Not just an X-ray Extractor

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Yaxx is a Perl script that facilitates batch data processing using Perl open source tools and commonly available astronomical software (CIAO/*Sherpa*, SAS, HEAsoft). For X-ray analysis it includes automated spectral extraction, fitting, and report generation. The primary emphasis is on having a simple tool that can be run without requiring an extensive learning curve. However, for those with the motivation, *yaxx* is highly configurable and can be customized to support complex pipeline data analysis. In particular the *yaxx* processing flow is fully configurable, easily allowing automation of data reduction steps. *Yaxx* includes default processing threads and output templates for *Chandra* and *XMM* spectral analysis.

Key Features:

- Uses Perl and CIAO/*Sherpa*, plus other packages (e.g. SAS and FTOOLS) as needed
- Full reference manual, quick-start guide, and installation guide
- Easily customized via configuration and template files
- Takes advantage of the powerful *Sherpa* and *S-lang* scripting capability
- Uses file dependencies in each processing step for selective reprocessing
- Processing summaries created in HTML and postscript formats
- Files and *Sherpa* fit script ready for interactive analysis
- Generates convenient FITS tables of all spectral fitting results
- Completely free, as in beer and open source software
- Supported on Linux and Solaris platforms

Yaxx is Simple

Fit one source with default models

Copy input data

```
mkdir obs3102
cp $YAXX/Test/Data/obs3102/acisf03102_evt2.fits obs3102/
cp $YAXX/Test/Data/obs3102/pcadf_asol1.fits obs3102/
```

Create object list file

Create a new file named `sample.dat` and insert the following lines:

```
obsid    src      redshift        X       Y      object
3102     1        0.32          4167    4085     Q1250+568
```

Copy standard configuration file

```
cp $YAXX/User/yaxx.cfg ./yaxx.cfg
```

Run *yaxx*

```
$YAXX/yaxx
```

View fit results

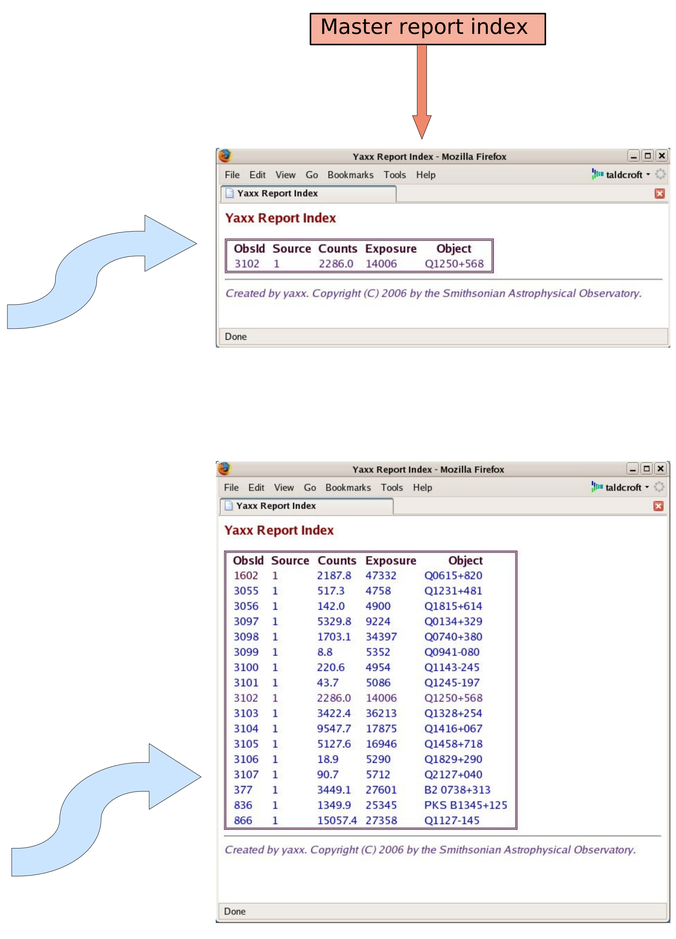
```
firefox report_index.html
```

Fit 17 sources with default models

Edit configuration file to set input data directory

```
input_dir = /data/chandra/obs%d # Input data dirs
```

Create object list file with all sources
Run *yaxx*
View fit results



Yaxx is not Simple

Processing threads

- Configuration and processing flow required for each type of data analysis is encapsulated in a processing thread
- Yaxx currently has standard threads for *Chandra* and *XMM*
- A derivative of the *Chandra* thread is used for ChaMP processing
- A major survey project could develop a standard thread to provide a unified configuration for outputs, models etc for individual studies within the project
- Detailed data processing is defined entirely within the thread configuration file, making it possible to tweak existing threads or define completely new threads (possibly having nothing to do with spectral fitting!)

```
# Generate GTI table to filter out times of high rates
<process_step>
  name          create_pn_gti_table
  dir           %FILE{obsid_dir}%
  detector      pn
  depend_file   %FILE{obsid_bgd_rate}%
  target_file   %FILE{obsid_gti}%
  command <<COMMAND
    tabgtigen
      table=%FILE{obsid_bgd_rate}%
      expression='RATE<=1.0'
      gtiset=%FILE{obsid_gti}%
      -w 0
COMMAND
</process_step>
```

Target dependencies

%FILE{}% macro lets *yaxx* worry about file names. In this case the name includes the XMM detector.

Command definition

Configuration

- Most behavior of *yaxx* is controlled by a hierarchical set of configuration files, from System level (global options not changed after installation) down to options specific to a single source.
- Spectral fitting models are completely configurable
- Choice of fit model(s) determined by *yaxx* for each source based on user-supplied criteria
- Output report format specified by HTML and LaTeX templates

Example: Calculate unabsorbed luminosity

- Using the powerful combination of the Sherpa, S-lang, and the yaxx macro language, one can do complex manipulations and capture the results in the output data
- The following is placed in the *Sherpa* fit template:

```
# Calculate the 2-8 keV (rest-frame) unabsorbed luminosity
e2_rest = 2.0 / (1+%VALUE{redshift}%)
e8_rest = 8.0 / (1+%VALUE{redshift}%)
unabs_flux2_8_rest = get_eflux(1, [e2_rest, e8_rest], "pow1")
flux_cmd = "flux_to_luminosity.pl -redshift %VALUE{redshift}% -flux "
           + string(unabs_flux2_8_rest.value)
fp = popen(flux_cmd, "r")
lines = fgetslines(fp)
()=pclose(fp)
unabs_lumin_2_8_rest = lines[0]
fits_update_key(fp, “LU20_80R”, unabs_lumin_2_8_rest, “Luminosity”)
```

Grab results

Run an external perl script

Save as MDL file header keyword

Yaxx real-world use and development

- *Yaxx* supplies core data processing functions for three major survey projects: ChaMP, ANCHORS, and C-COSMOS
- Several analysis projects which are **not** concerned with X-ray data now take advantage of the capabilities of *yaxx*
- Substantial development continues, primarily focused on refactoring code to split out core processing functions from specific analysis thread (e.g. X-ray spectral fitting)
- *Yaxx* is no longer Yet Another X-ray Xtractor. *Yaxx* is *yaxx*!

Download

- Further information and source download at **http://cxc.harvard.edu/contrib/yaxx**

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