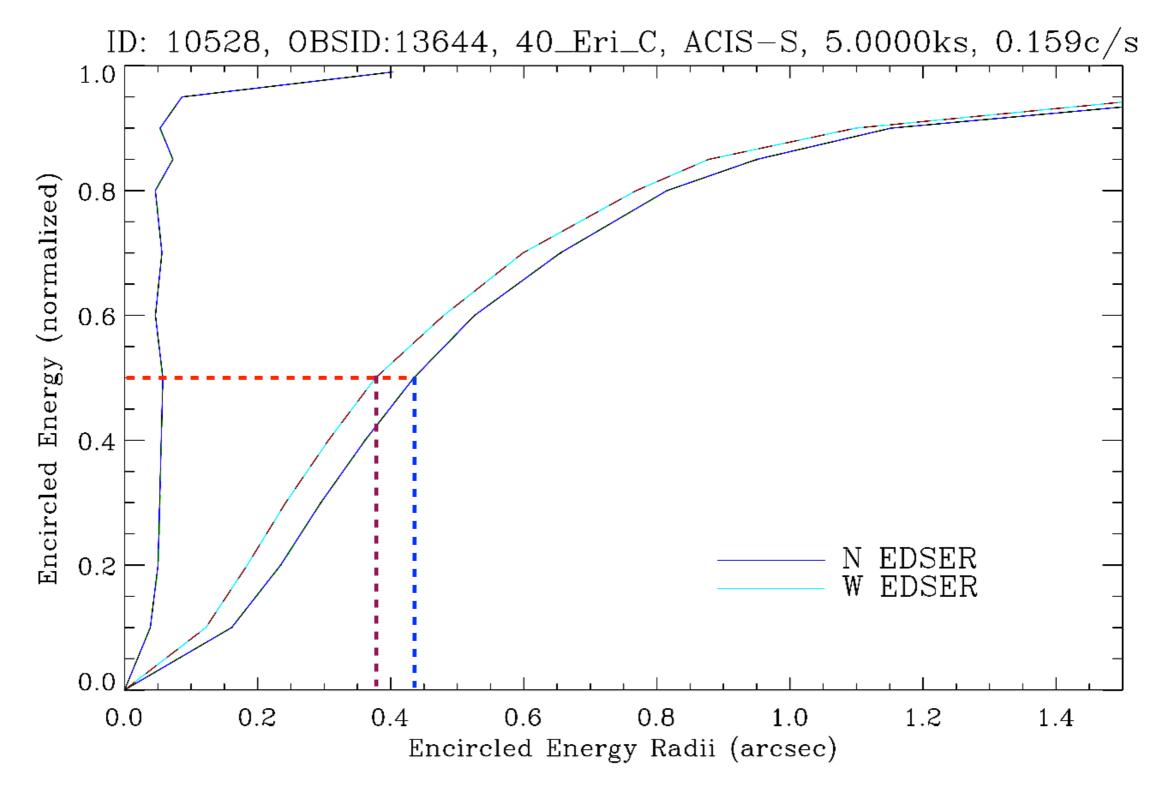
- Main concern: secular change in both ACIS and HRC performance
- Point Spread Function
 - calibrating EDSER; empirical PSFs
- ACIS
 - mid-chip gain droop; contamination
- HETG 0th order throughput
- HRC-S,I
 - QE decline; gain decline

HRMA: POINT Spread Function

ACIS EMPIRICAL PSF (V. KASHYAP, P. ZHAO, D. JERIUS)

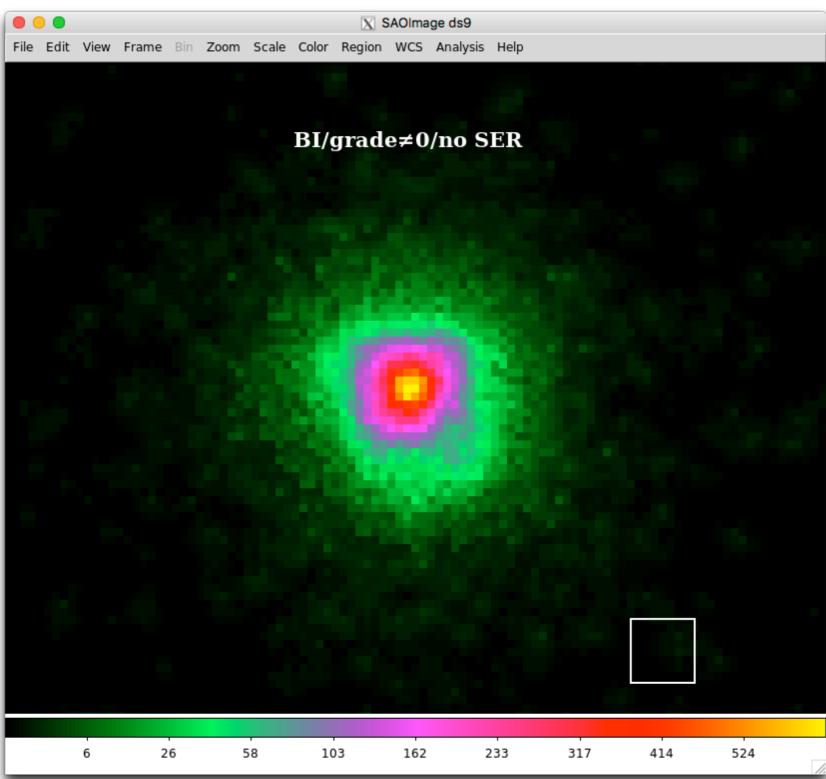
- Verify and calibrate HRMA raytrace model; calibrate EDSER
 - Energy Dependent Subpixel Event Repositioning ACIS images can be sharpened significantly at sub-pixel resolutions
 - Applies corrections to event locations based on photon energy and event grade (Li et al. 2004, ApJ 610, 1204)
 - BUT: EDSER'd PSFs have not yet been calibrated
- Generate an empirical "un-rolled" on-axis PSF using point sources (stars) culled from the CSC

QUANTIFYING EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)

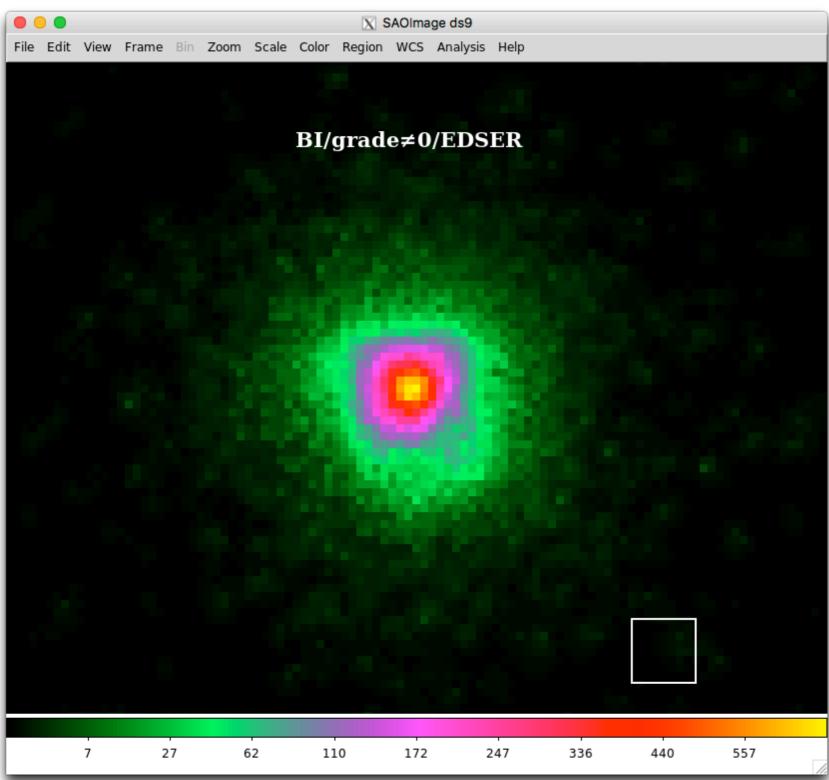


POINT SPREAD FUNCTION

EMPIRICAL PSF WITH EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)



EMPIRICAL PSF WITH EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)



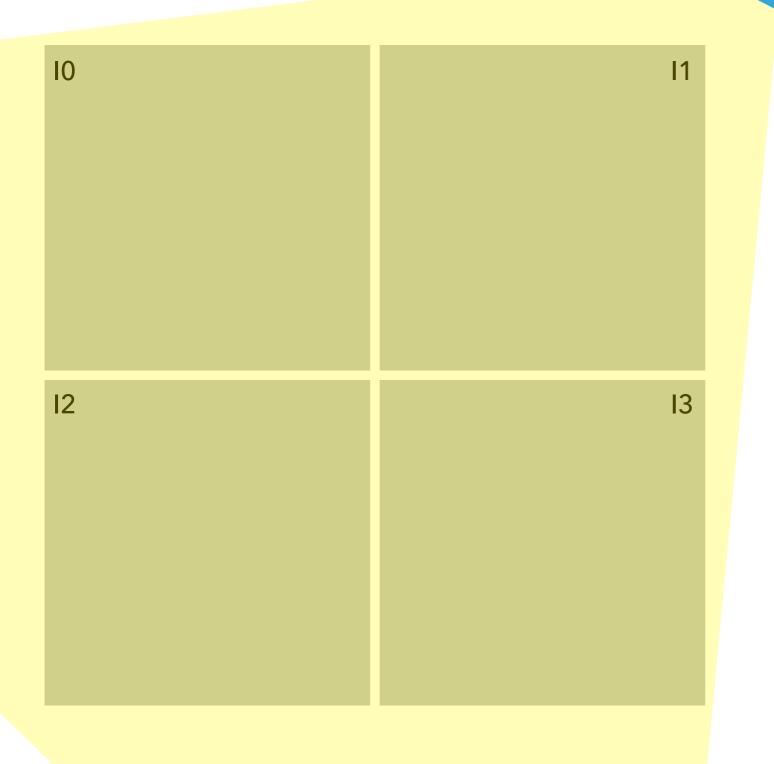
ACIS EMPIRICAL PSF (V. KASHYAP, P. ZHAO, D. JERIUS)

- Currently ~100 verified stars nearly on-axis within right count/frame range (< 0.02)</p>
 - Using Gaia to increase sample size
- To be release as FITS images in 3+ energy bands c. Summer 2019
- This release limited to on-axis

ADVANCED CCD IMAGING SPECTROMETER (ACIS)

ADVANCED CCD IMAGING SPECTROMETER (ACIS)

MID-CHIP GAIN DROOP (T. GAETZ)

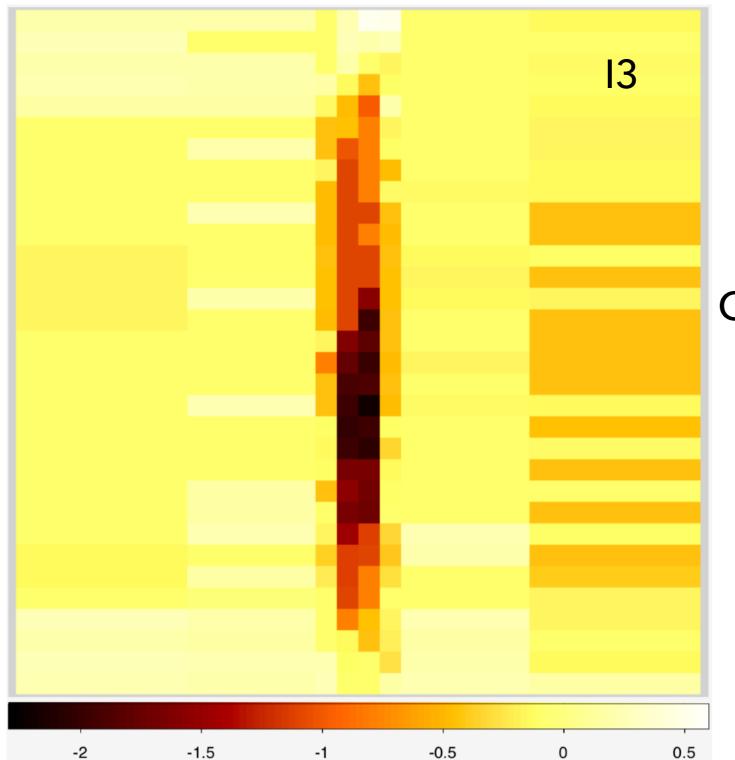


⁵⁵Fe source + Al/Ti target

EC,r

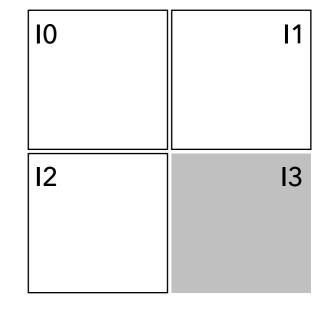
Mn K**α** (5.9 keV) Ti K**α** (4.5 keV) Al K**α** (1.5 keV)

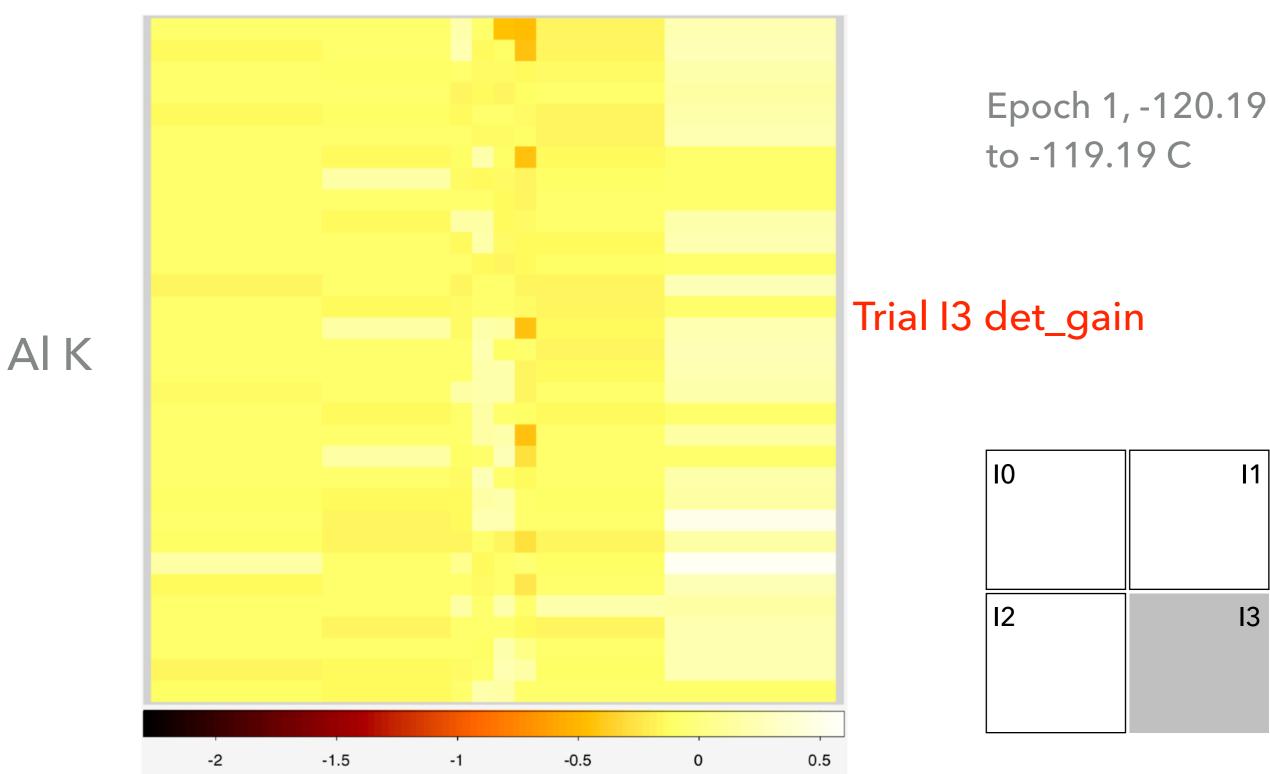
AK



Epoch 1, -120.19 to -119.19 C

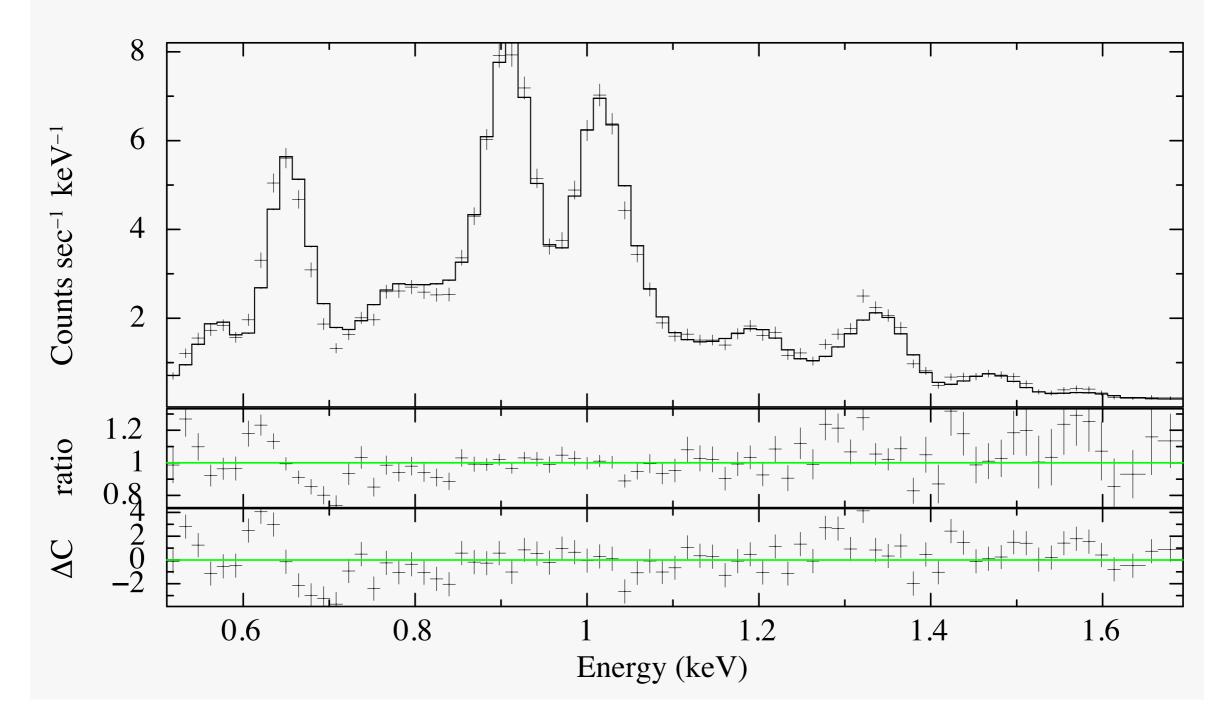
Current det_gain





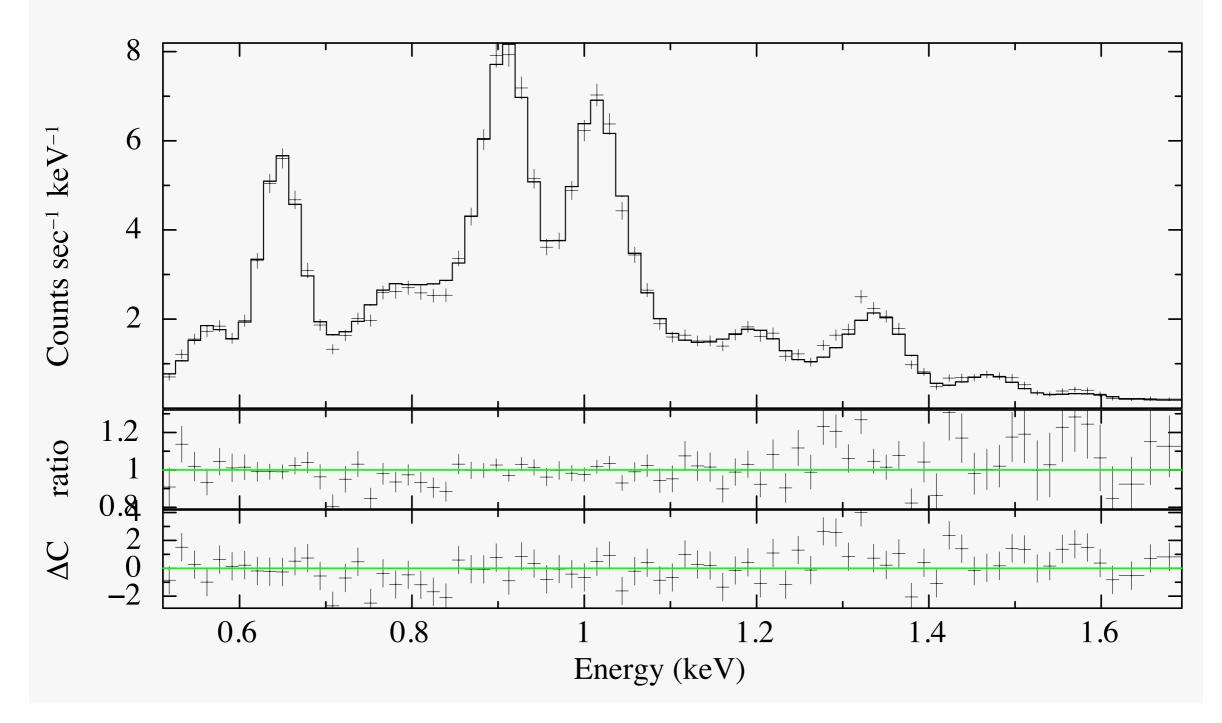
Extrapolation to lower E

E0102: obs01313: I3: (896.13, 103.20)



Extrapolation to lower E

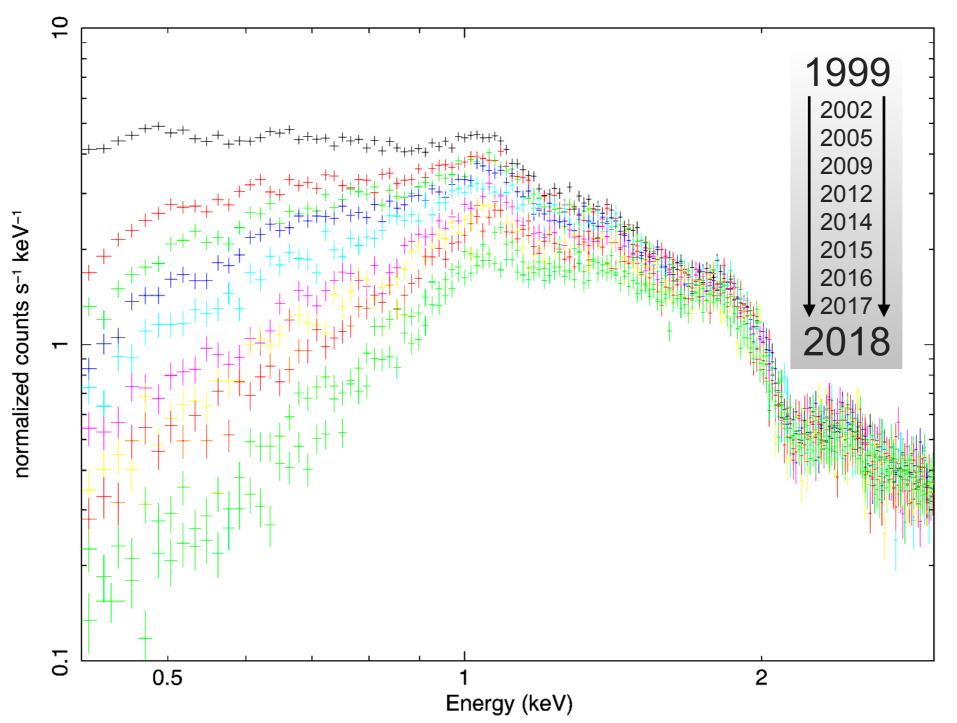
E0102: obs01313: I3: (896.13, 103.20): (O, Ne energy thawed)

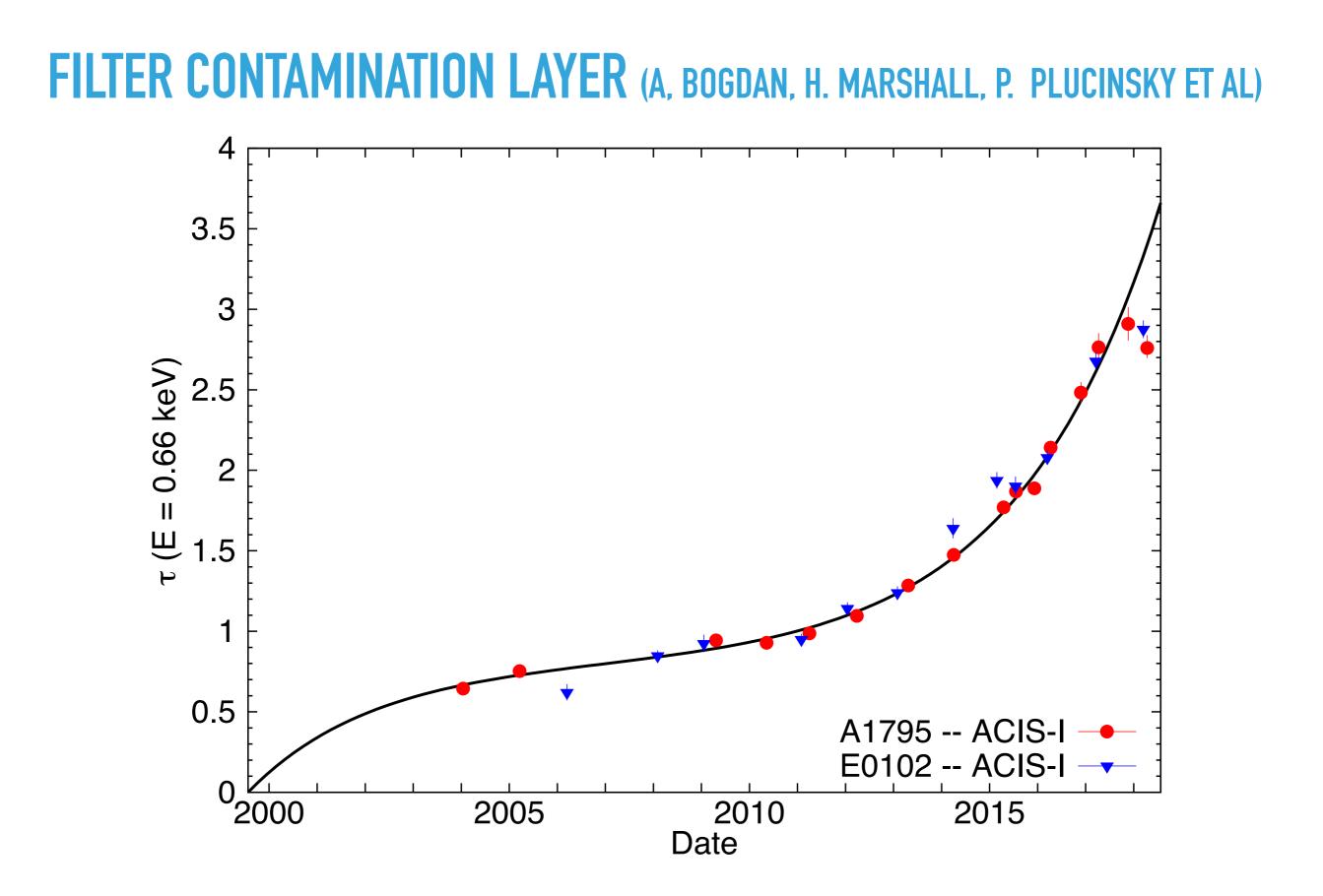


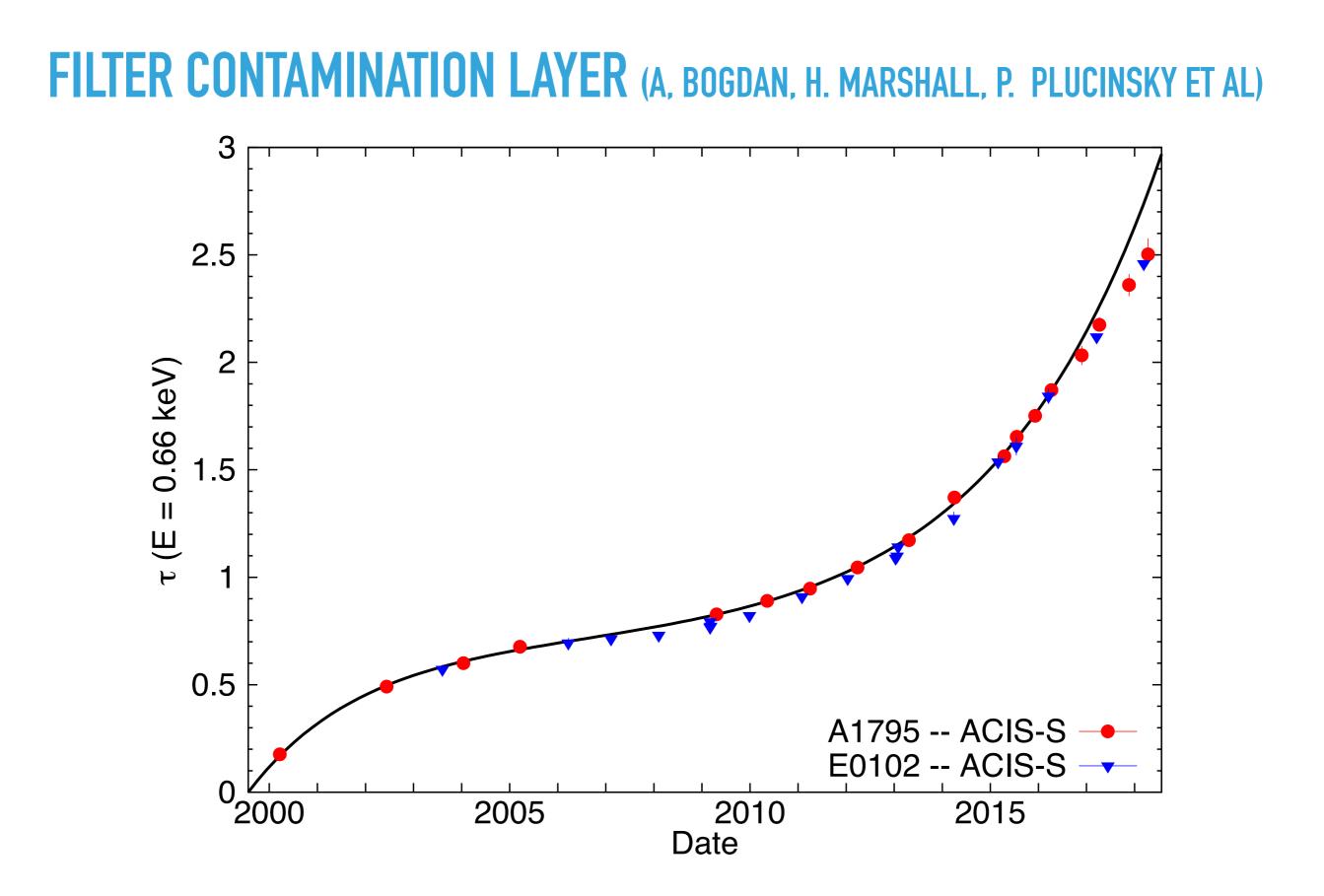
- New trial det_gain for I3 almost ready proof of concept
- Apply same procedure to other FI chips: I0,I1,I2,S2
- tdet_gain time dependence: currently unchanged
- Projected release Spring 2019

FILTER CONTAMINATION LAYER (A, BOGDAN, H. MARSHALL, P. PLUCINSKY ET AL)

data and folded model







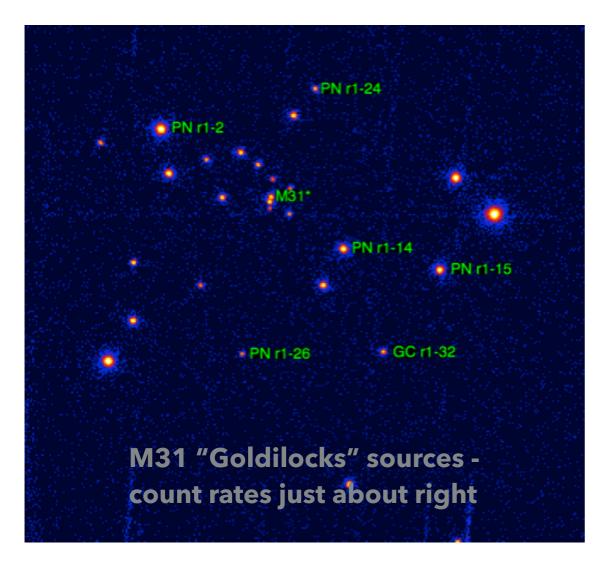
FILTER CONTAMINATION LAYER (A, BOGDAN, H. MARSHALL, P. PLUCINSKY ET AL)

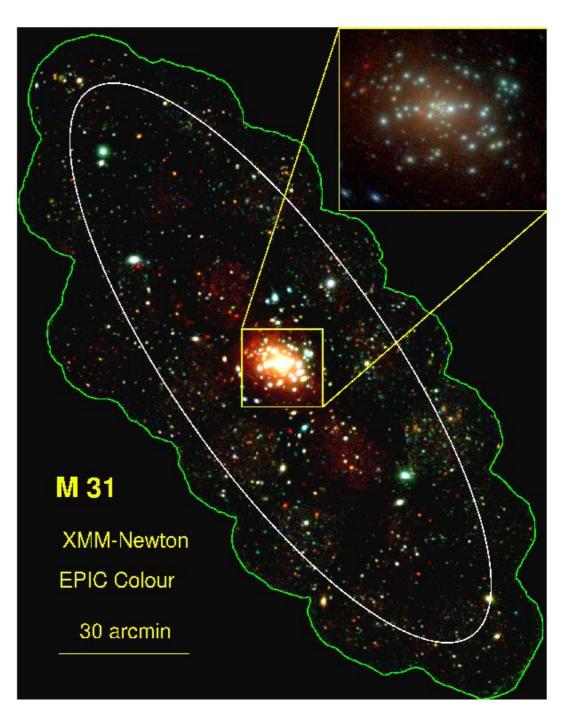
- Rate of accumulation on ACIS-I shows significant decrease compared with model prediction very little since 4/17
 - New contamination model released CalDB 4.7.9
- Rate of accumulation on ACIS-S is significantly lower than the current model prediction - essentially a linear trend rather than exponential
 - Model spatial variation still being tweaked new model release within next 3 months
- BAKEOUT? see Paul Plucinsky's presentation

HETG OTH ORDER (UPDATE)

OTH ORDER EFFECTIVE AREA (NORBERT SCHULZ)

M31 center with Chandra HETG:

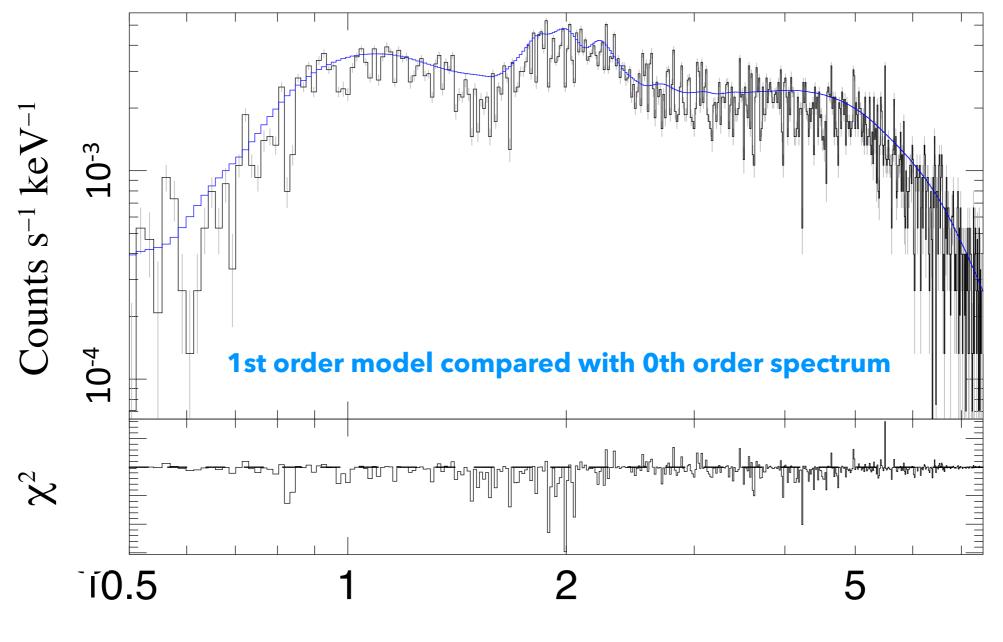




R. Supper et al. 2001, A&A, 373, 63: Soft diffuse emission at M31 bulge (~10³⁹ erg s⁻¹)

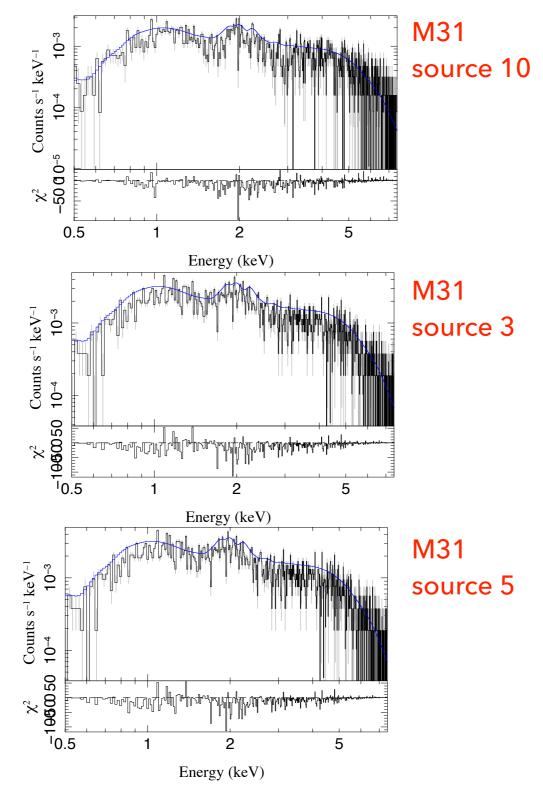
OTH ORDER EFFECTIVE AREA (NORBERT SCHULZ)

M31 source 6, ACIS 0^{th} order , R = 0 to 10 pixel circle extraction



Energy (keV)

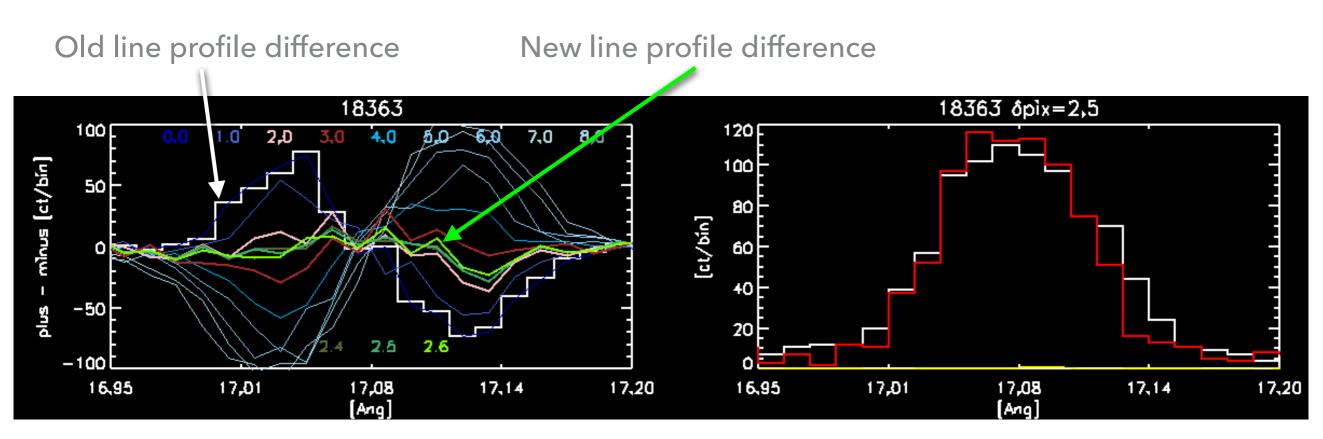
OTH ORDER EFFECTIVE AREA (NORBERT SCHULZ)



- More thorough analysis correcting for both pile-up and diffuse X-ray emission
- Current constraint is 0th order effective area good to better than 10%
- Analysis and full constraints using more sources will be concluded Spring 2019

HIGH RESOLUTION CAMERA

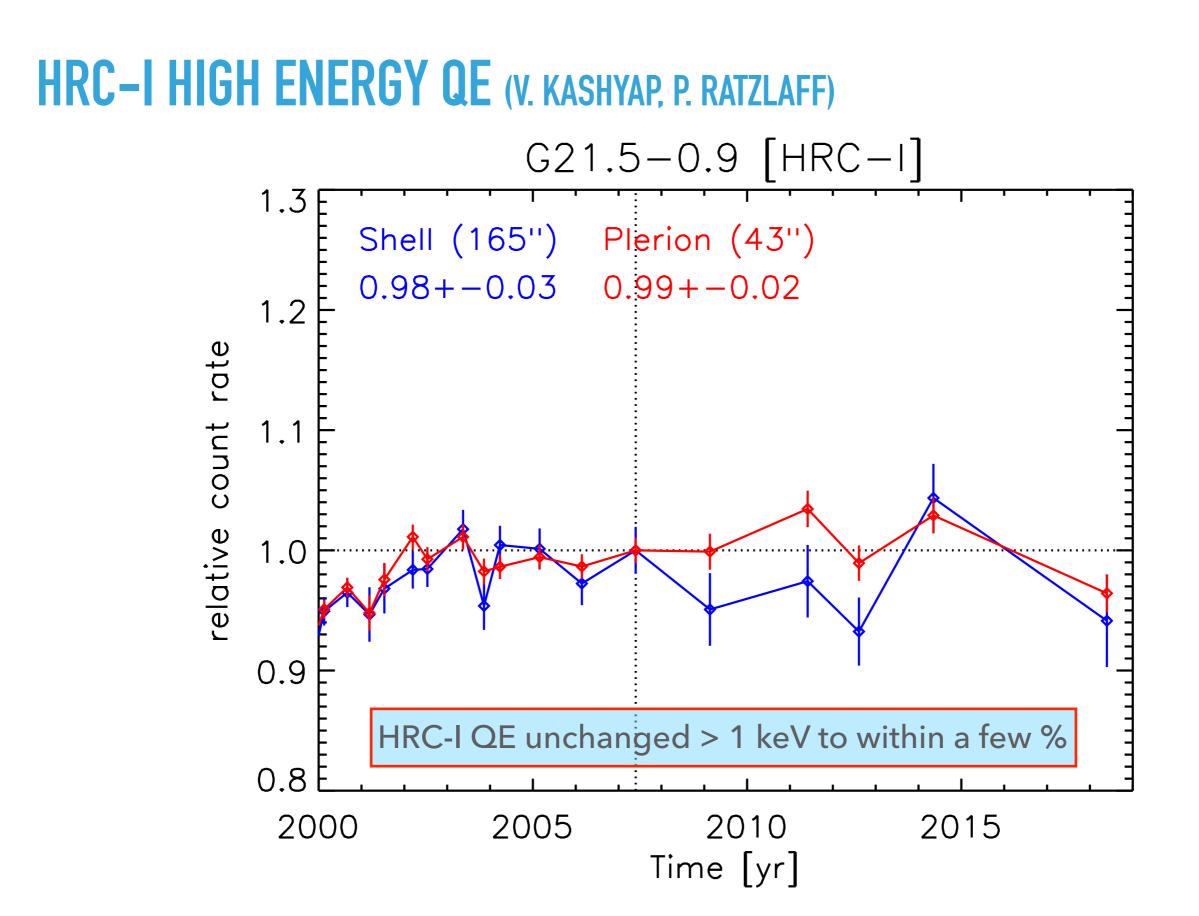
HRC-S DEGAP UPDATE (V. KASHYAP)



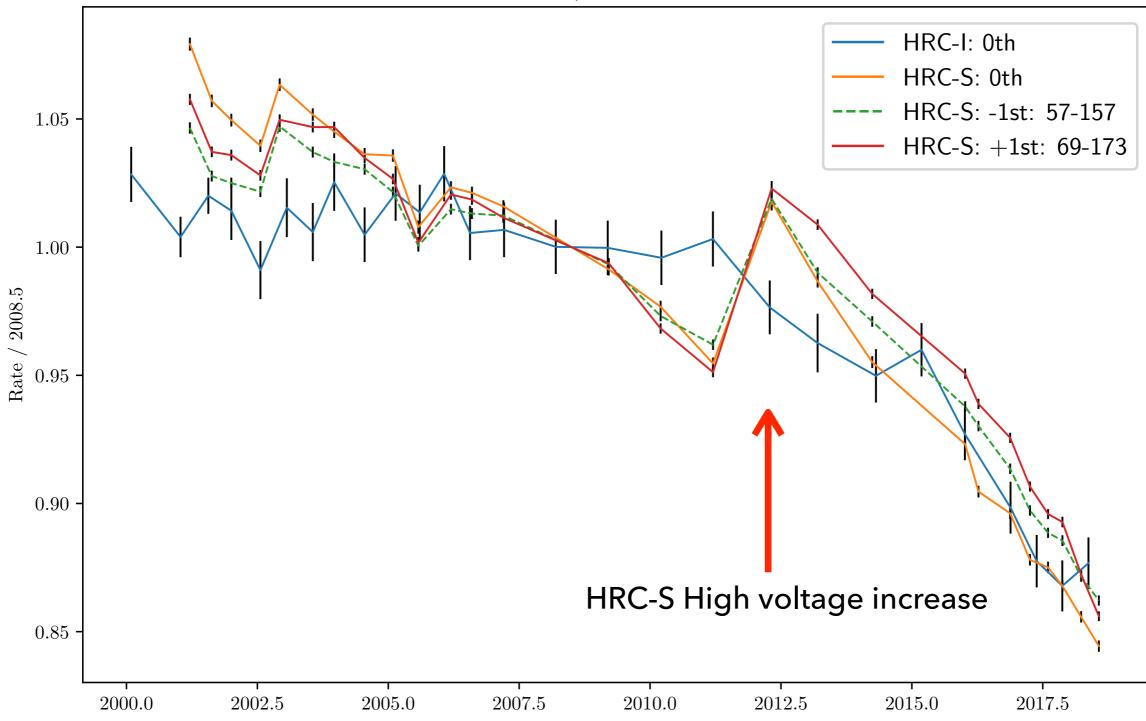
 HRC-S aim point degap solution improved in 2012 ==> sharper images

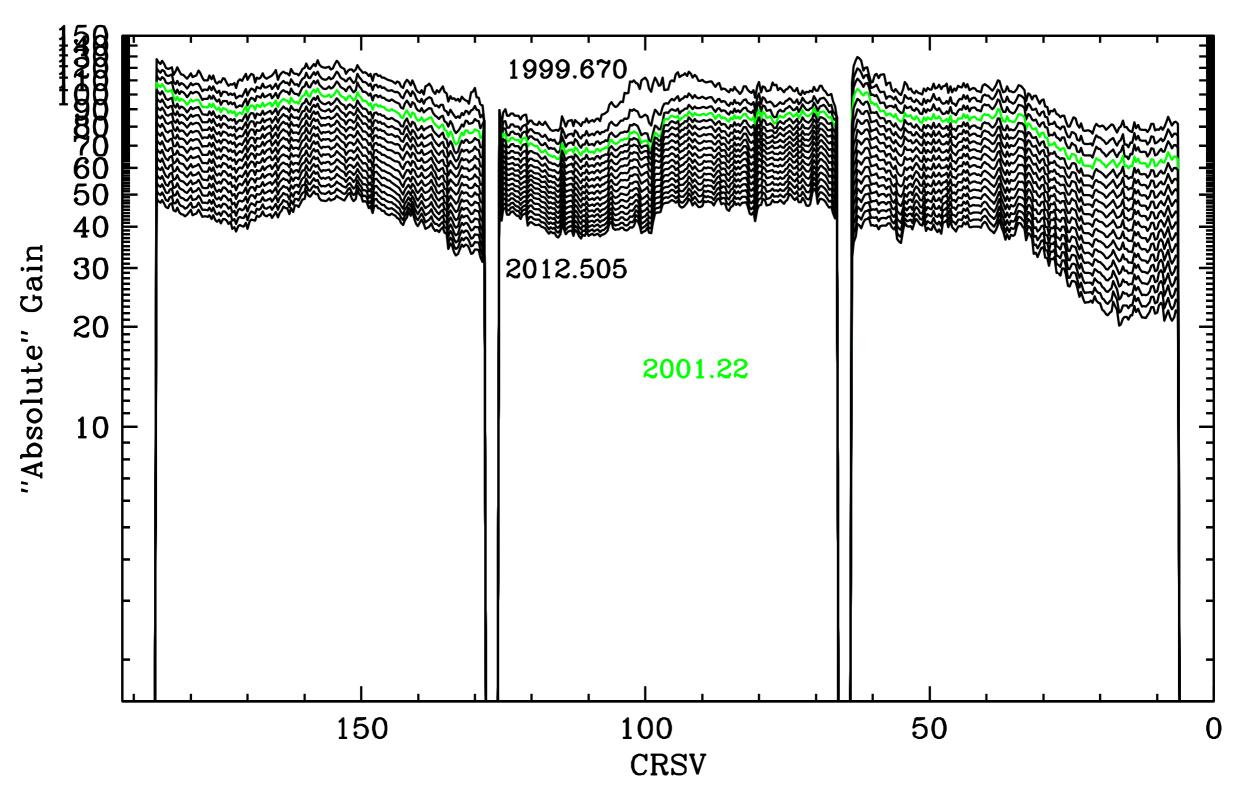
But...! Caused a shift in derived wavelengths relative to 0th order location...

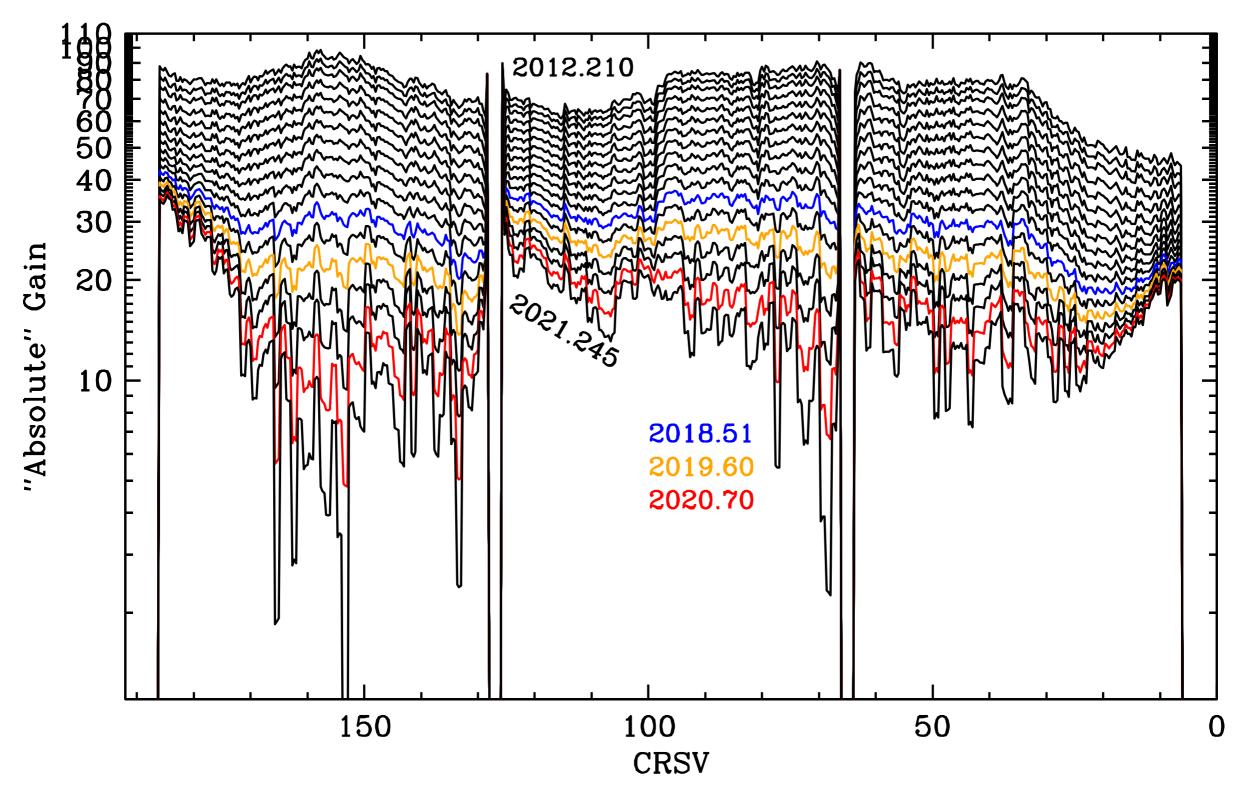
http://cxc.harvard.edu/cal/Hrc/Degap/hrcsdegap_centershift.html

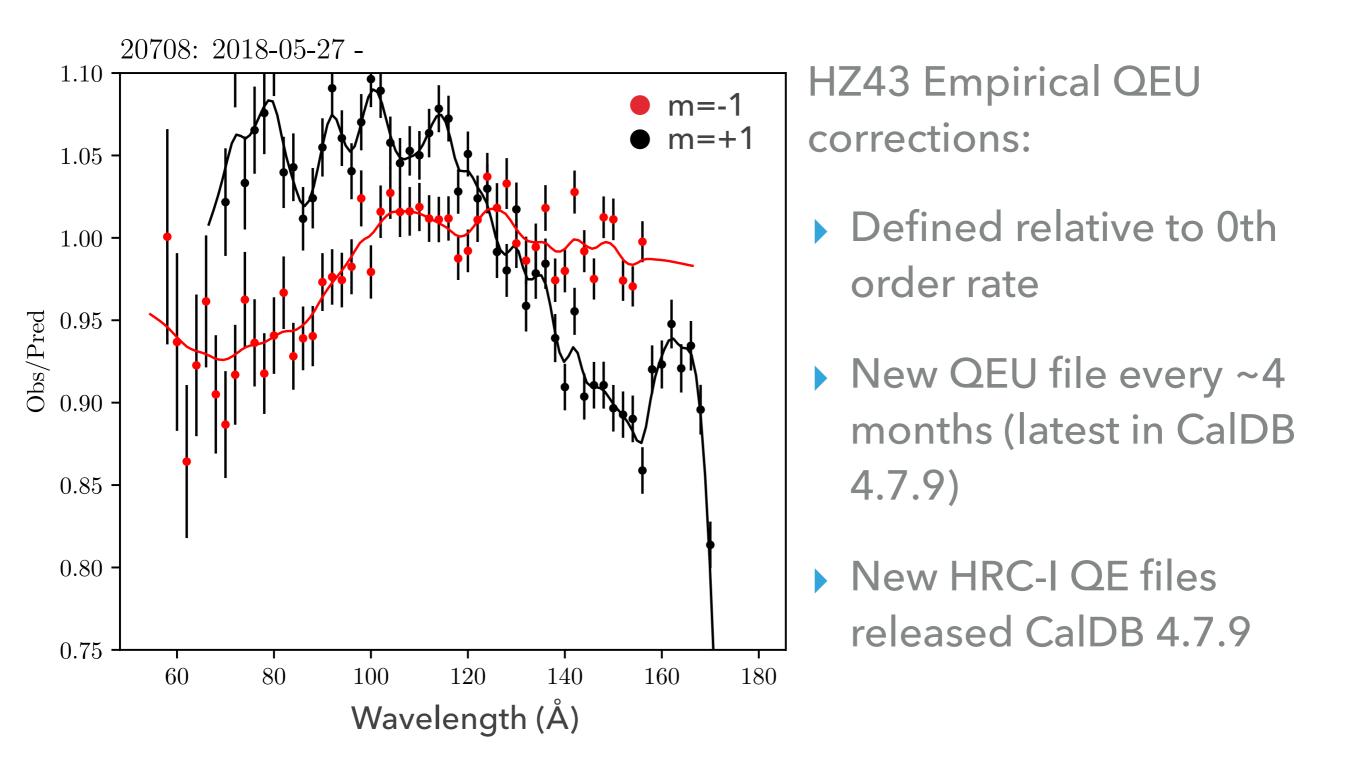


HZ 43: HRC/LETG Count Rates



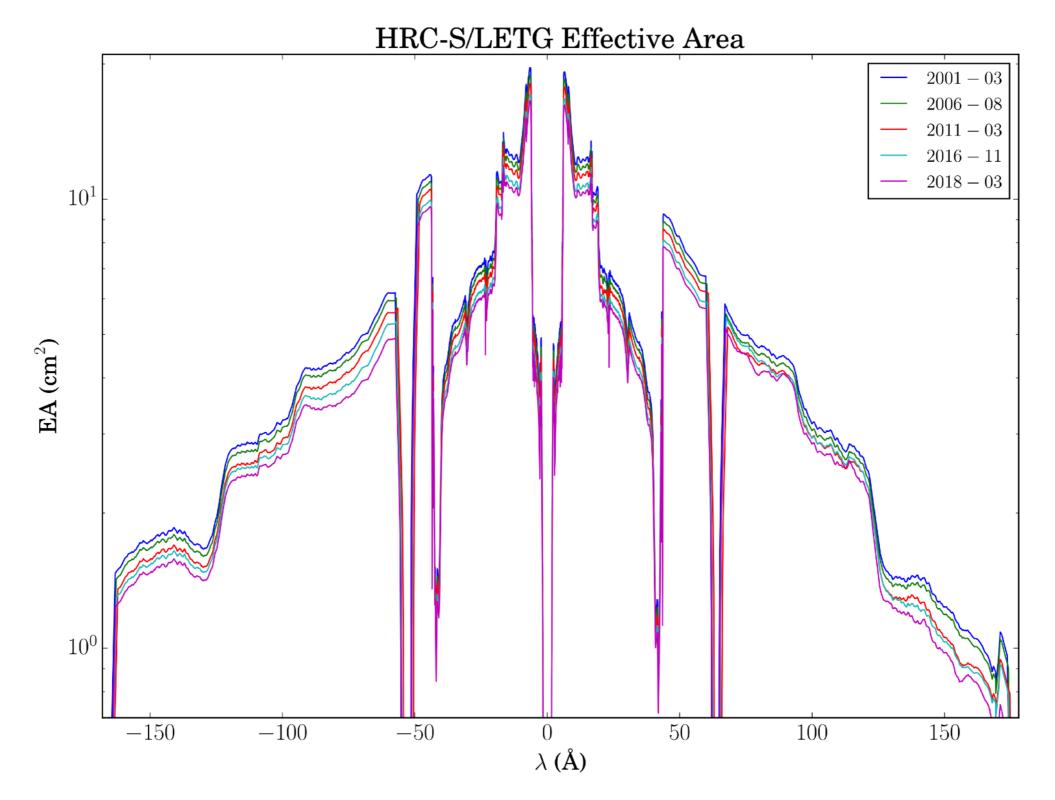






1.02 1.01 1.00 Rate / Predicted 0.99 0.98 HRC-I: 0th 0.97 HRC-S: 0th HRC-S: -1st: 57-157 HRC-S: +1st: 69-173 0.96 2000.0 2002.5 2012.5 2015.0 2017.5 2005.0 2007.5 2010.0

HZ 43: HRC/LETG Ratios to Predicted



SUMMARY AND FUTURE

- On-axis empirical PSF and EDSER calibration well underway; empirical PSF release c. Summer 2019
- ACIS mid-chip gain droop new det_gain release c. Spring 2019
- ACIS contamination model reflecting slower rate of increase has been updated and released for I; release pending for S.
- Continuing HRC-S QE secular changes calibrated
- New HRC-I QE released
- HV increase on HRC-S.... Only a matter of time
- New HRC-S observing mode on thin Al filter being tested