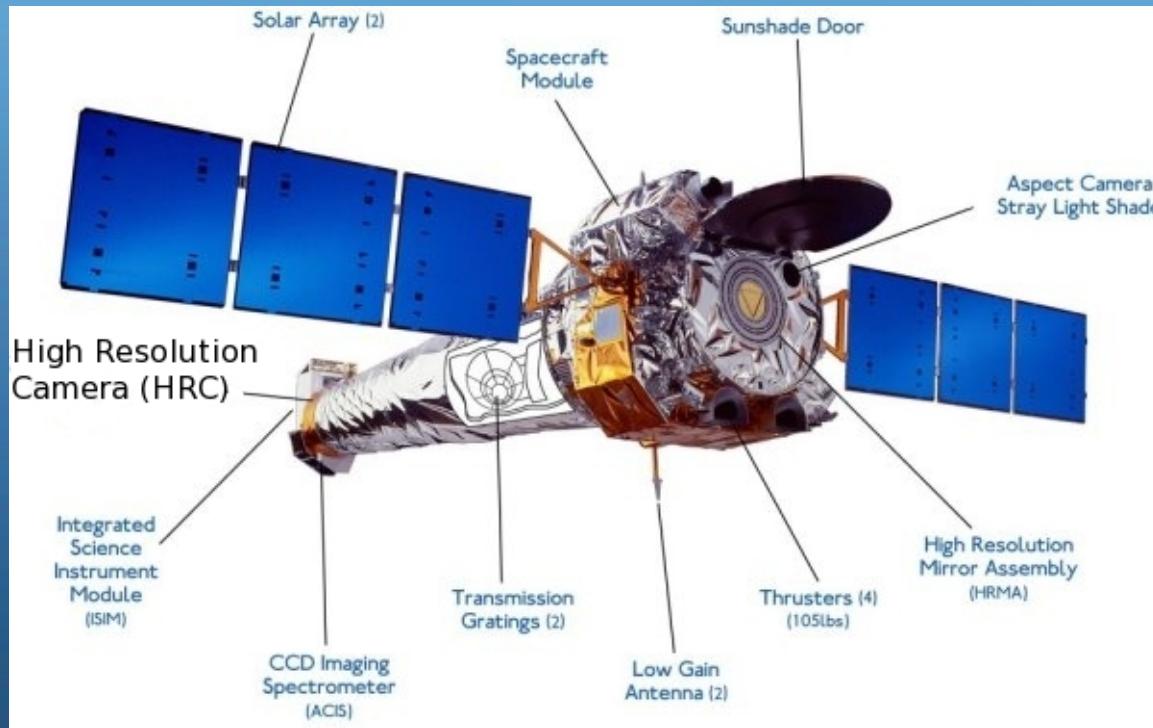


Revisions to the HRC-S Quantum Efficiency at Energies above the Carbon Edge

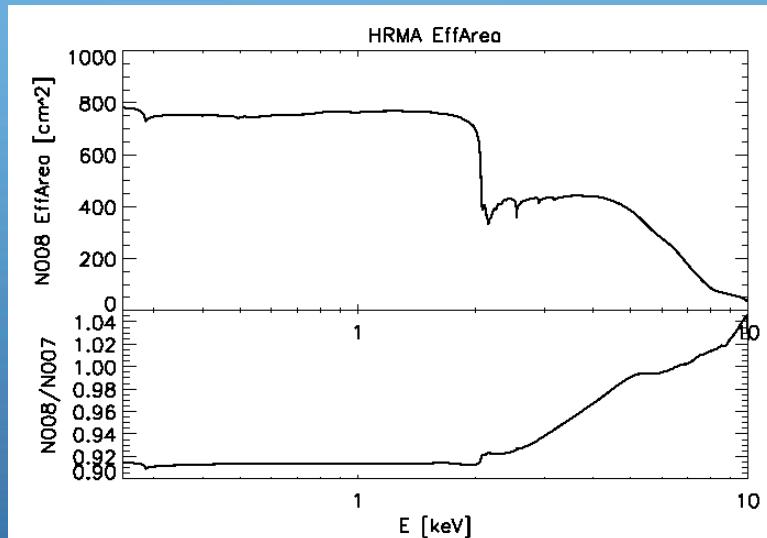
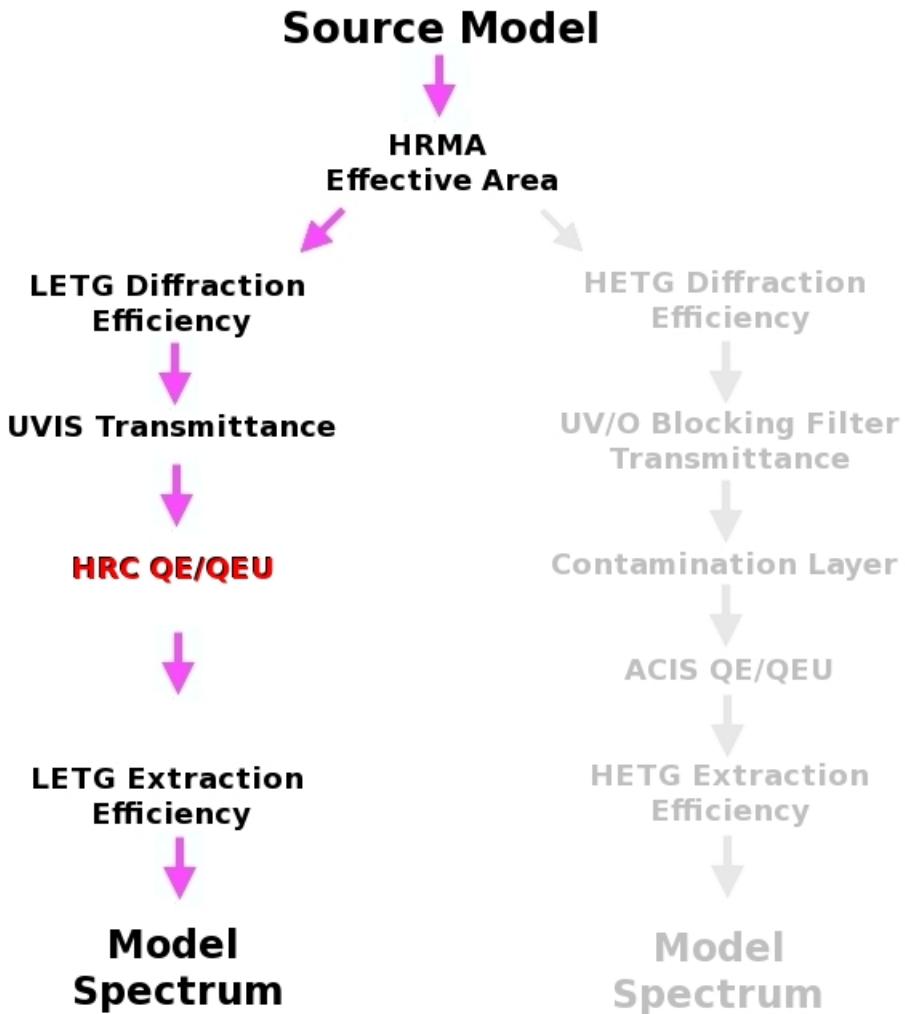


R. Nicholas Durham
Jeremy J. Drake
CXC Calibration Group

Outline

- Motivation for HRC-S QE Revision
 - New HRMA Effective Area Model
 - Accumulated In-flight Observation Database
- In-flight Analysis
 - QE Revision Procedure
 - Calibration Source
 - Models & Fits
- Results
 - Prototype QE
- Future
 - Cross-calibration

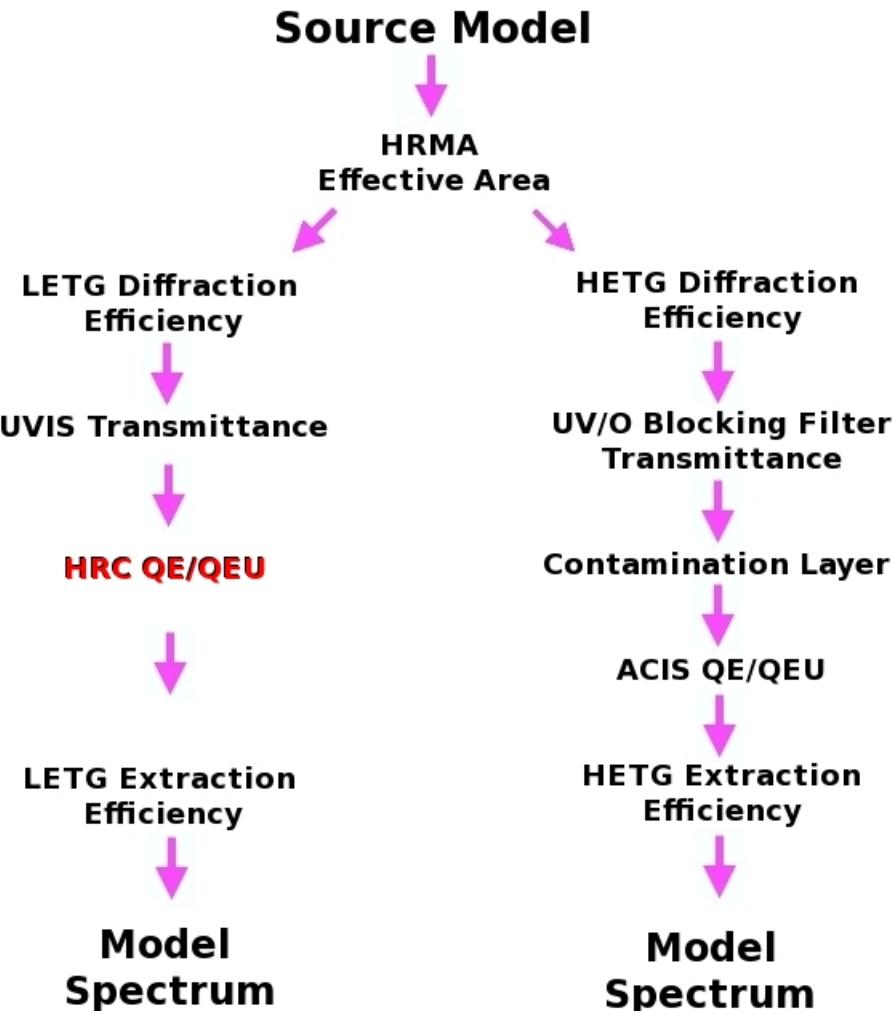
New HRMA Model Effects



*New (N008) HRMA Eff_area
Model Released: Jan 2009*

*In-Flight HRC QE Calibration =
Sensitive to any HRMA change*

In-flight HRC QE Calibration Procedure



HETG/ACIS-S Observation

- Assume a Source Model
- Fold SM through HETG Calibration and fit in XSpec

LETG/HRC-S Observation (of same Source)

- Use HETG/ACIS-S Source Model Parameters (ie PhoIndex, Normalization, etc.)
- Fold through LETG Calibration
- Compare Residuals of Observation Spectrum to Model Spectrum

**HETG Calibration
Uncertainties?**

Calibration Source Selection Criteria

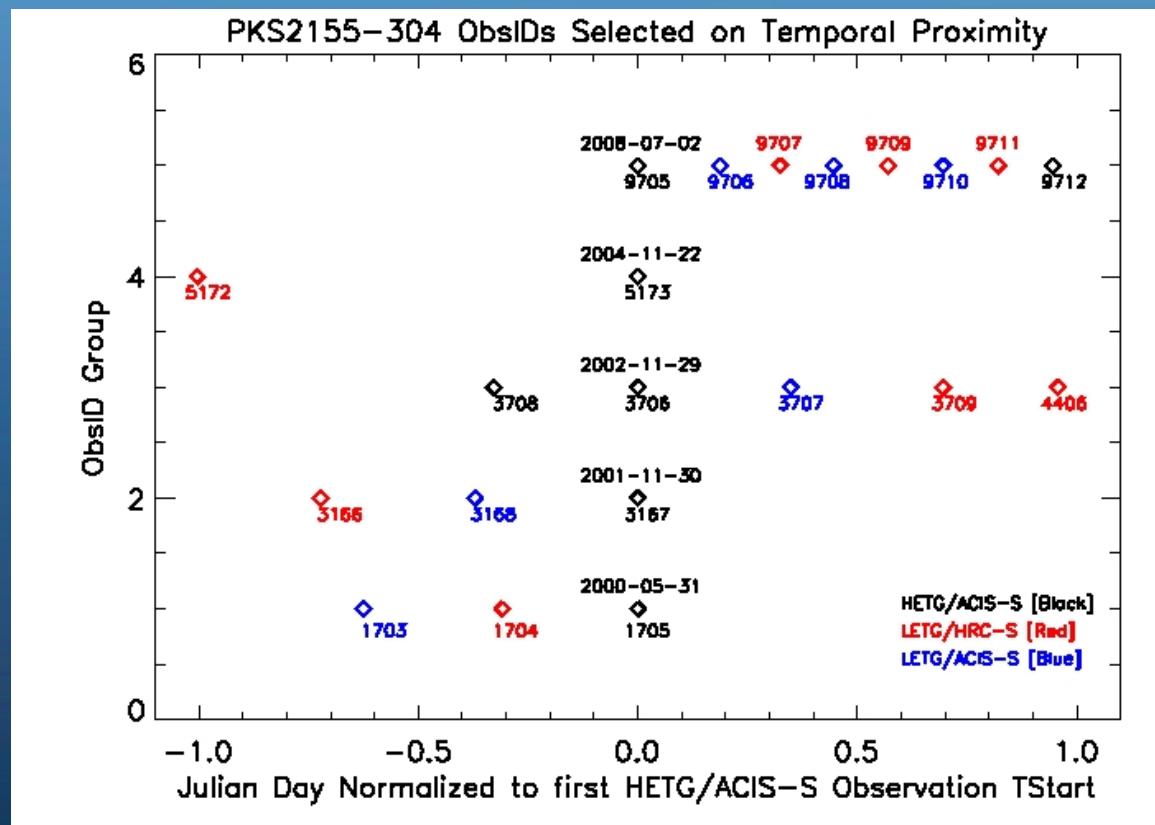
Soft Spectrum
Low Absorption
Multiple Configurations
Consistent Aimpoint

}

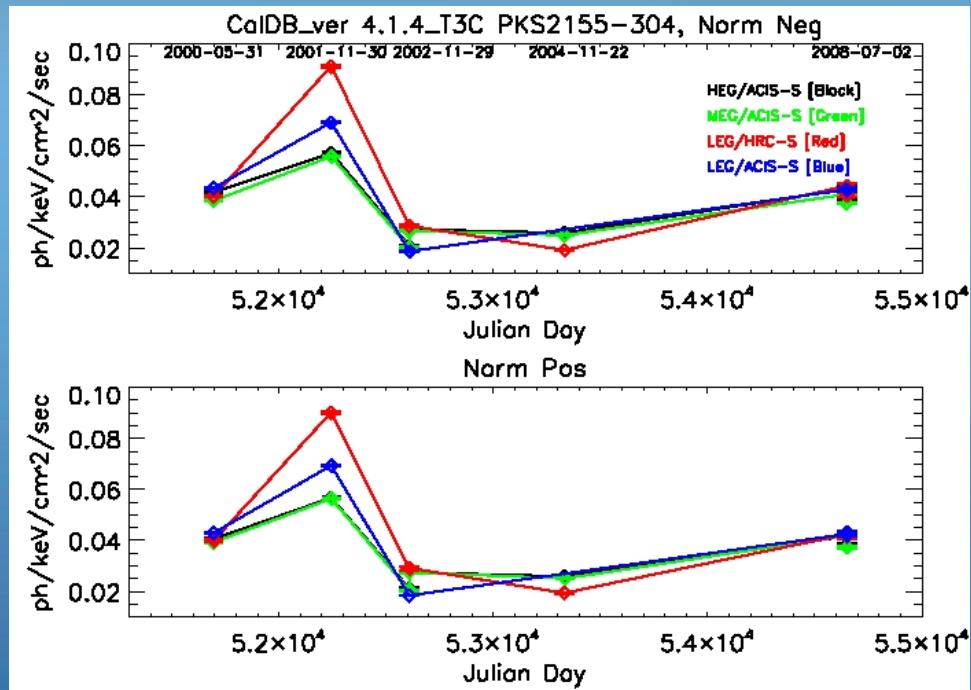
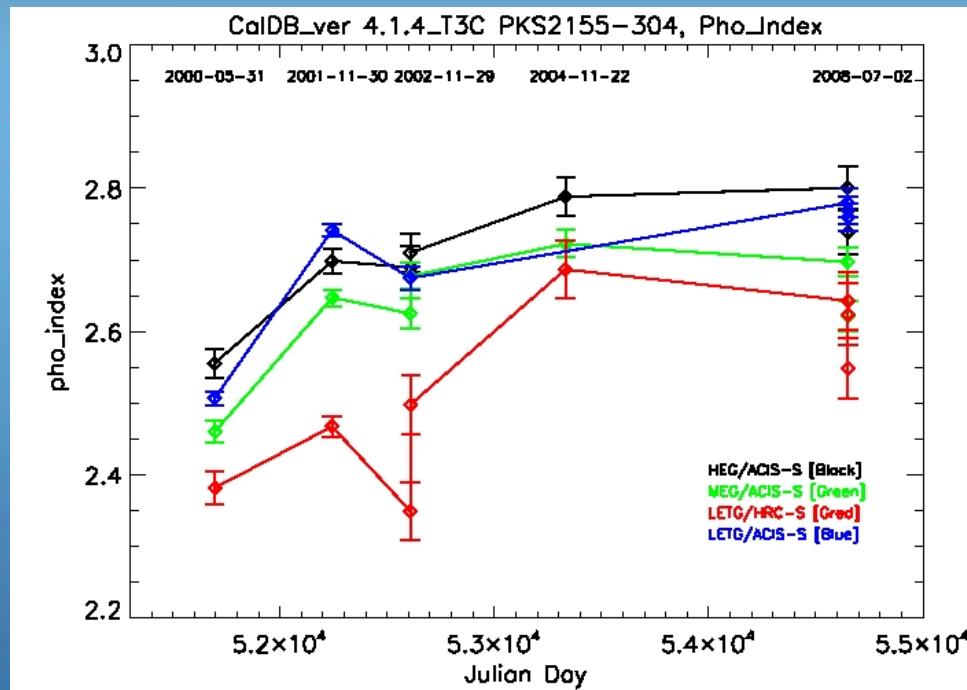
PKS 2155-304

Problems:
Variability

Multiple Configuration
Observations within
close temporal proximity



Fit nH-Absorbed PowerLaw



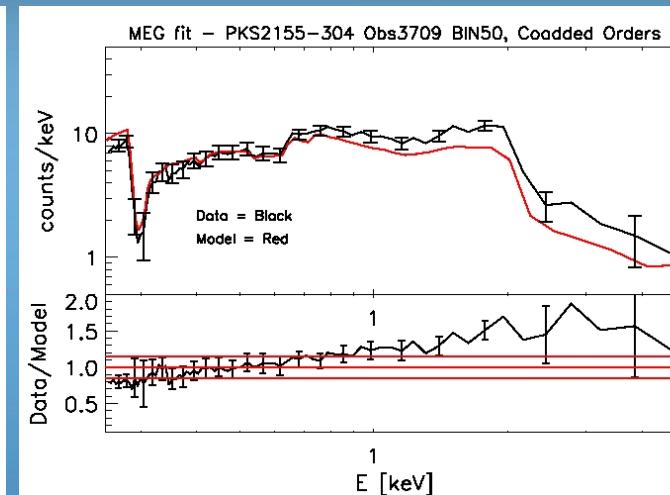
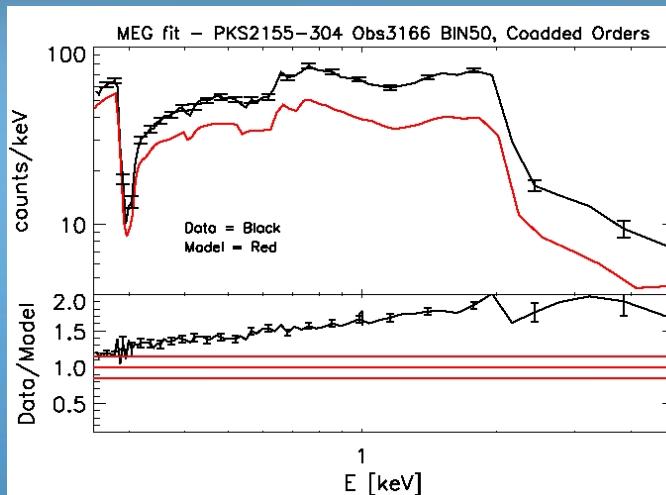
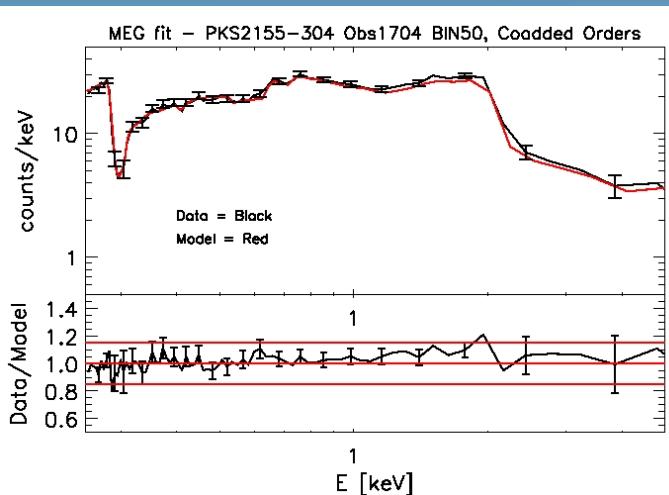
LEG PhoIndex < MEG

Inconsistent LEG-MEG
variance

PhoIndex sensitive to
variability

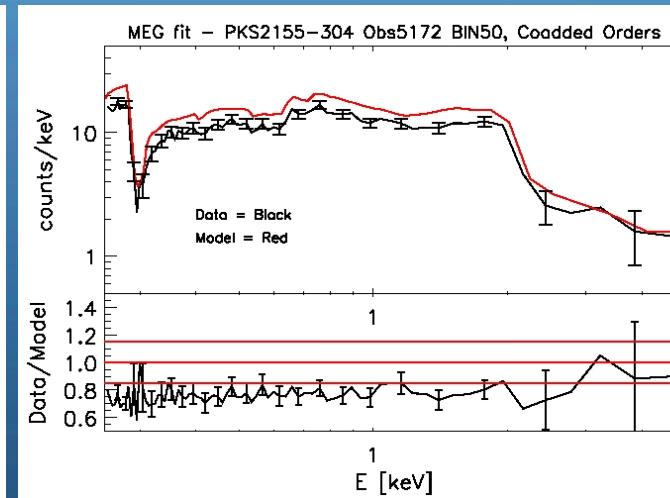
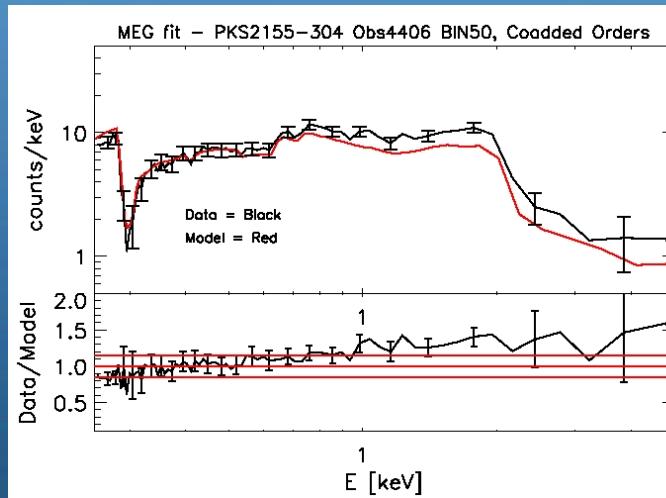
- Xspec & Sherpa
- Chi-Gehrels & Cash-Statistic
- Separate Fits for Positive / Negative Orders
- Energy Range: 0.8 – 5.0 keV
- $NH = 1.42e20 \text{ cm}^2$ <- LAB Survey

HRC-S/LEG Using ACIS-S/MEG Parameters



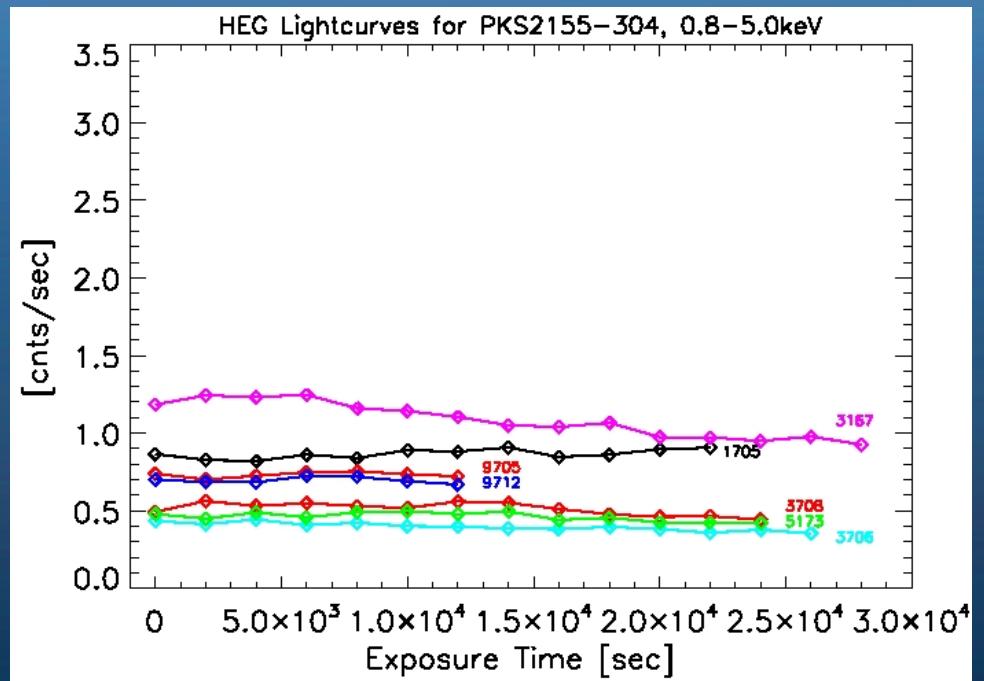
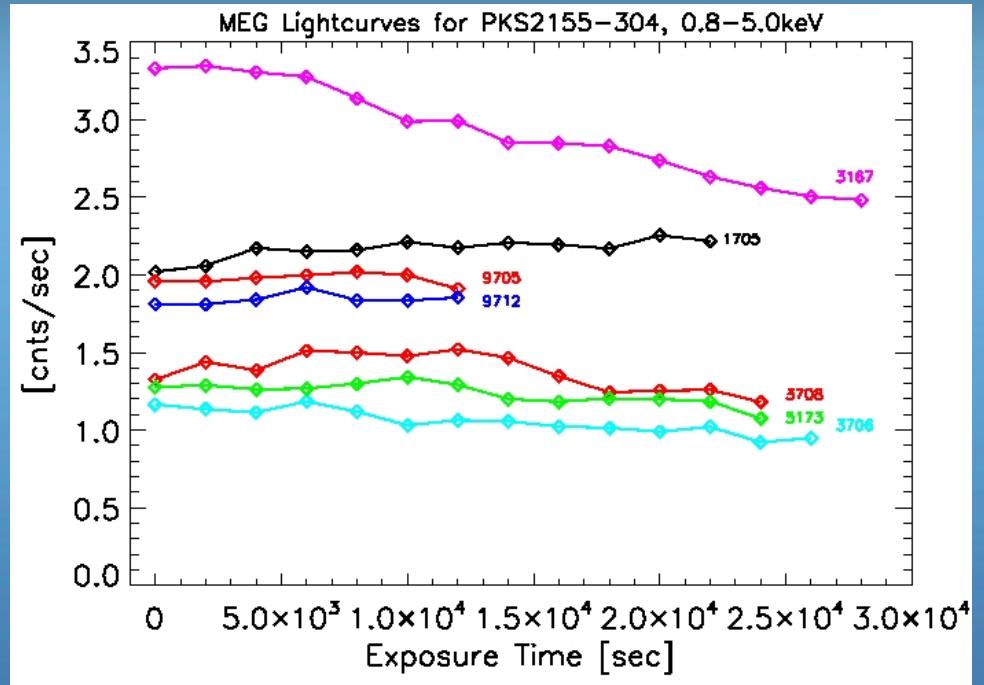
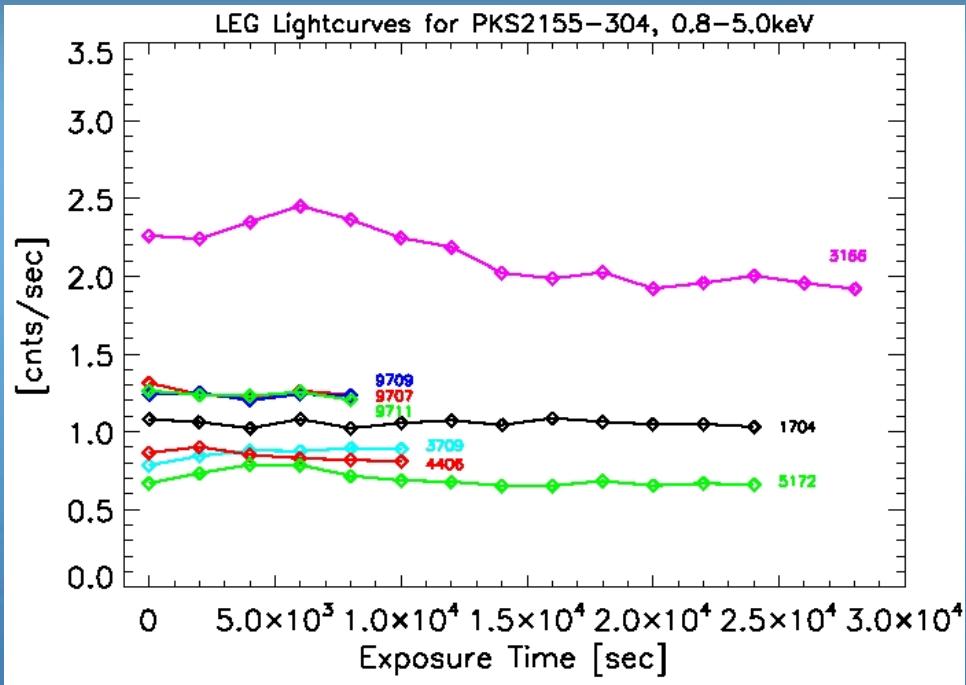
Bad Fits!

- MEG Pho_Index, Norm
- Coadded +/- Orders
- Residual Slope = Positive
- Due to LEG Pho_Index Systematics



**Residual Inconsistency
Due to Source
Variability**

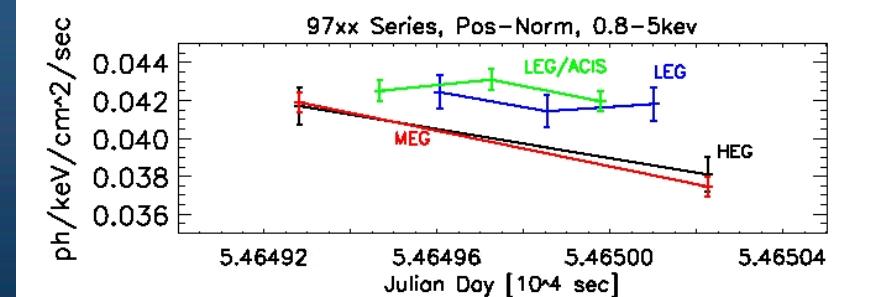
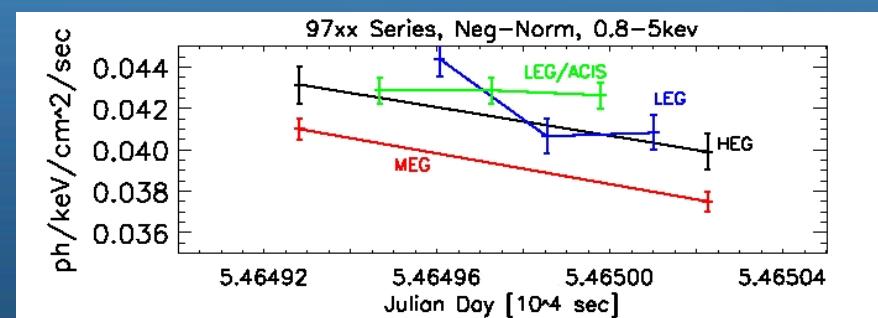
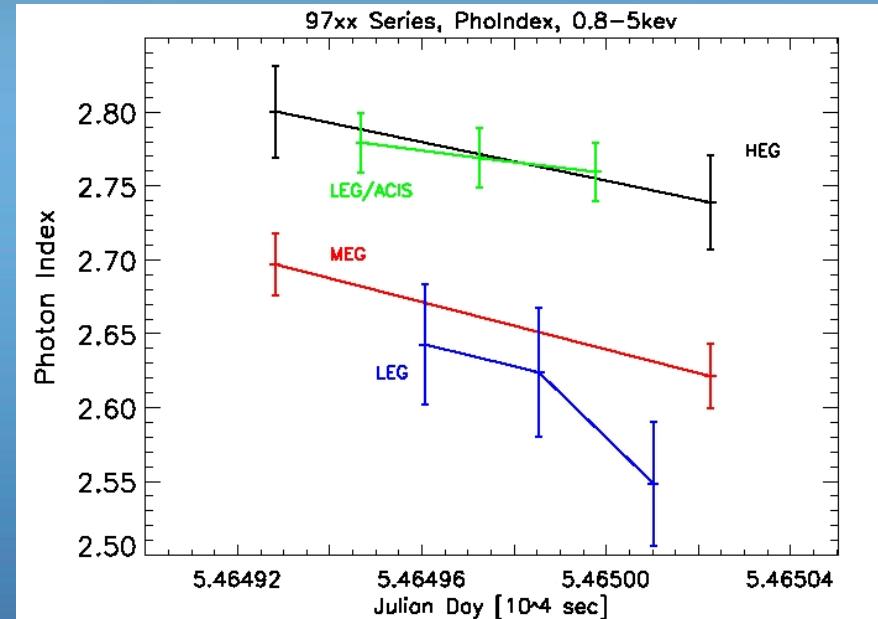
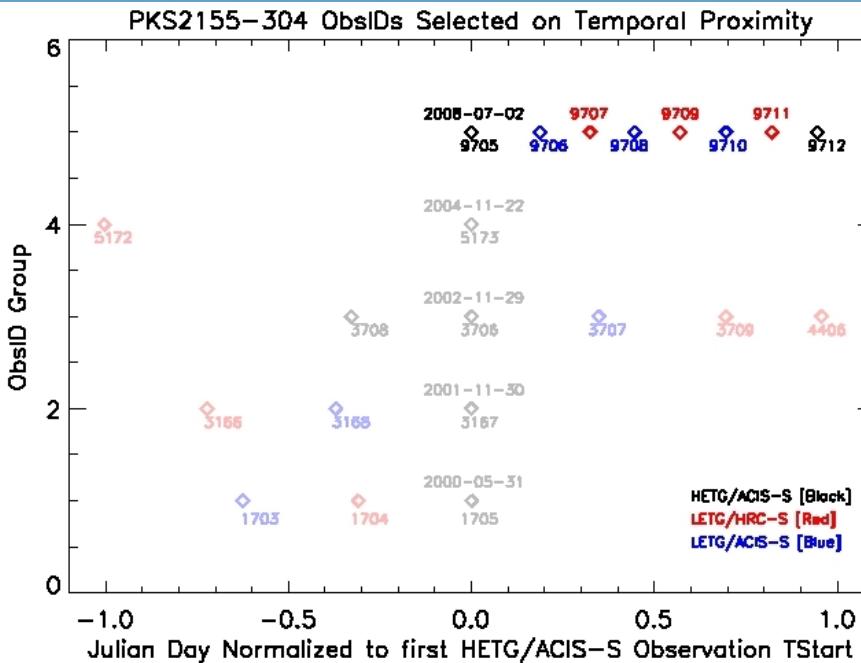
PKS 2155-304 Lightcurves



PKS 2155-304 = VARIABLE

$\Delta\text{Flux} >> \text{A Few Percent}$
is a Problem

Conquering Variability: July 2008 Series

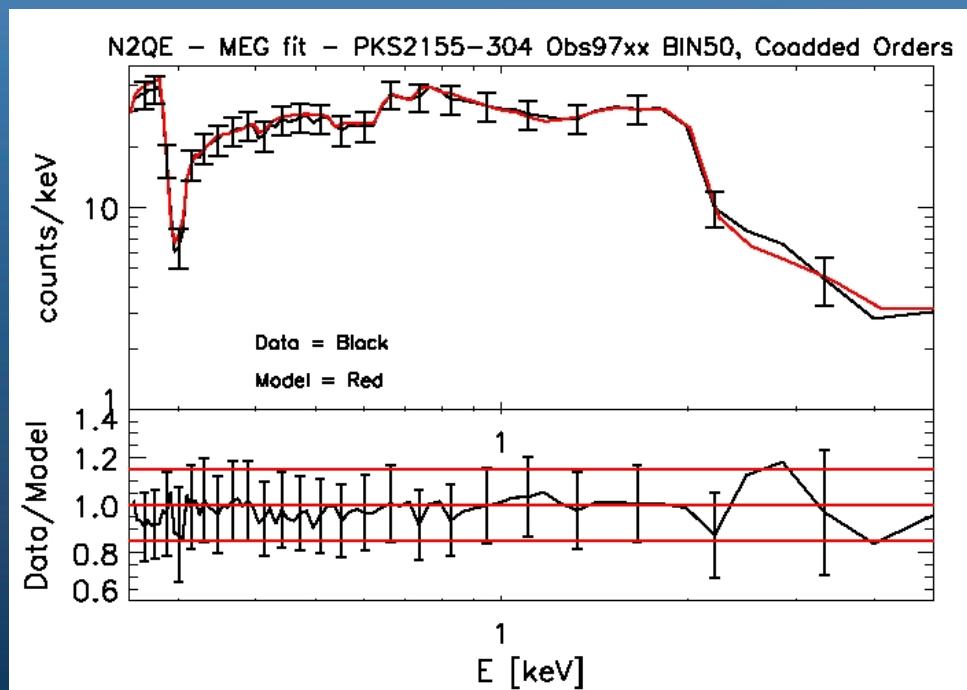
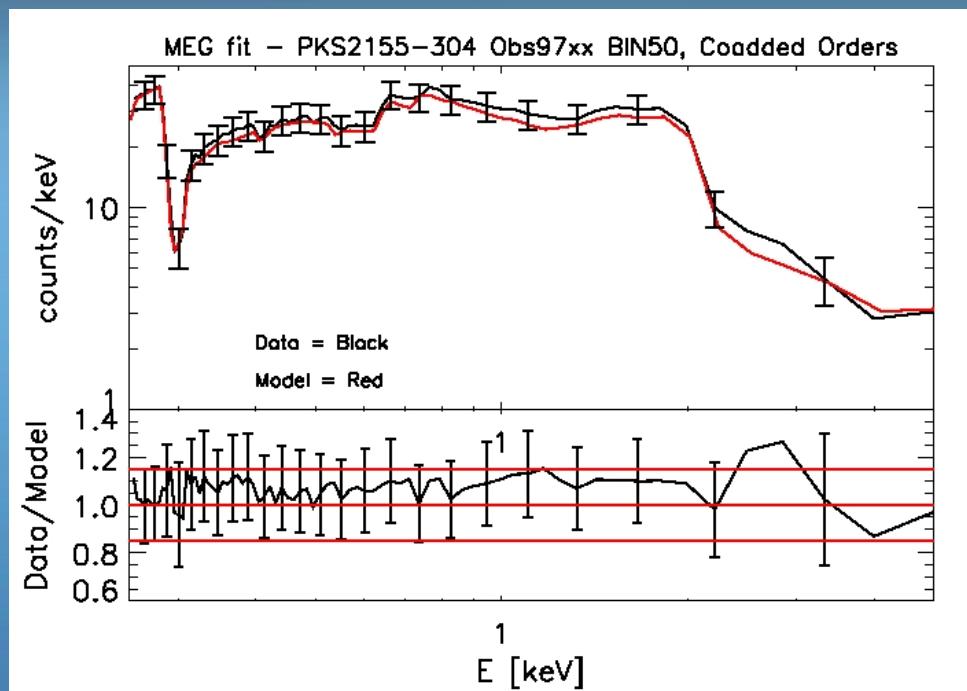
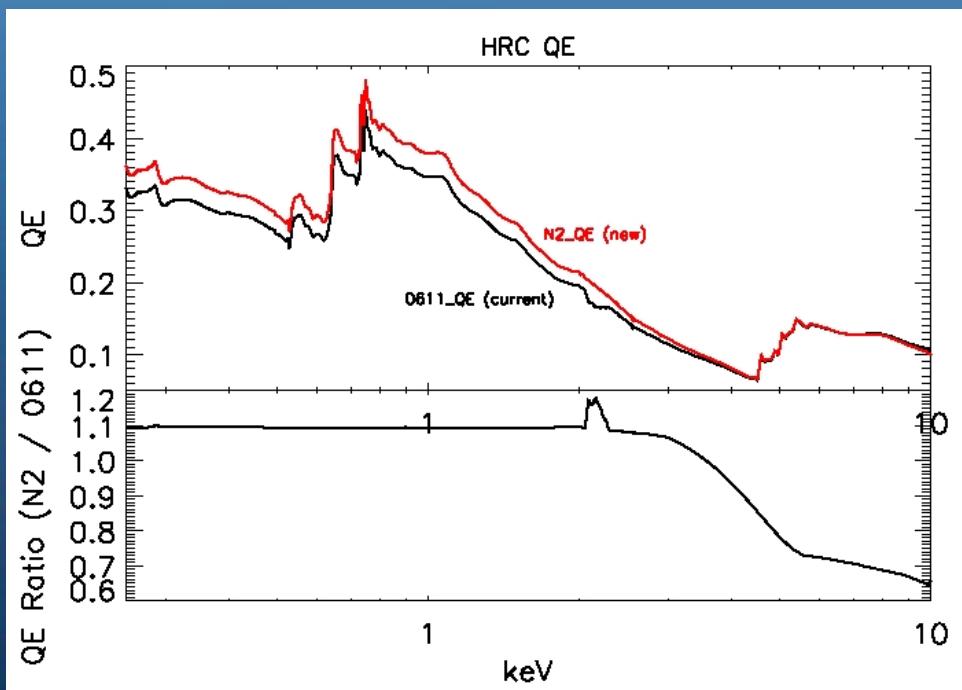


- 97xx Series
- MEG showing ~ monotonic declines
- Interpolate to LETG observation dates

QE Version - N2

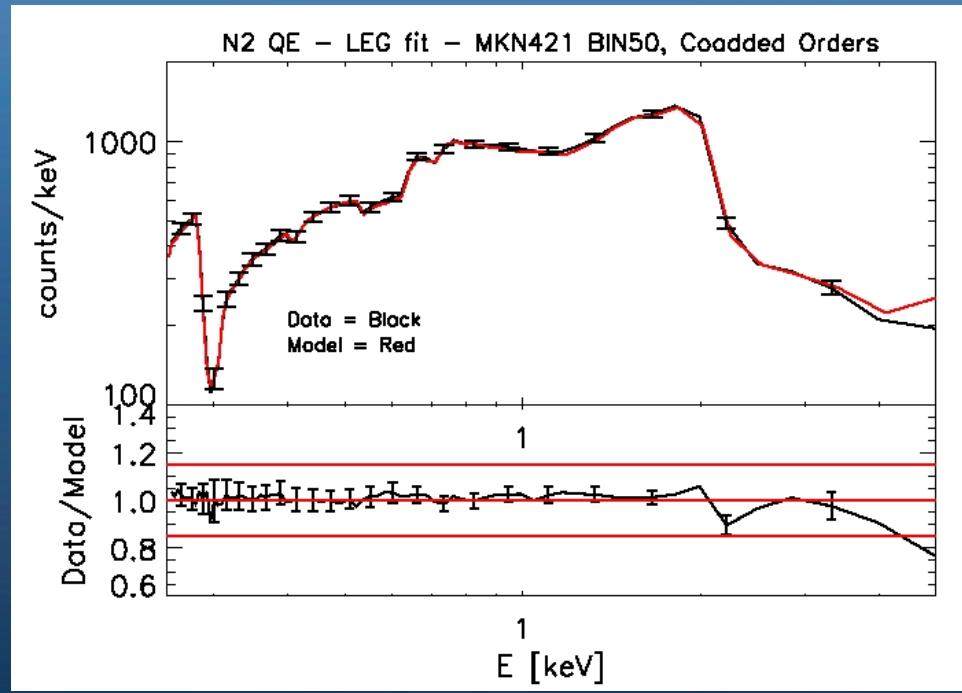
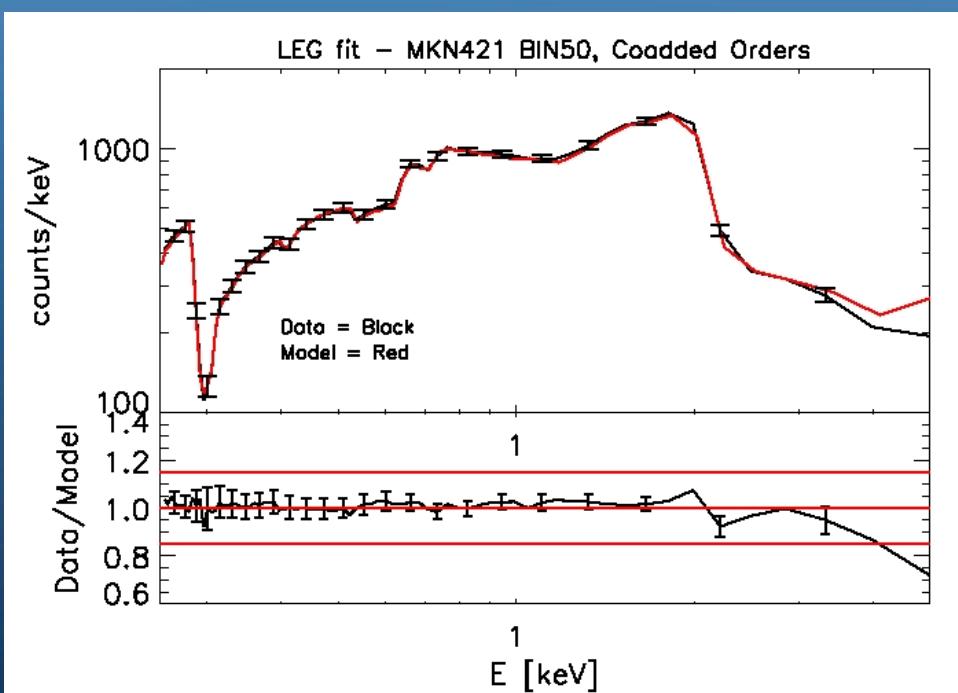
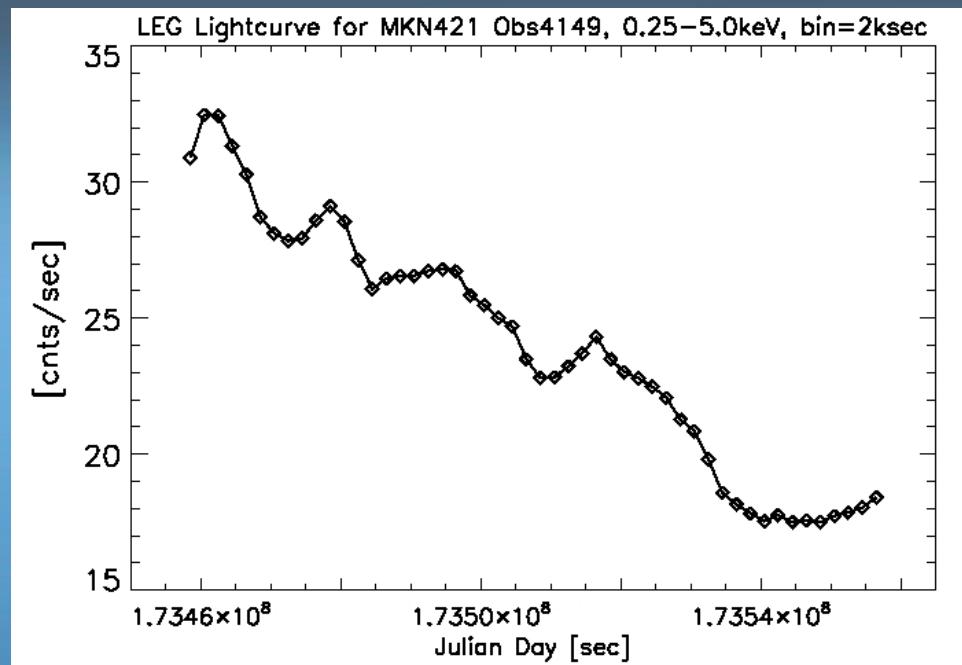
- Multiply QE by the change in the HRMA
- Smooth 2 keV bump

We arrive at our first QE revision

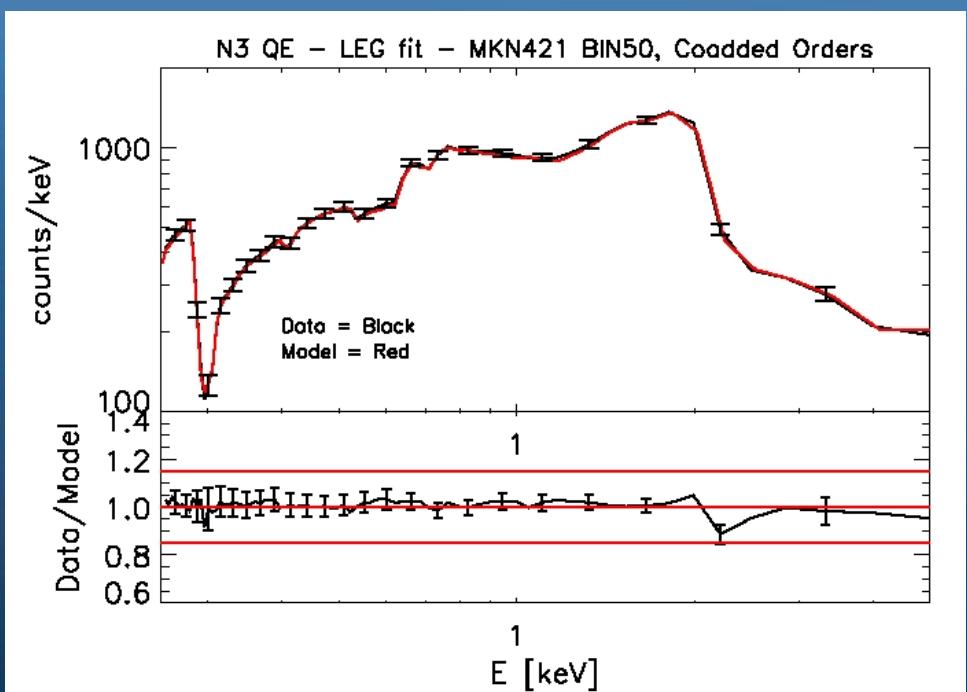
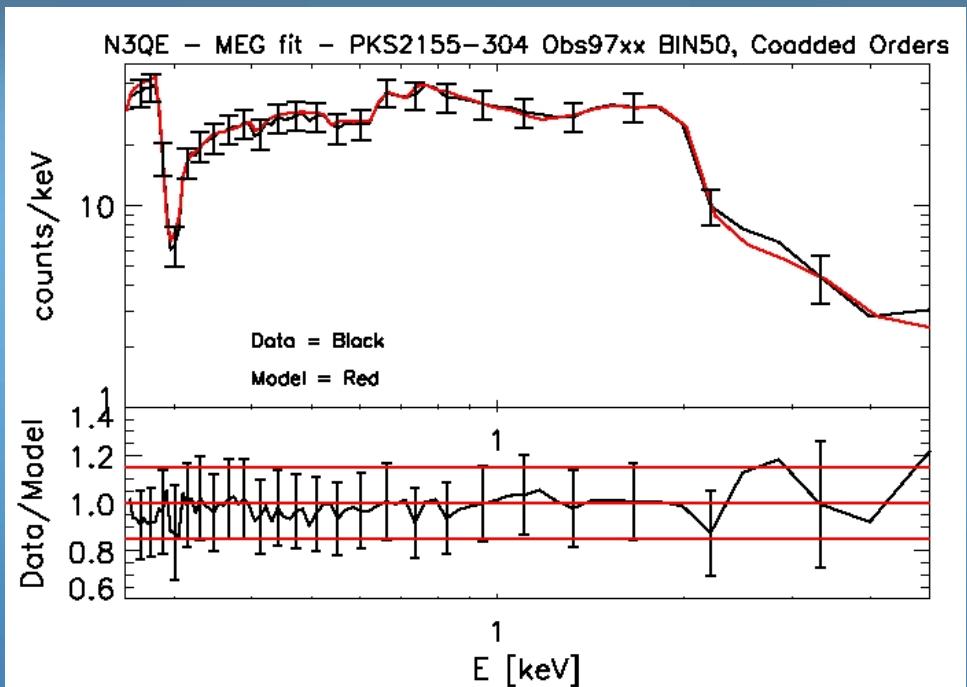
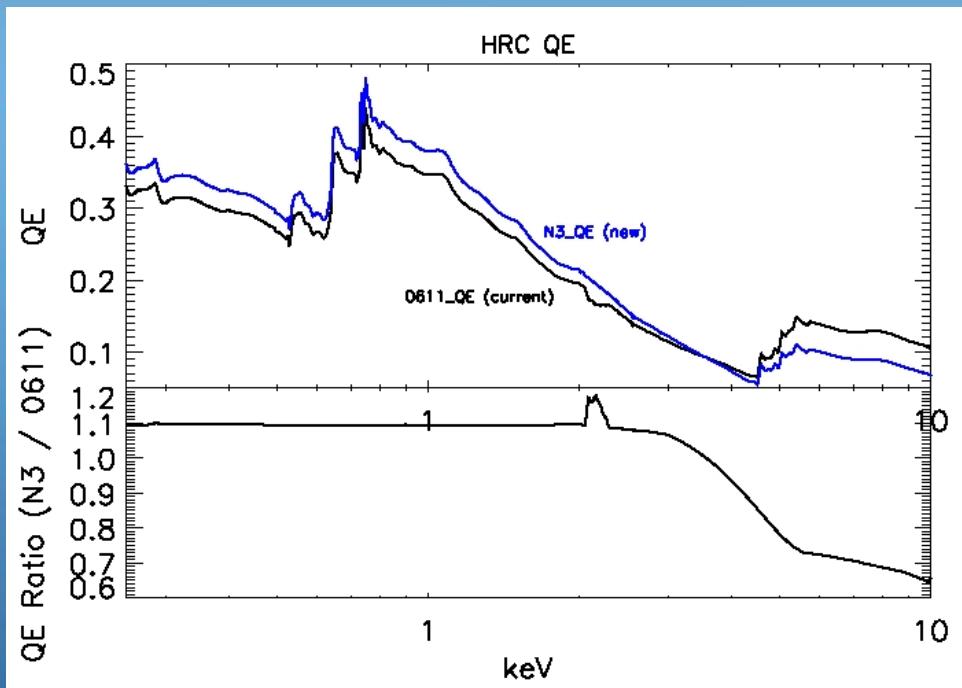


Serendipity! Mkn 421

- 100 ksec Observation (ObsID 4149)
- Variable LC:
- Split into four 25 ksec pieces
- Fit in XSpec – Absorbed PowerLaw Energy range 0.25 – 5.0 keV
- Fine QE correction



QE Version N3



- Attenuated at $E > 3$ keV
- Based on Mkn 421 smoothed residual fit
- Improvement in both PKS 2155-304 and Mkn 421

Future

- Further Revisions to our current prototype
- Release Data Late September / Early October
In time for POG
- Prototype Available for testing Early September
Prototype is usable in XSpec / Sherpa

Further Along

- Revisions Utilizing Simultaneous XMM Observations?