

Do Ultraluminous X-ray Sources Exist in Dwarf Galaxies?

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Do Ultraluminous X-ray Sources Exist in Dwarf Galaxies?

Dwarf Galaxies \longleftrightarrow Giant Galaxies

Star Formation

External Pressure \longleftrightarrow Density Waves, Bars

Metallicity

Low \longleftrightarrow High

Gas Fraction

High \longleftrightarrow Low

A CATALOG OF NEIGHBORING GALAXIES

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ABSTRACT

C, N, G CATALOG OF NEIGHBORING GALAXIES

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**$B_t < 17.5$ mag
 $D < 8$ Mpc**

ABSTRACT

TABLE I
CATALOG OF LV GALAXIES

Name (1)	R.A. (J2000.0) (2)	Decl. (J2000.0) (3)	α (4)	b/a (5)	B_r (6)	A_N (7)	T (8)	T_e (9)	V_k (10)	V_{LG} (11)	M_{50} (12)	$\log F$ (13)	D (14)	Method (15)
WLM, DDO 221	00 01 58.1	-15 27 40	11.5	0.35	11.03	0.16	9	A	-116	-10	57	2.30	0.92	rgb
E349-031, SDIG	00 08 13.3	-34 34 42	1.1	0.82	15.48	0.05	10	A	207	216	20	0.36	4.1	tf
N55	00 15 08.5	-39 13 13	32.4	0.17	8.84	0.06	8	A	129	111	172	3.43	1.8	tf
N59, KK 2	00 15 25.1	-21 26 38	2.7	0.48	12.97	0.09	-3	KB	361	431	50	0.43	5.30	sbf
E410-005, KK 3	00 15 31.4	-32 10 48	1.3	0.77	14.90	0.06	-1	A	1.92	rgb
I10	00 20 24.5	+59 17 30	6.8	0.87	12.20	3.65	10	A	-344	-60	63	2.98	0.66	cep
Sc 22	00 23 51.7	-24 42 18	0.9	0.78	17.73	0.06	-3	A	4.21	rgb
Cetus, KKSG 1	00 26 11.0	-11 02 40	5.0	0.86	14.4	0.12	-2	A	0.78	rgb
E294-010	00 26 33.3	-41 51 20	1.1	0.64	15.53	0.02	-3	A	117	81	1.92	rgb
U288	00 29 04.0	+43 25 54	1.3	0.62	15.64	0.33	10	A	187	464	36	0.71	6.7	bs
N147	00 33 11.6	+48 30 28	13.5	0.61	10.36	0.75	-3	A	-193	85	0.76	rgb
And III, KK 5	00 35 33.8	+36 29 52	3.0	0.60	15.20	0.10	-3	A	-355	-92	0.76	rgb
N185	00 38 58.0	+48 20 10	12.2	0.85	9.99	0.79	-3	A	-202	73	...	0.23	0.62	rgb
N205	00 40 22.5	+41 41 11	19.4	0.61	8.72	0.27	-5	K	-244	24	13	0.34	0.83	rgb
And IV	00 42 30.1	+40 34 33	1.3	0.77	16.60	0.27	10	A	256	521	6.11	rgb
N221	00 42 42.1	+40 51 59	9.0	0.71	8.73	0.26	-5	B	-145	121	0.77	rgb
M31	00 42 44.5	+41 16 09	189.0	0.33	4.17	0.27	3	B	-301	-35	510	4.53	0.77	cep
DDO 226	00 43 03.8	-22 15 01	2.2	0.36	14.36	0.07	10	A	357	408	48	1.01	4.92	rgb
And I, KK 8	00 45 40.0	+38 02 14	4.0	0.95	13.90	0.23	-3	A	-380	-120	0.81	rgb
N247	00 47 08.3	-20 45 36	21.4	0.32	9.86	0.08	7	K	160	215	210	2.83	4.1	tf
N253	00 47 34.3	-25 17 32	26.7	0.22	7.92	0.08	5	B	241	274	410	2.90	3.94	rgb
KDG 2, E540-030	00 49 21.1	-18 04 28	1.2	0.92	16.37	0.10	-1	A	3.40	rgb
DDO 6	00 49 49.3	-21 00 58	1.7	0.41	15.19	0.07	10	A	295	348	22	0.53	3.34	rgb
E540-032, FG 24	00 50 24.6	-19 54 25	1.3	0.92	16.44	0.09	-3	A	3.42	rgb
SMC	00 52 38.0	-72 48 01	319	0.64	2.75	0.17	9	A	158	-22	90	4.95	0.06	cep
N300	00 54 53.5	-37 40 57	21.9	0.71	8.95	0.06	7	K	144	114	149	3.31	2.15	cep
Sculptor	01 00 09.4	-33 42 33	35.2	0.79	10.05	0.05	-3	A	110	96	...	0.11	0.09	rgb
LGS-3	01 03 56.6	+21 53 41	1.2	0.58	16.18	0.18	-1	A	-286	-74	18	0.23	0.62	rgb
I1613	01 04 54.1	+02 07 60	16.4	0.90	9.92	0.11	10	A	-232	-89	25	2.67	0.73	cep

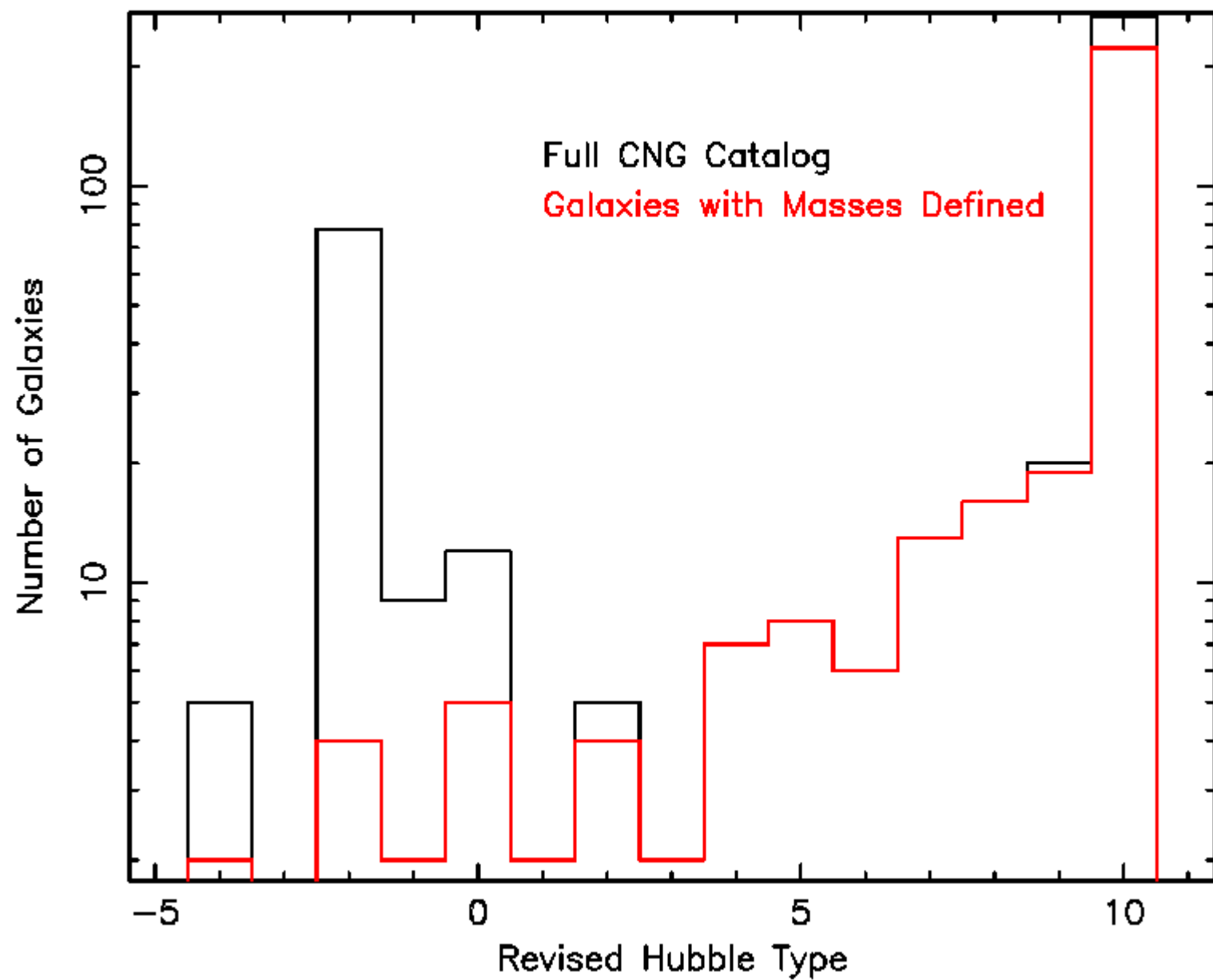
TABLE I
CATALOG OF LV GALAXIES

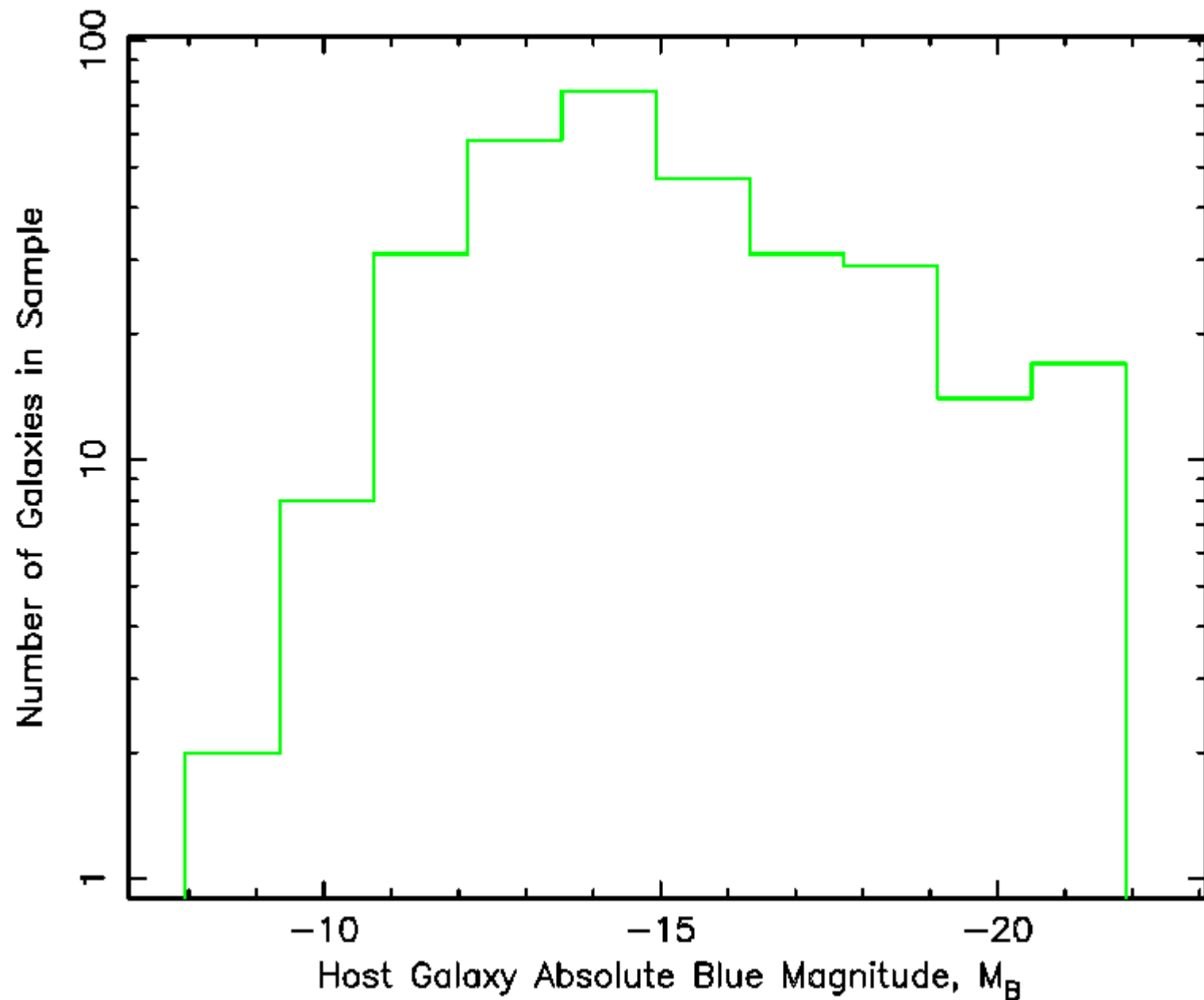
Name (1)	R.A. (J2000.0) (2)	Decl. (J2000.0) (3)	a (4)	b/a (5)	B_r (6)	A_V (7)	T (8)	T_e (9)	Y_k (10)	Y_{LG} (11)	M_{50} (12)	$\log F$ (13)	D (14)	Method (15)
WLM, DDO 221	00 01 58.1	-15 27 40	11.5	0.35	11.03	0.16	9	A	-116	-10	57	2.30	0.92	rgb
E349-031, SDIG	00 08 13.3	-34 34 42	1.1	0.82	15.48	0.05	10	A	207	216	20	0.36	4.1	tf
N55	00 15 08.5	-39 13 13	32.4	0.17	8.84	0.06	8	A	129	111	172	3.43	1.8	tf
N59, KK 2	00 15 25.1	-21 26 38	2.7	0.48	12.97	0.09	-3	KB	361	431	50	0.43	5.30	sbf
E410-005, KK 3	00 15 31.4	-32 10 48	1.3	0.77	14.90	0.06	-1	A	1.92	rgb
I10	00 20 24.5	+59 17 30	6.8	0.87	12.20	3.65	10	A	-344	-60	63	2.98	0.66	cep
Sc 22	00 23 51.7	-24 42 18	0.9	0.78	17.73	0.06	-3	A	4.21	rgb
Cetus, KK9G 1	00 26 11.0	-11 02 40	5.0	0.86	14.4	0.12	-2	A	0.78	rgb
E294-010	00 26 33.3	-41 51 20	1.1	0.64	15.53	0.02	-3	A	117	81	1.92	rgb
U288	00 29 04.0	+43 25 54	1.3	0.62	15.64	0.33	10	A	187	464	36	0.71	6.7	bs
N147	00 33 11.6	+48 30 28	13.5	0.61	10.36	0.75	-3	A	-193	85	0.76	rgb
And III, KK 5	00 35 33.8	+36 29 52	3.0	0.60	15.20	0.10	-3	A	-355	-92	0.76	rgb
N185	00 38 58.0	+48 20 10	12.2	0.85	9.99	0.79	-3	A	-202	73	...	0.23	0.62	rgb
N205	00 40 22.5	+41 41 11	19.4	0.61	8.72	0.27	-5	K	-244	24	13	0.34	0.83	rgb
And IV	00 42 30.1	+40 34 33	1.3	0.77	16.60	0.27	10	A	256	521	6.11	rgb
N221	00 42 42.1	+40 51 59	9.0	0.71	8.73	0.26	-5	B	-145	121	0.77	rgb
M31	00 42 44.5	+41 16 09	189.0	0.33	4.17	0.27	3	B	-301	-35	510	4.53	0.77	cep
DDO 226	00 43 03.8	-22 15 01	2.2	0.36	14.36	0.07	10	A	357	408	48	1.01	4.92	rgb
And I, KK 8	00 45 40.0	+38 02 14	4.0	0.95	13.90	0.23	-3	A	-380	-120	0.81	rgb
N247	00 47 08.3	-20 45 36	21.4	0.32	9.86	0.08	7	K	160	215	210	2.83	4.1	tf
N253	00 47 34.3	-25 17 32	26.7	0.22	7.92	0.08	5	B	241	274	410	2.90	3.94	rgb
KDG 2, E540-030	00 49 21.1	-18 04 28	1.2	0.92	16.37	0.10	-1	A	3.40	rgb
DDO 6	00 49 49.3	-21 00 58	1.7	0.41	15.19	0.07	10	A	295	348	22	0.53	3.34	rgb
E540-032, FG 24	00 50 24.6	-19 54 25	1.3	0.92	16.44	0.09	-3	A	3.42	rgb
SMC	00 52 38.0	-72 48 01	319	0.64	2.75	0.17	9	A	158	-22	90	4.95	0.06	cep
N300	00 54 53.5	-37 40 57	21.9	0.71	8.95	0.06	7	K	144	114	149	3.31	2.15	cep
Sculptor	01 00 09.4	-33 42 33	35.2	0.79	10.05	0.05	-3	A	110	96	...	0.11	0.09	rgb
LGS-3	01 03 56.6	+21 53 41	1.2	0.58	16.18	0.18	-1	A	-286	-74	18	0.23	0.62	rgb
I1613	01 04 54.1	+02 07 60	16.4	0.90	9.92	0.11	10	A	-232	-89	25	2.67	0.73	cep

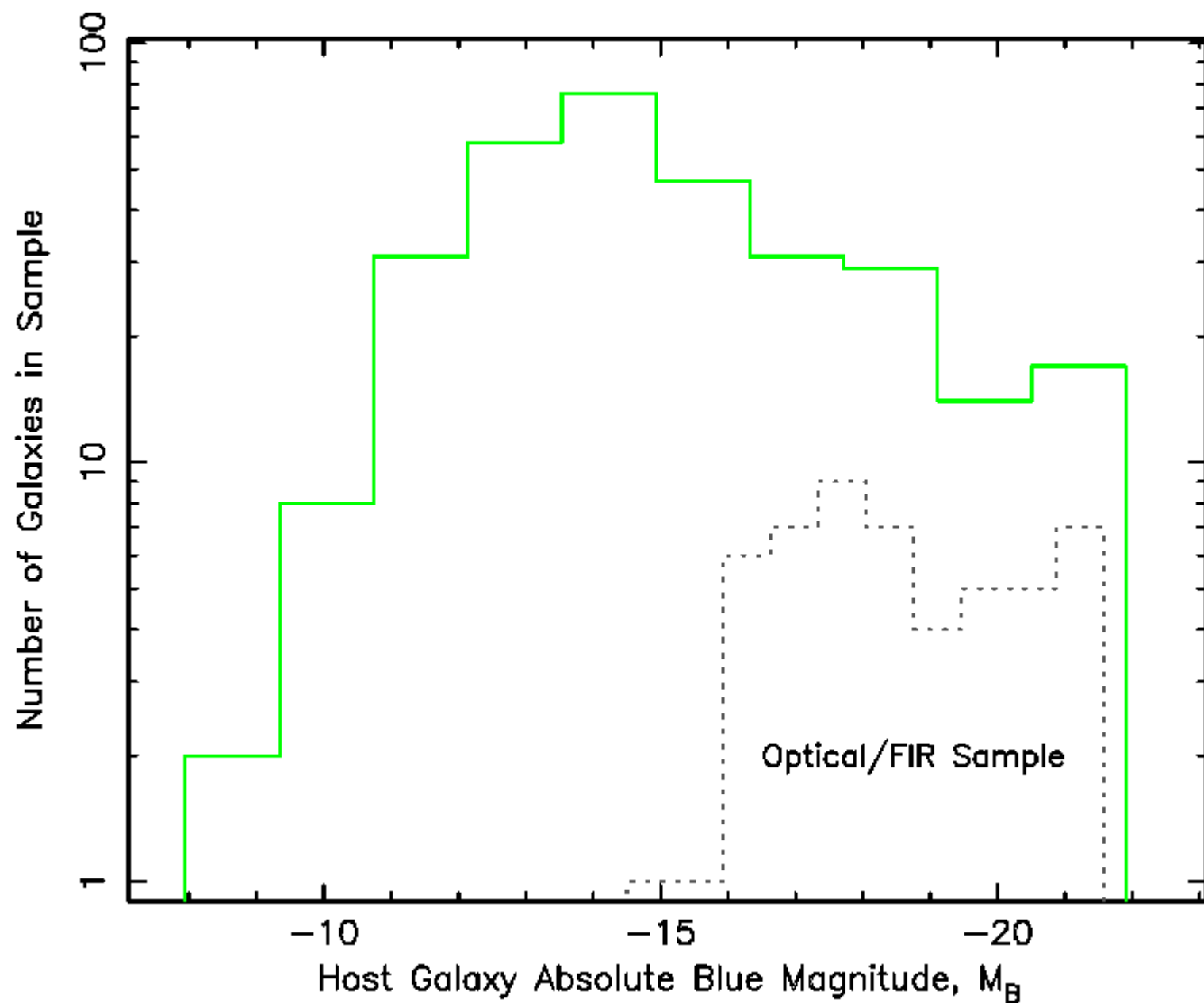
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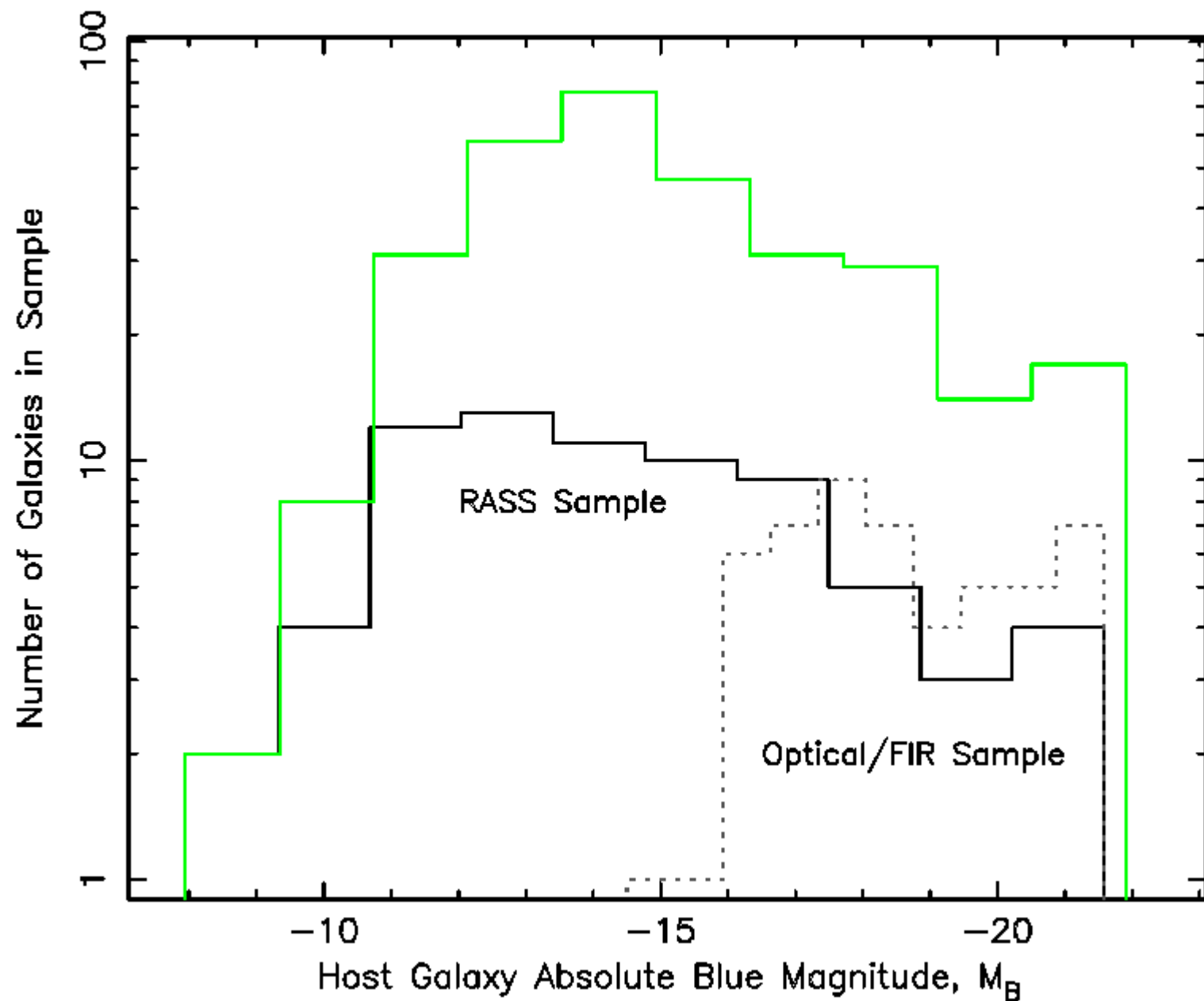
Name (1)	R.A. (J2000.0) (2)	Decl. (J2000.0) (3)	a (4)	b/a (5)	B_i (6)	A_V (7)	T (8)	T_e (9)	Y_k (10)	Y_{LG} (11)	M_{50} (12)	$\log F$ (13)	D (14)	Method (15)
WLM, DDO 221	00 01 58.1	-15 27 40	11.5	0.35	11.03	0.16	9	A	-116	-10	57	2.30	0.92	rgb
E349-031, SDIG	00 08 13.3	-34 34 42	1.1	0.82	15.48	0.05	10	A	207	216	20	0.36	4.1	tf
N55	00 15 08.5	-39 13 13	32.4	0.17	8.84	0.06	8	A	129	111	172	3.43	1.8	tf
N59, KK 2	00 15 25.1	-21 26 38	2.7	0.48	12.97	0.09	-3	KB	361	431	50	0.43	5.30	sbf
E410-005, KK 3	00 15 31.4	-32 10 48	1.3	0.77	14.90	0.06	-1	A	1.92	rgb
110	00 20 24.5	+59 17 30	6.8	0.87	12.20	3.65	10	A	-344	-60	63	2.98	0.66	cep
Sc 22	00 23 51.7	-24 42 18	0.9	0.78	17.73	0.06	-3	A	4.21	rgb
Cetus, KK9G 1	00 26 11.0	-11 02 40	5.0	0.86	14.4	0.12	-2	A	0.78	rgb
E294-010	00 26 33.3	-41 51 20	1.1	0.64	15.53	0.02	-3	A	117	81	1.92	rgb
U288	00 29 04.0	+43 25 54	1.3	0.62	15.64	0.33	10	A	187	464	36	0.71	6.7	bs
N147	00 33 11.6	+48 30 28	13.5	0.61	10.36	0.75	-3	A	-193	85	0.76	rgb
And III, KK 5	00 35 29.8	+56 42 52	3.0	0.60	12.20	0.10	9	A	-355	-92	0.76	rgb
N185	00 38 21.1	+8 21 10	9.5	0.9	9.5	0.9	9	A	-202	73	0.62	rgb
N205	00 40 22.5	-41 41 11	19.4	0.61	8.7	0.17	5	K	-244	24	13	0.34	0.83	rgb
And IV	00 42 30.1	+40 34 33	1.3	0.77	16.60	0.27	10	A	256	521	6.11	rgb
N221	00 42 42.1	+40 51 59	9.0	0.71	8.73	0.26	-5	B	-145	121	0.77	rgb
M31	00 42 44.5	+41 16 09	189.0	0.33	4.17	0.27	3	B	-301	-35	510	4.53	0.77	cep
DDO 226	00 43 03.8	-22 15 01	2.2	0.36	14.36	0.07	10	A	357	408	48	1.01	4.92	rgb
And I, KK 8	00 45 40.0	+38 02 14	4.0	0.95	13.90	0.23	-3	A	-380	-120	0.81	rgb
N247	00 47 08.3	-20 45 36	21.4	0.32	9.86	0.08	7	K	160	215	210	2.83	4.1	tf
N253	00 47 34.3	-25 17 32	26.7	0.22	7.92	0.08	5	B	241	274	410	2.90	3.94	rgb
KDG 2, E540-030	00 49 21.1	-18 04 28	1.2	0.92	16.37	0.10	-1	A	3.40	rgb
DDO 6	00 49 49.3	-21 00 58	1.7	0.41	15.19	0.07	10	A	295	348	22	0.53	3.34	rgb
E540-032, FG 24	00 50 24.6	-19 54 25	1.3	0.92	16.44	0.09	-3	A	3.42	rgb
SMC	00 52 38.0	-72 48 01	319	0.64	2.75	0.17	9	A	158	-22	90	4.95	0.06	cep
N300	00 54 53.5	-37 40 57	21.9	0.71	8.95	0.06	7	K	144	114	149	3.31	2.15	cep
Sculptor	01 00 09.4	-33 42 33	35.2	0.79	10.05	0.05	-3	A	110	96	...	0.11	0.09	rgb
LGS-3	01 03 56.6	+21 53 41	1.2	0.58	16.18	0.18	-1	A	-286	-74	18	0.23	0.62	rgb
11613	01 04 54.1	+02 07 60	16.4	0.90	9.92	0.11	10	A	-232	-89	25	2.67	0.73	cep

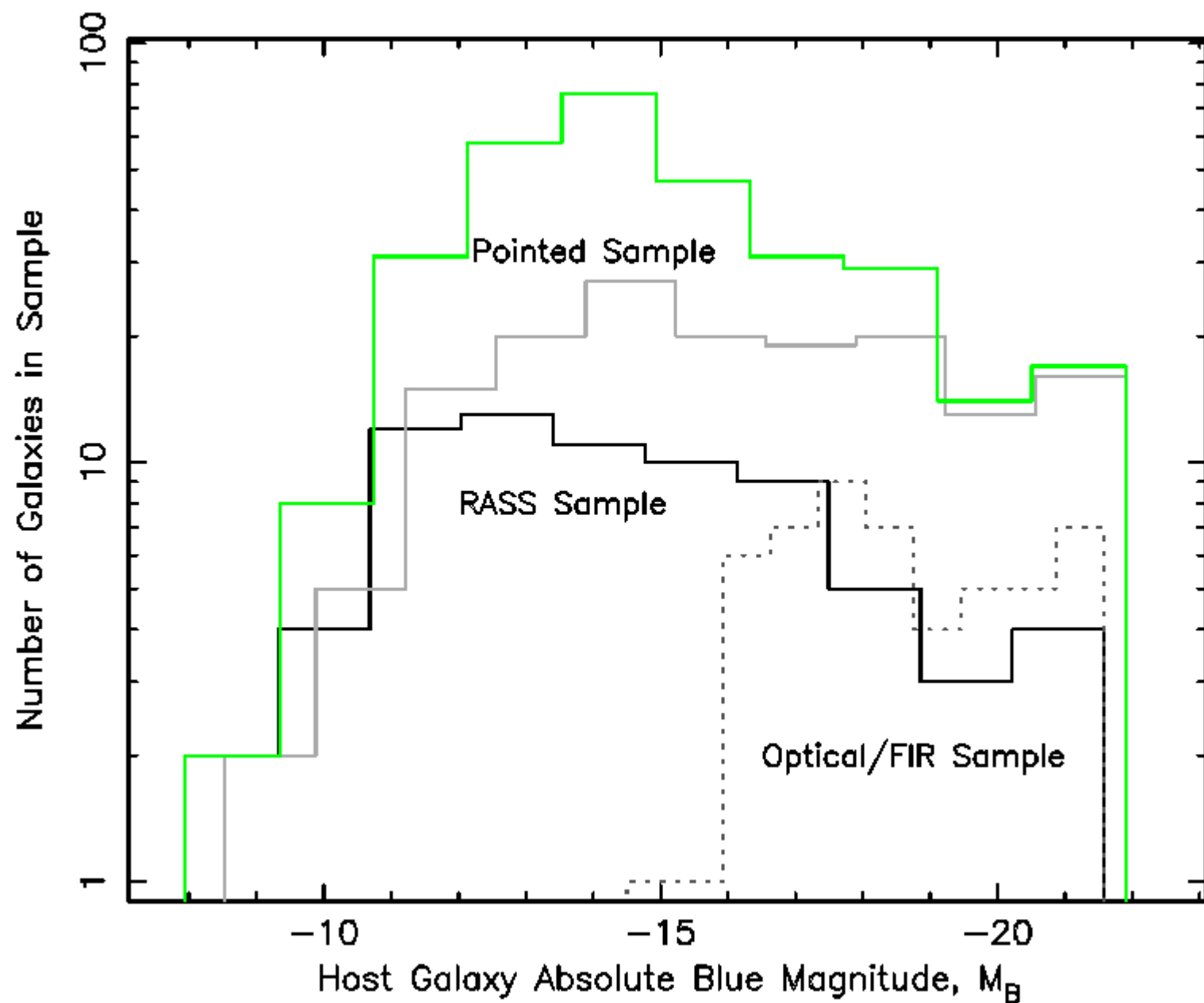
451 → 313

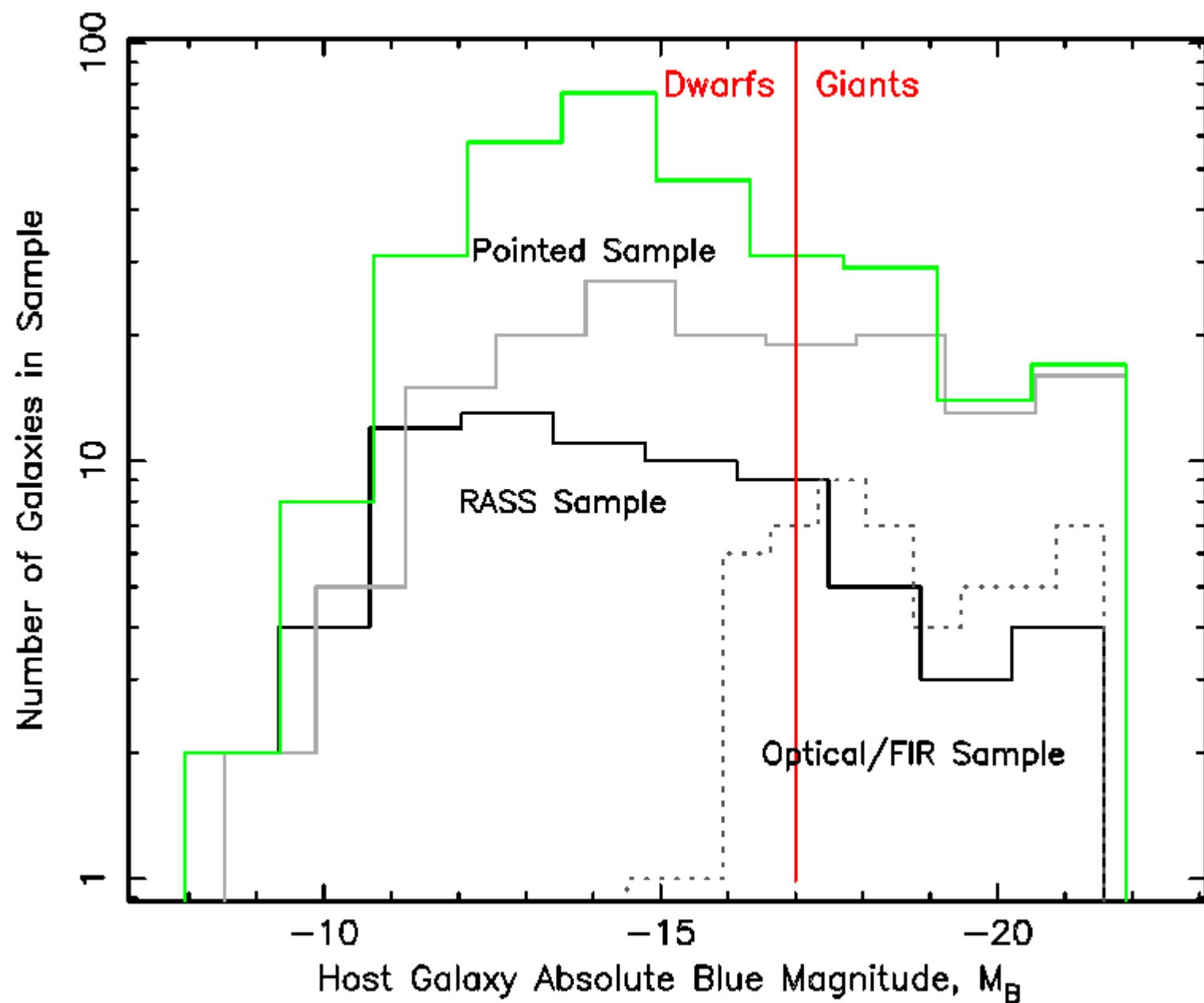


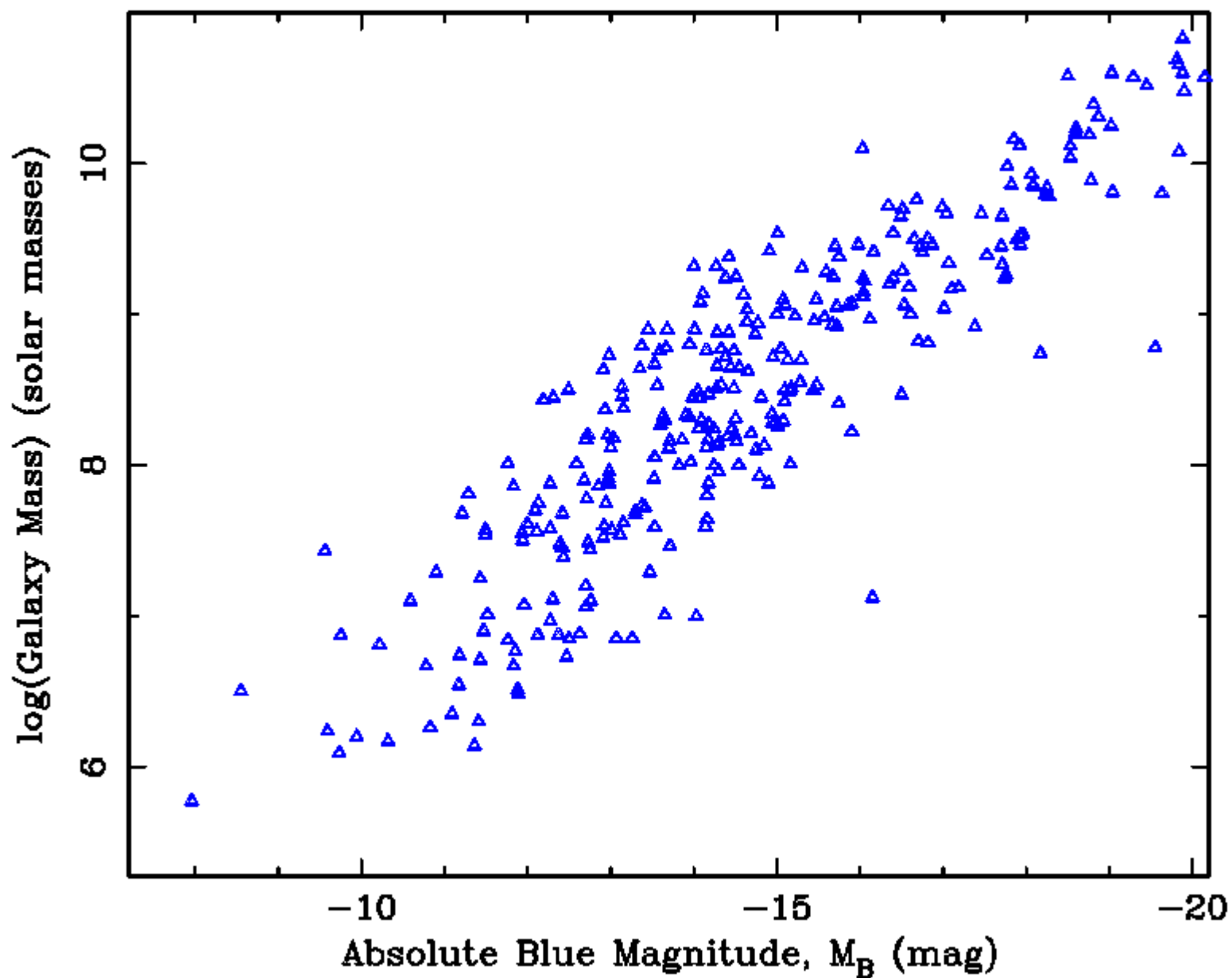


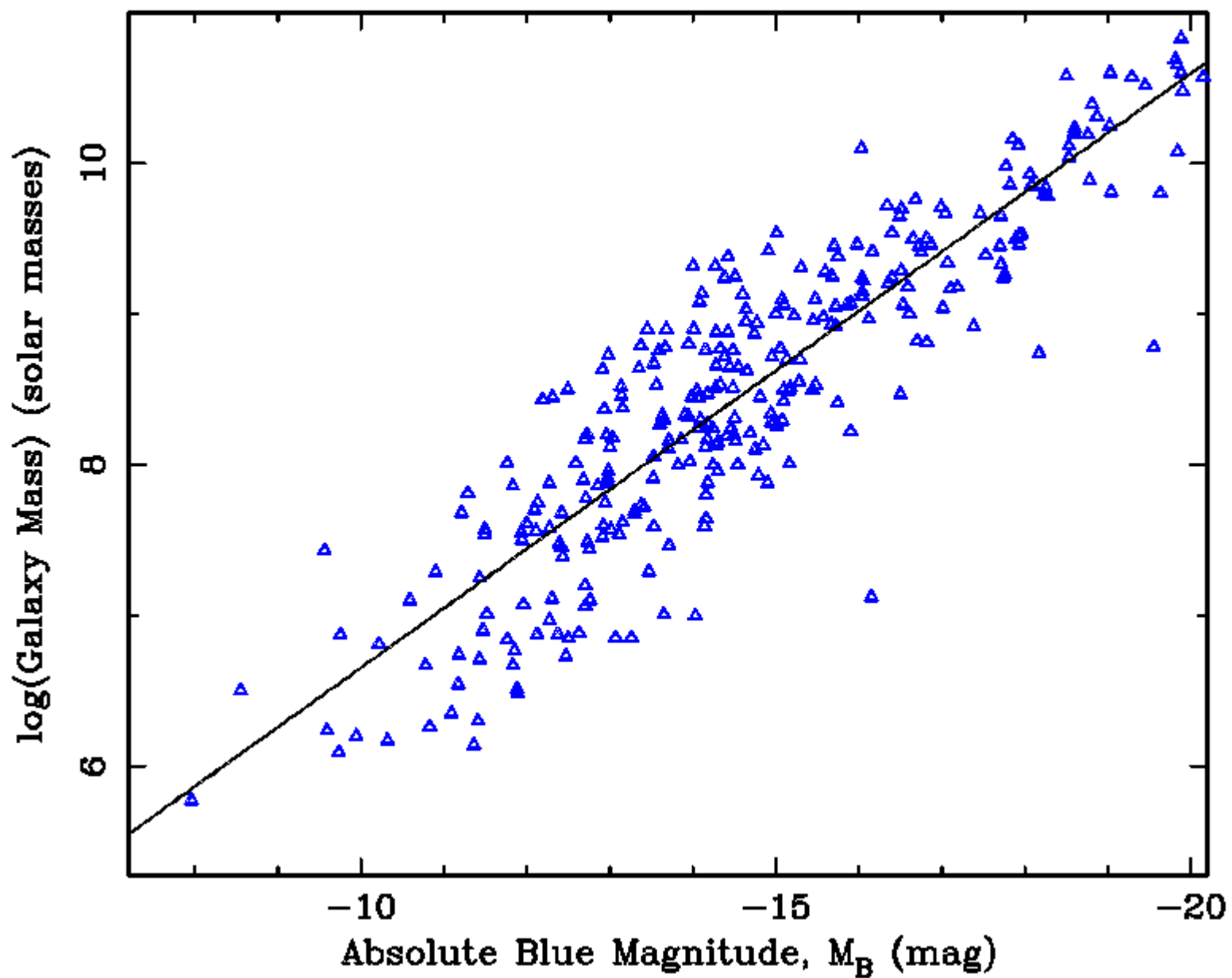


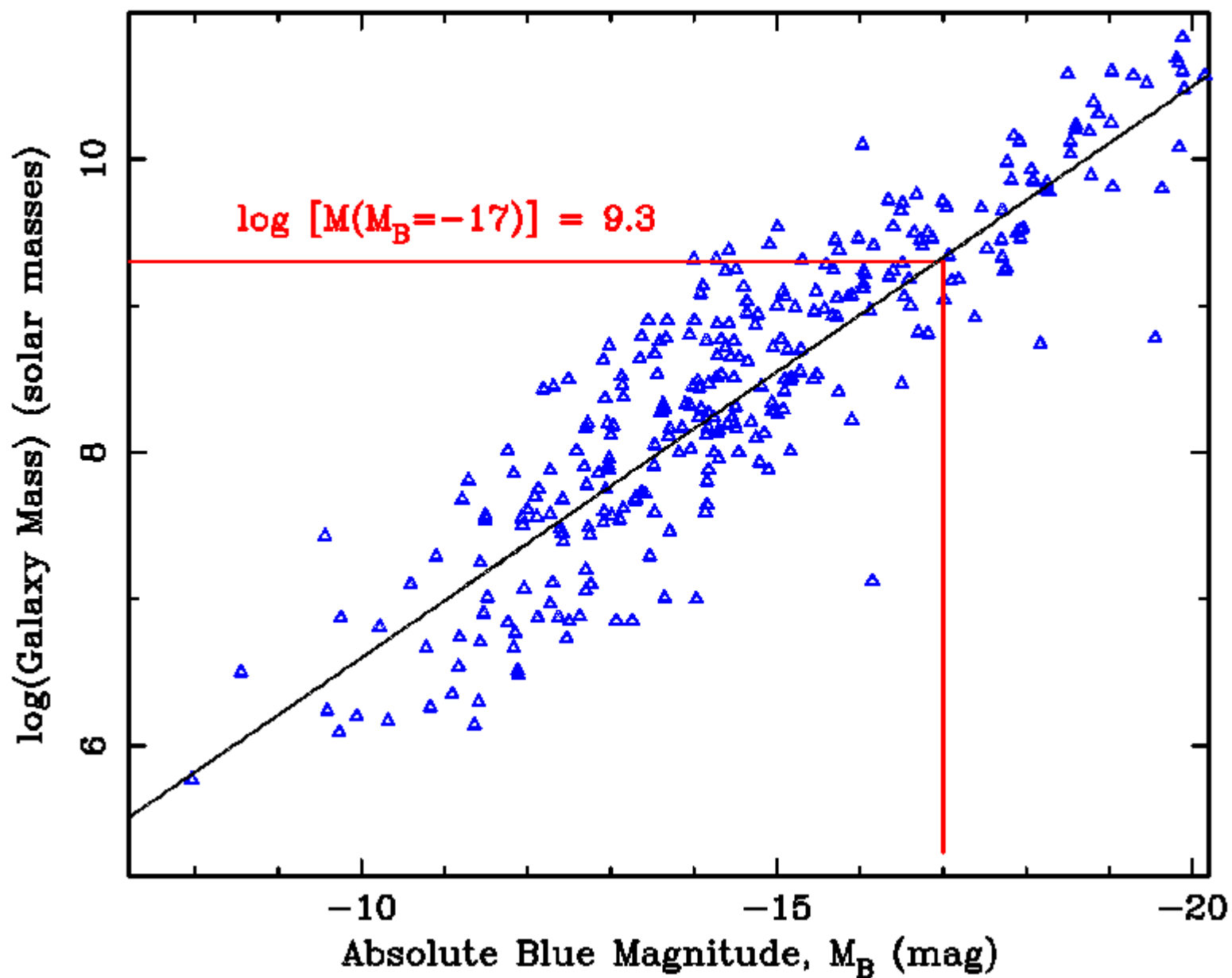




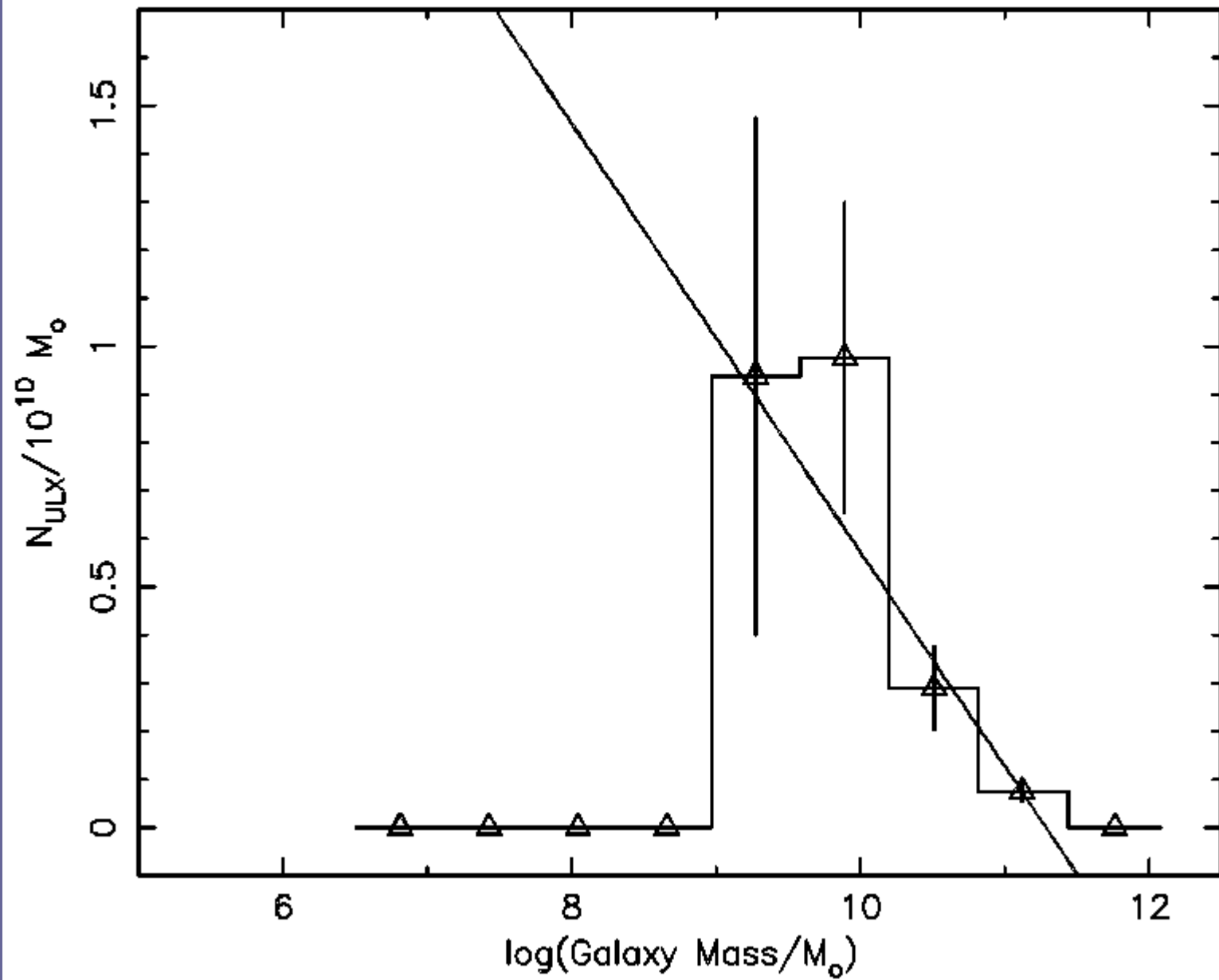




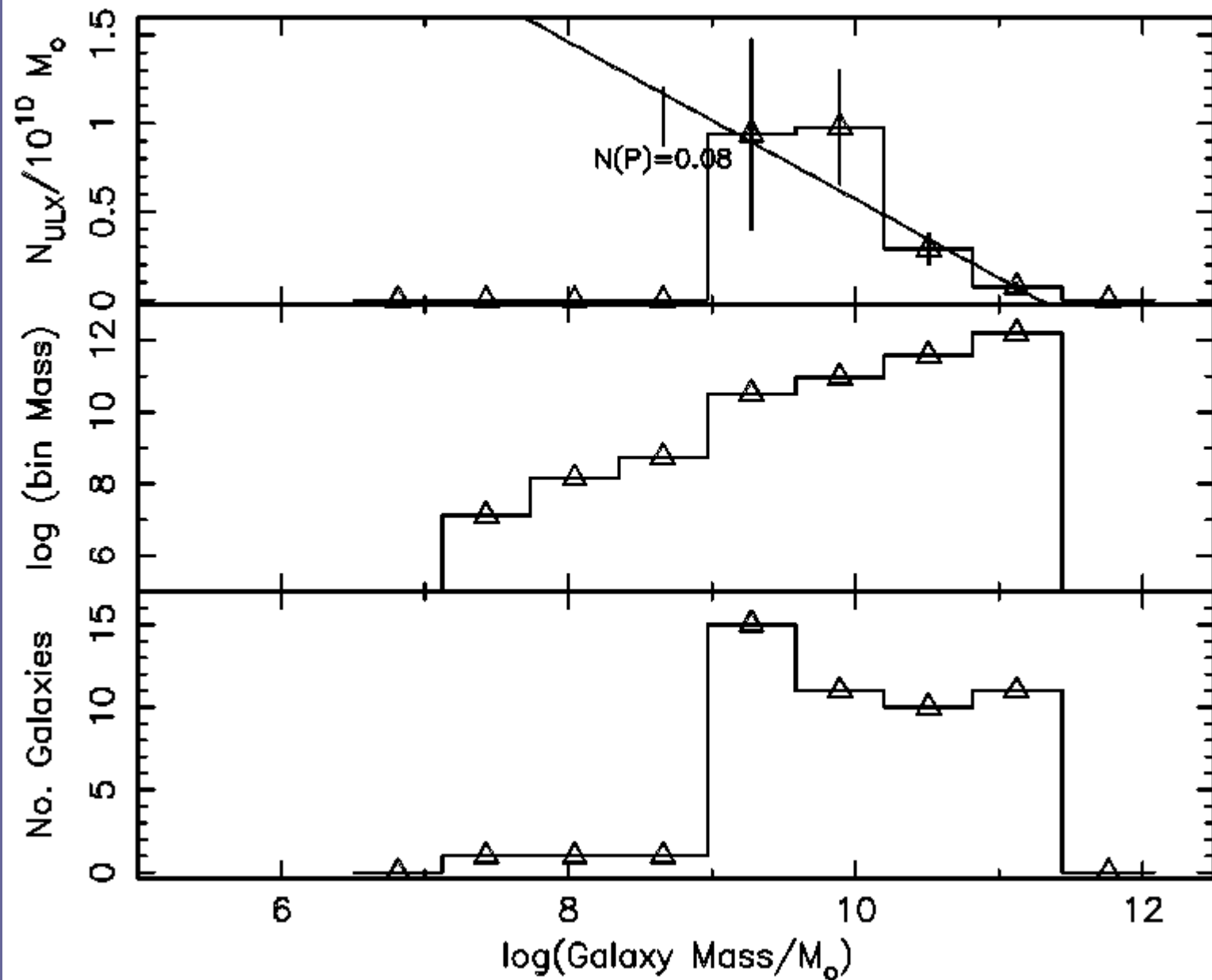




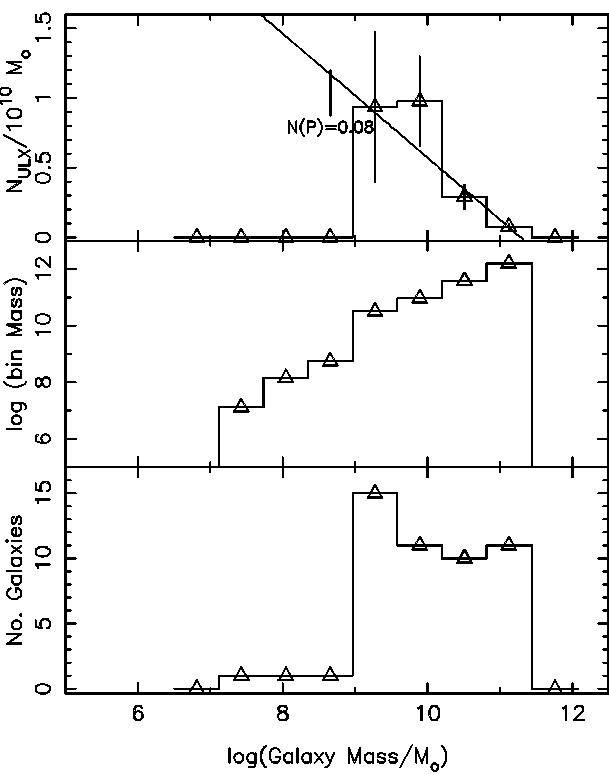
Optical/FIR Subsample



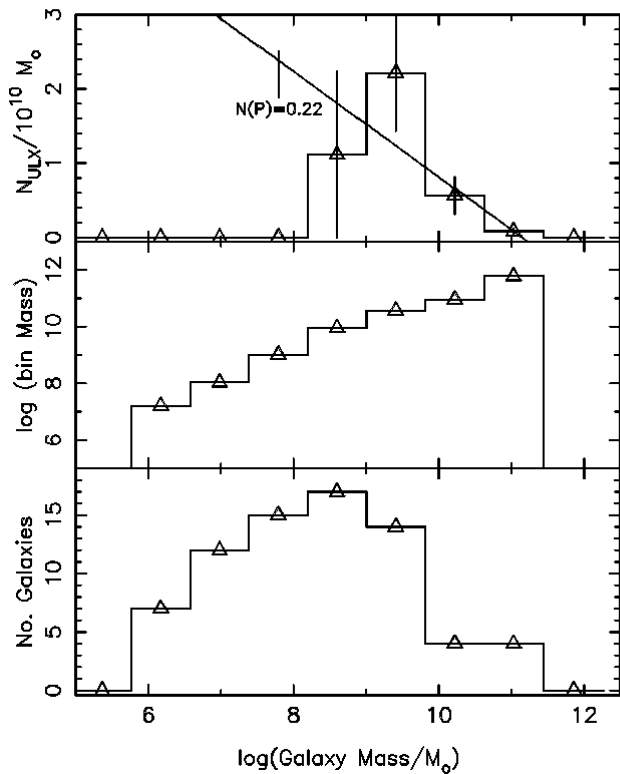
Optical/FIR Subsample



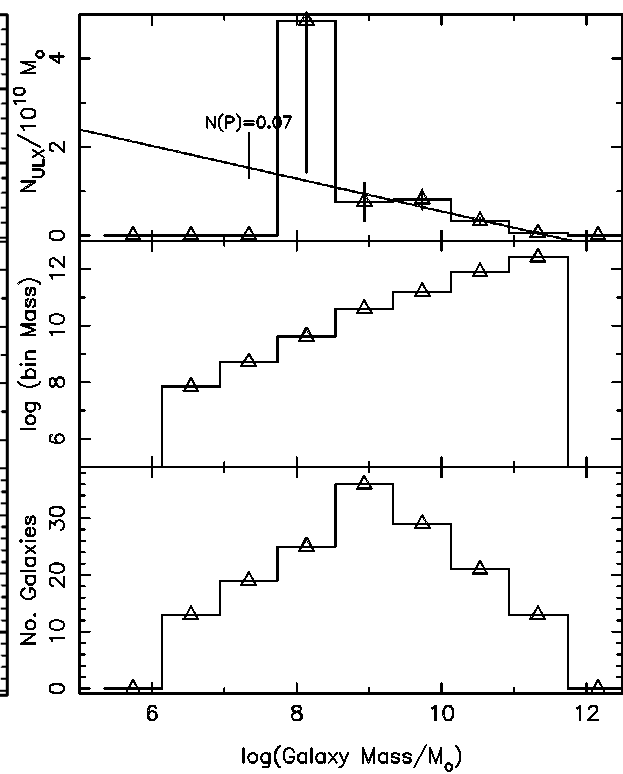
Optical/FIR Subsample



RASS Subsample



Catalog Subsample



Do Ultraluminous X-ray Sources Exist in Dwarf Galaxies?

YES....specific frequency increases
with decreasing host galaxy mass....

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N_{ULX} scales with SFR:

$$N_{\text{ULX}} = 1.1 * \text{SFR}$$

(Grimm+03, Swartz+04, Liu+06)

SFR proportional to L_{FIR}

Do Ultraluminous X-ray Sources Exist in Dwarf Galaxies?

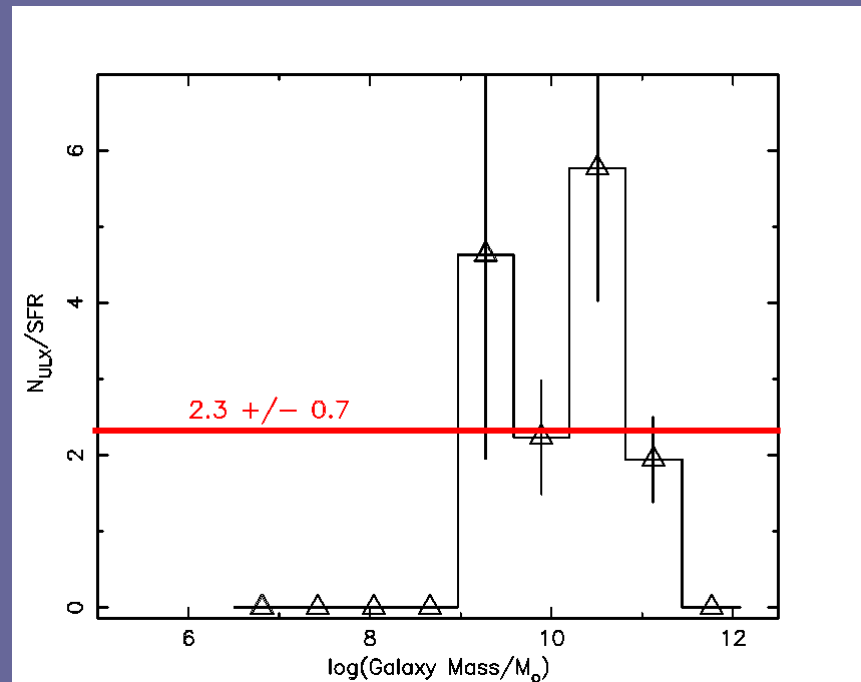
YES....specific frequency increases with decreasing host galaxy mass....

N_{ULX} scales with SFR:

$$N_{\text{ULX}} = 1.1 * \text{SFR}$$

(Grimm+03, Swartz+04, Liu+06)

SFR proportional to L_{FIR}



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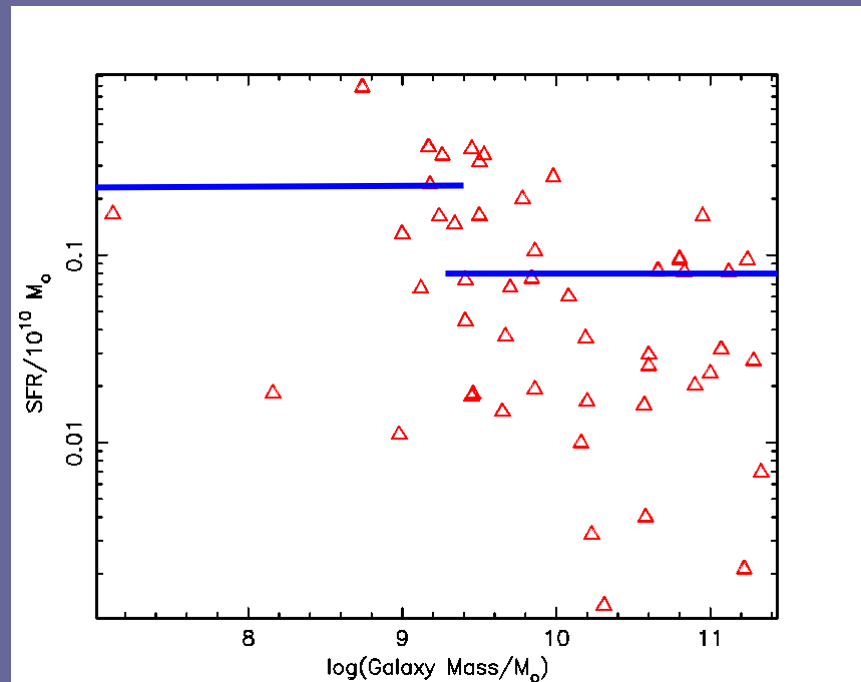
YES....specific frequency increases with decreasing host galaxy mass....

N_{ULX} scales with SFR:

$$N_{\text{ULX}} = 1.1 * \text{SFR}$$

(Grimm+03, Swartz+04, Liu+06)

SFR proportional to L_{FIR}



What may favor ULX formation in Dwarfs?

- **Low metallicity -> high mass SNe -> BH
higher L_{Edd} (Heger+03)**
- **Low shear -> higher mass MC -> sample high
end of IMF (Billet+02)**
- **Cold, massive clusters -> protostar mergers
-> top-heavy IMF (Peretto+07)**