

The Merger-Free Co-Evolution of Galaxies and Supermassive Black Holes

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UC San Diego

Note: talk originally given October 18th, 2016; date at bottom right of some slides reflects date talk was exported to PDF.

@the_zooniverse

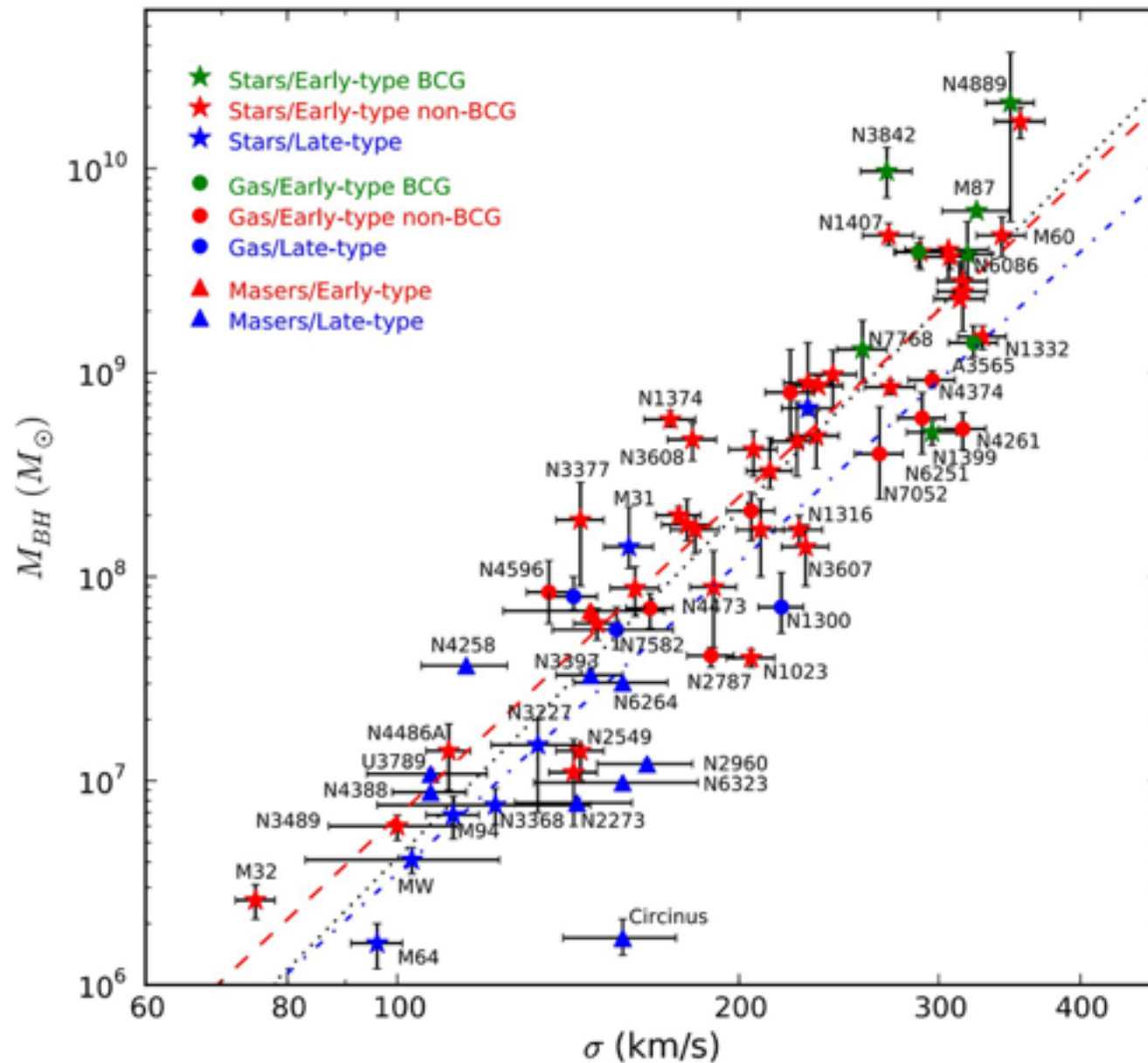
@galaxyzoo

@vrooje



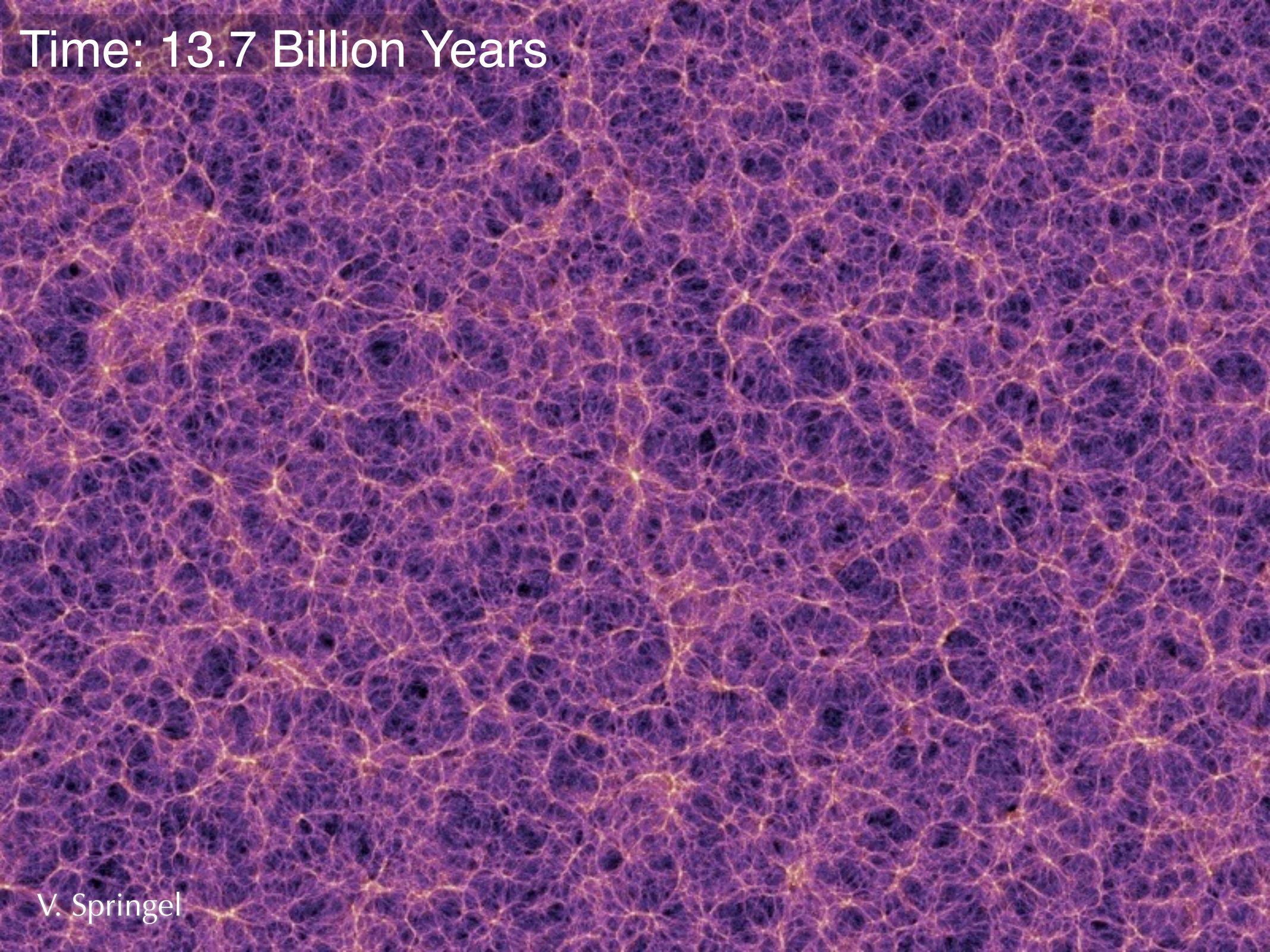
Galaxies come in a diverse array of shapes, sizes, evolutionary histories.
And yet...

Black Holes & Galaxies Co-Evolve



Time: 0.21 Billion Years

Time: 13.7 Billion Years



V. Springel

Time: 13.7 Billion Years



Do Mergers Drive Co-Evolution?

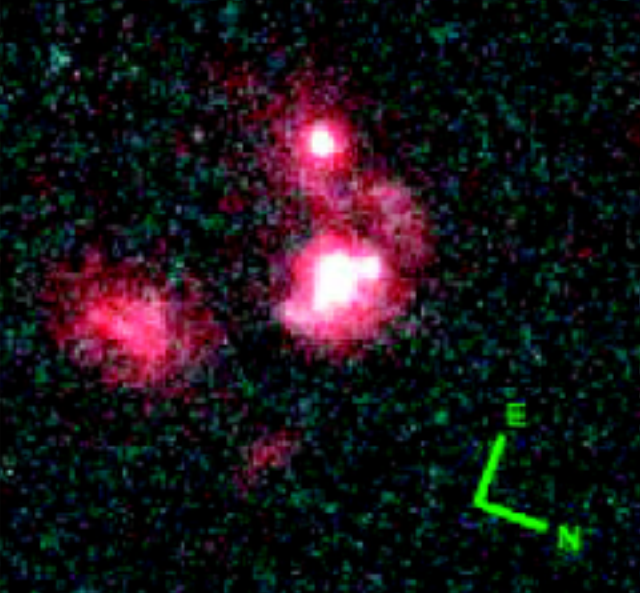


See: Sanders+88, Hopkins+05,06 etc.

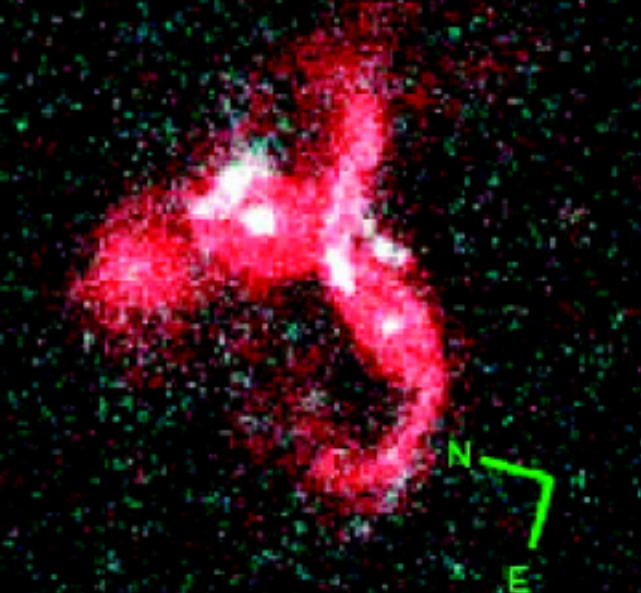
Credit: HST/NASA/STScI/ESA

Quasars: Trainwreck Mergers?

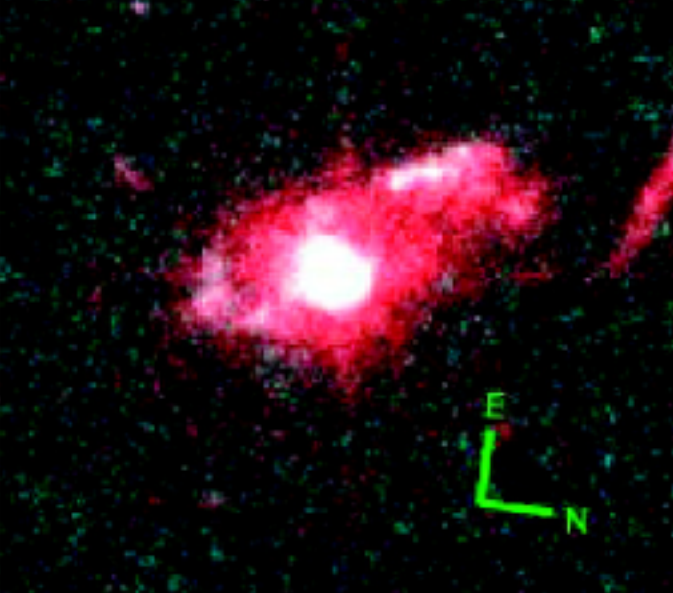
F2M0825+4716



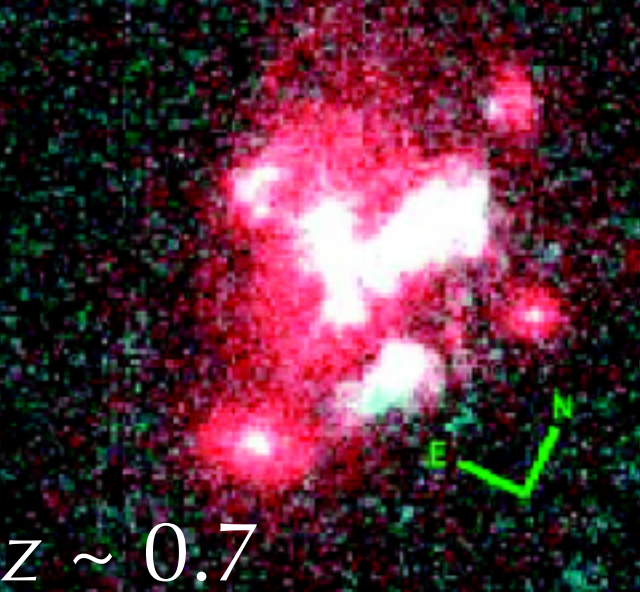
F2M0841+3604



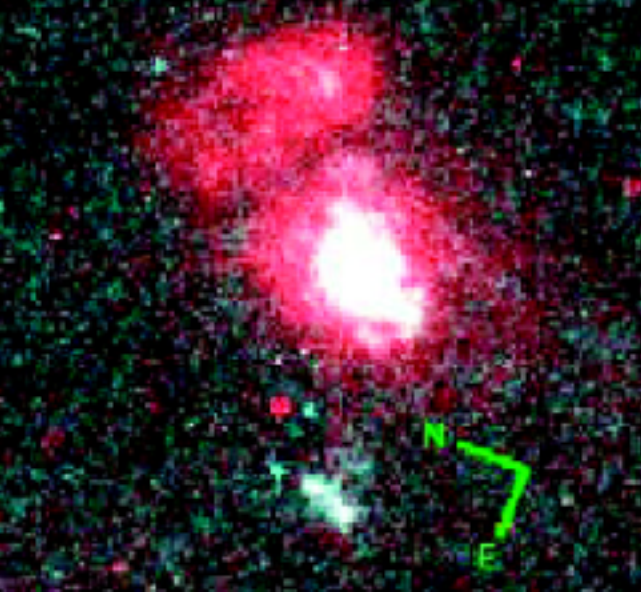
F2M0729+3336



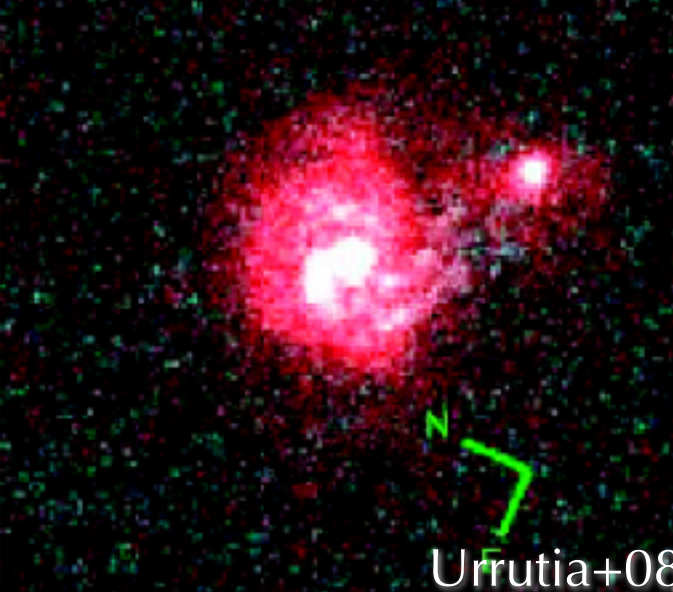
F2M1532+2415



F2M1507+3129



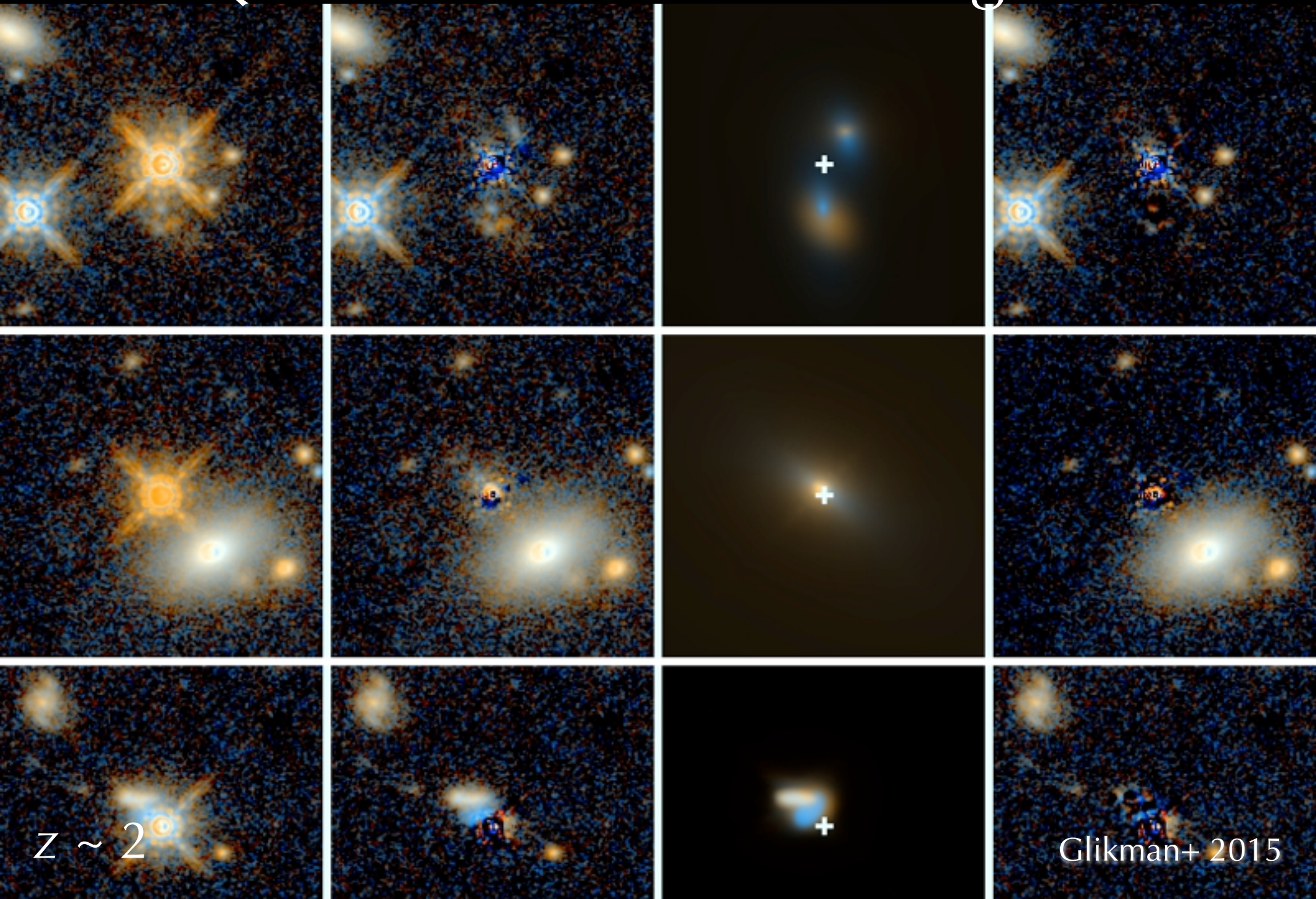
F2M1118 0033



$z \sim 0.7$

Urrutia+08

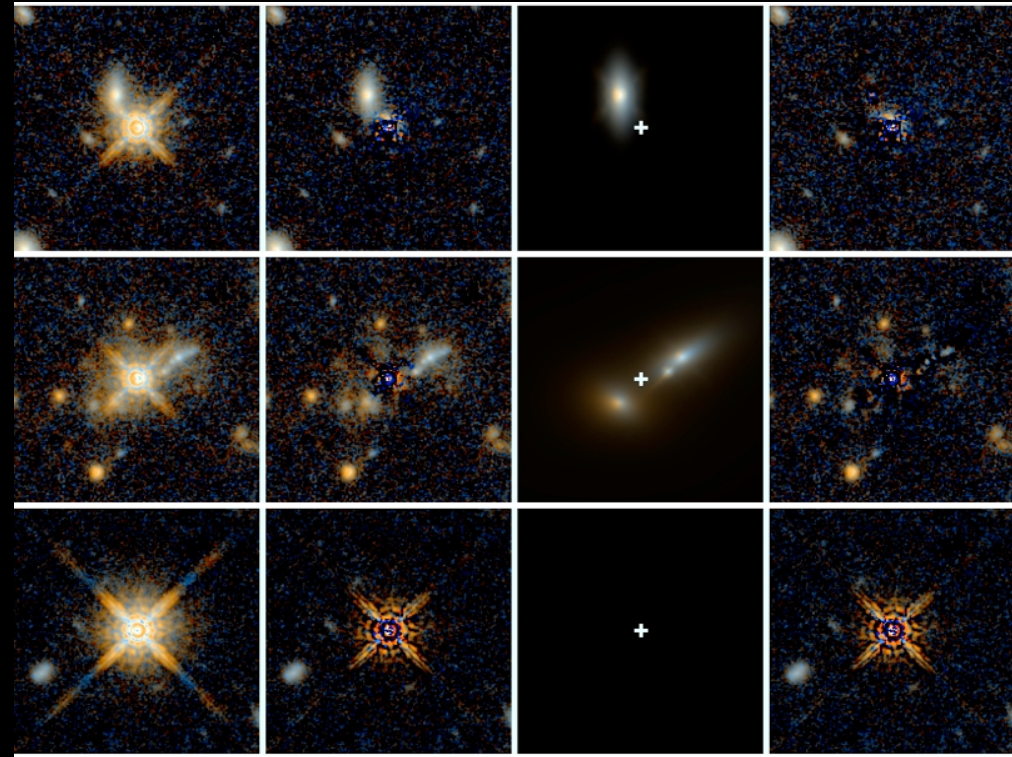
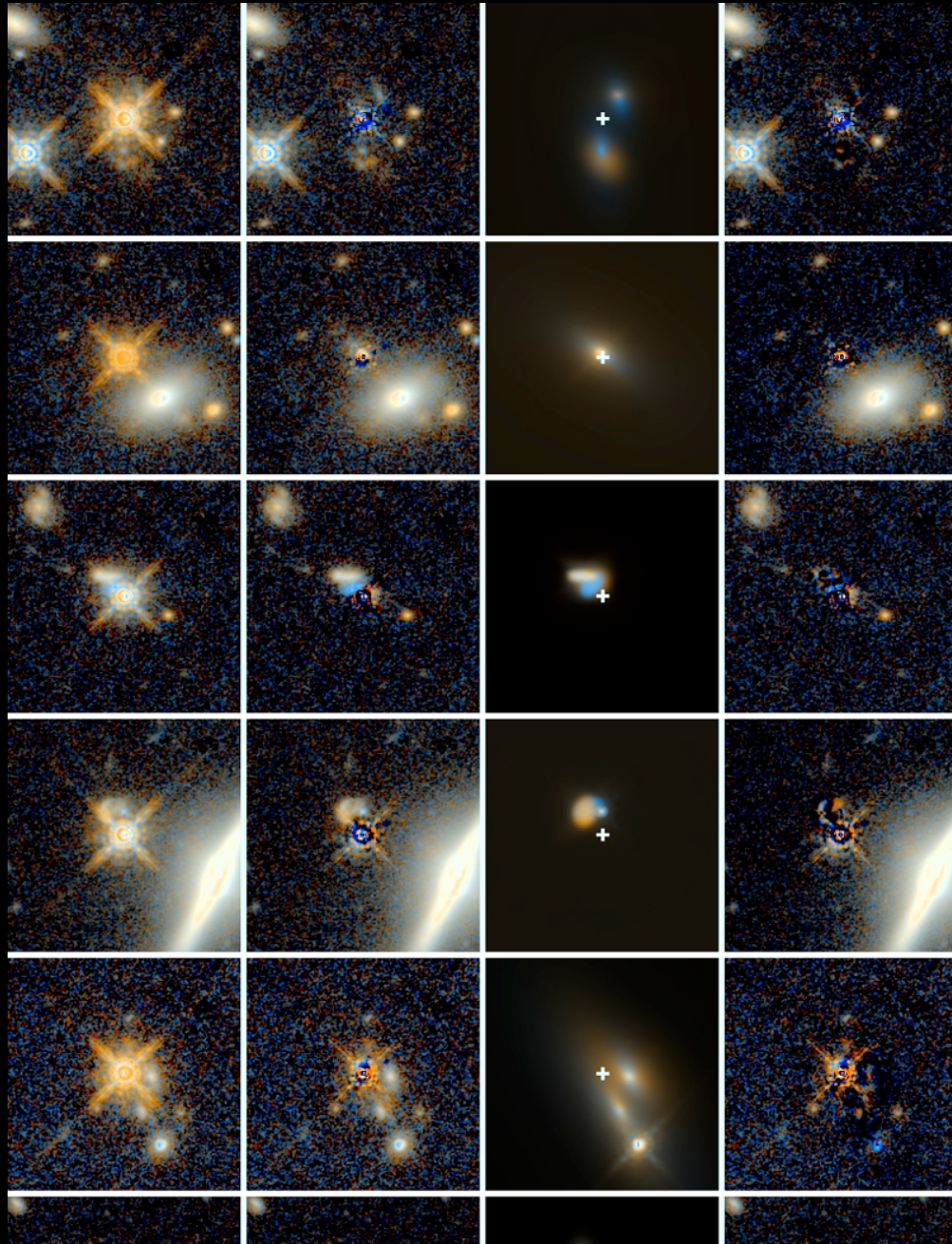
Quasars: Trainwreck Mergers?



$z \sim 2$

Glikman+ 2015

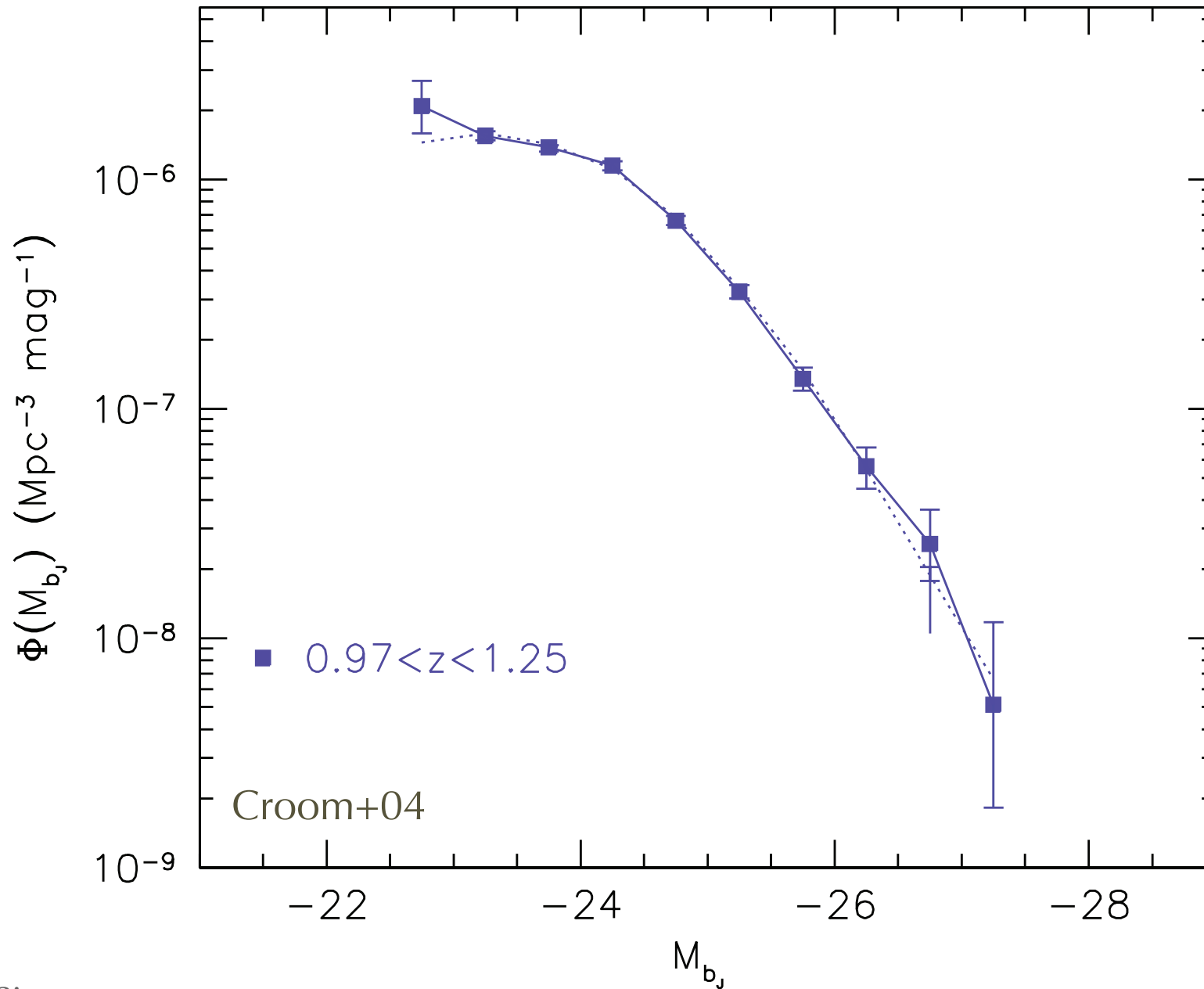
Quasars: Trainwreck Mergers?



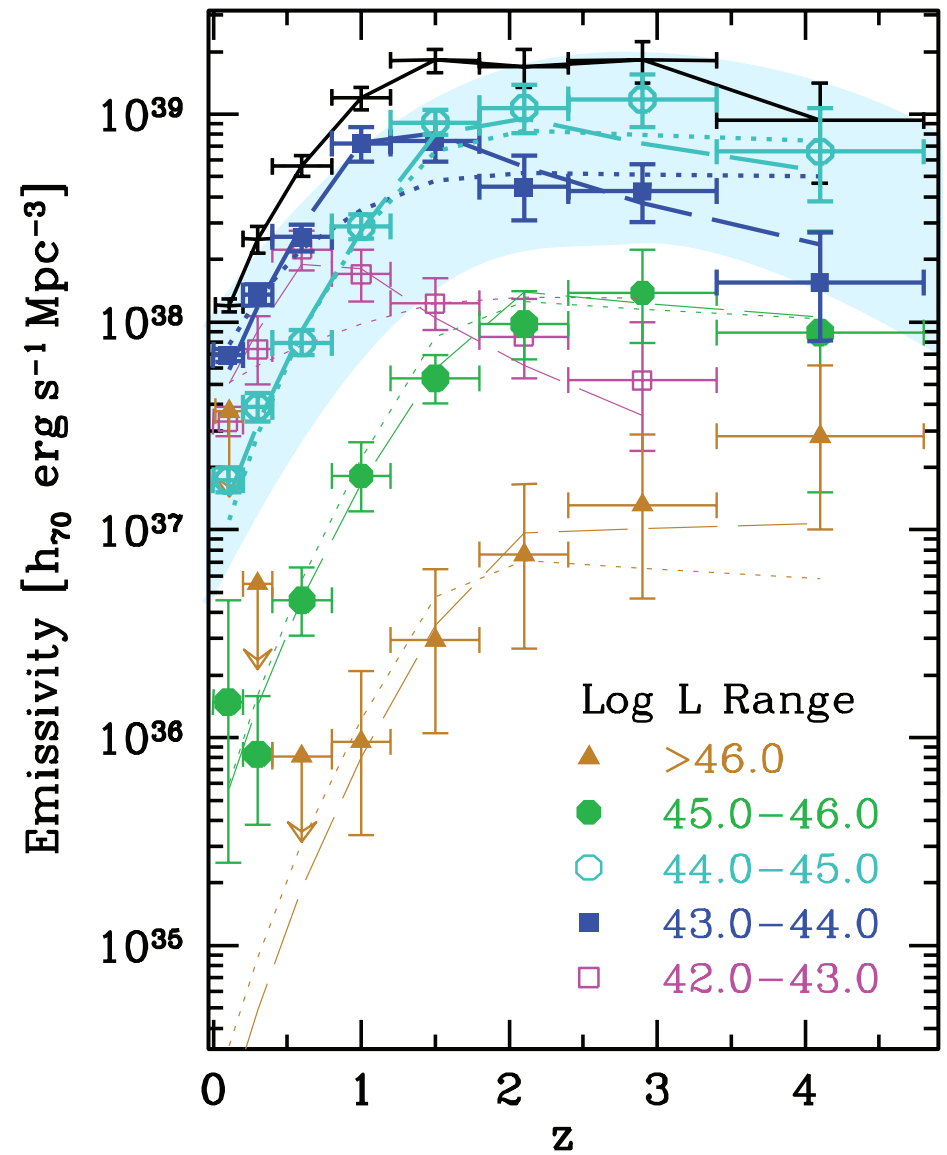
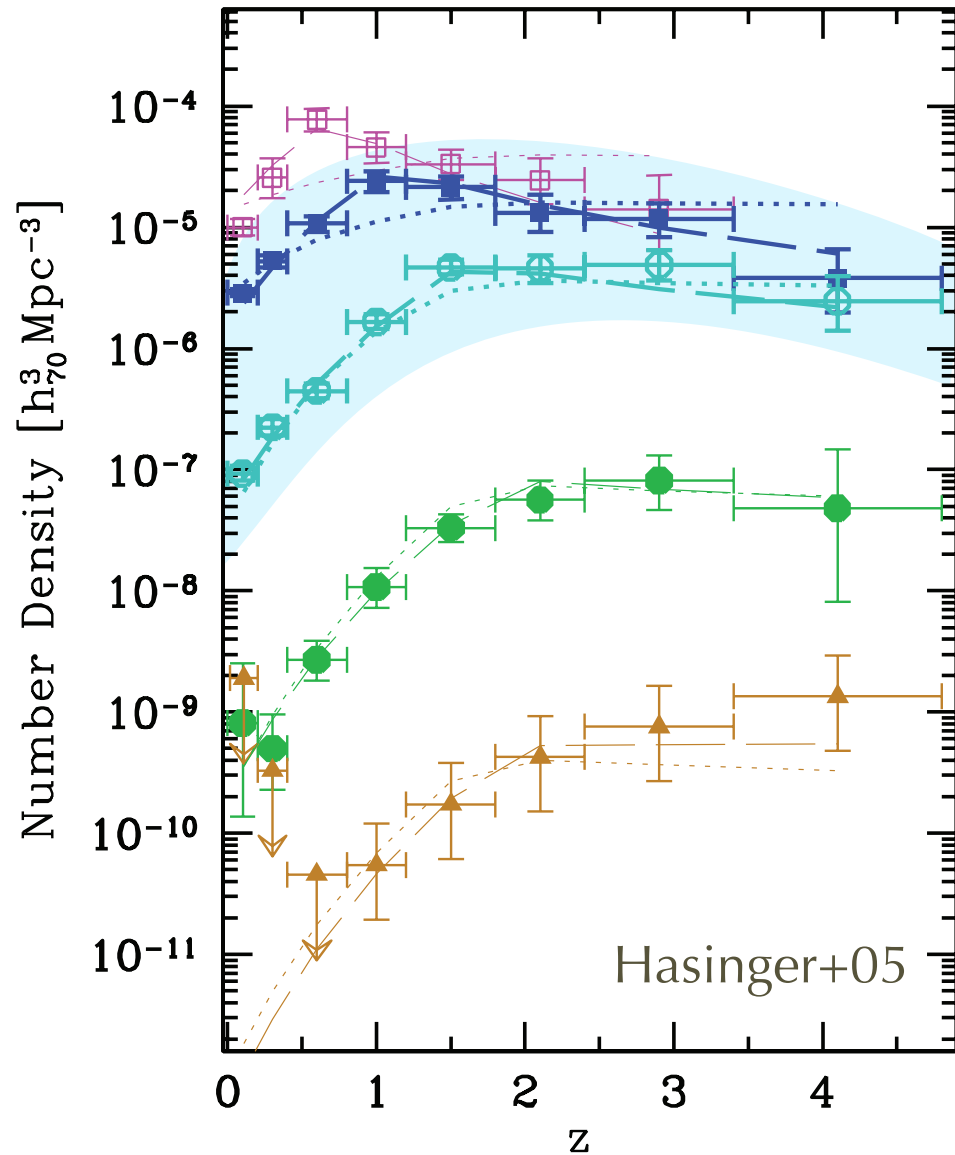
Glikman+15

Nov 2, 2016

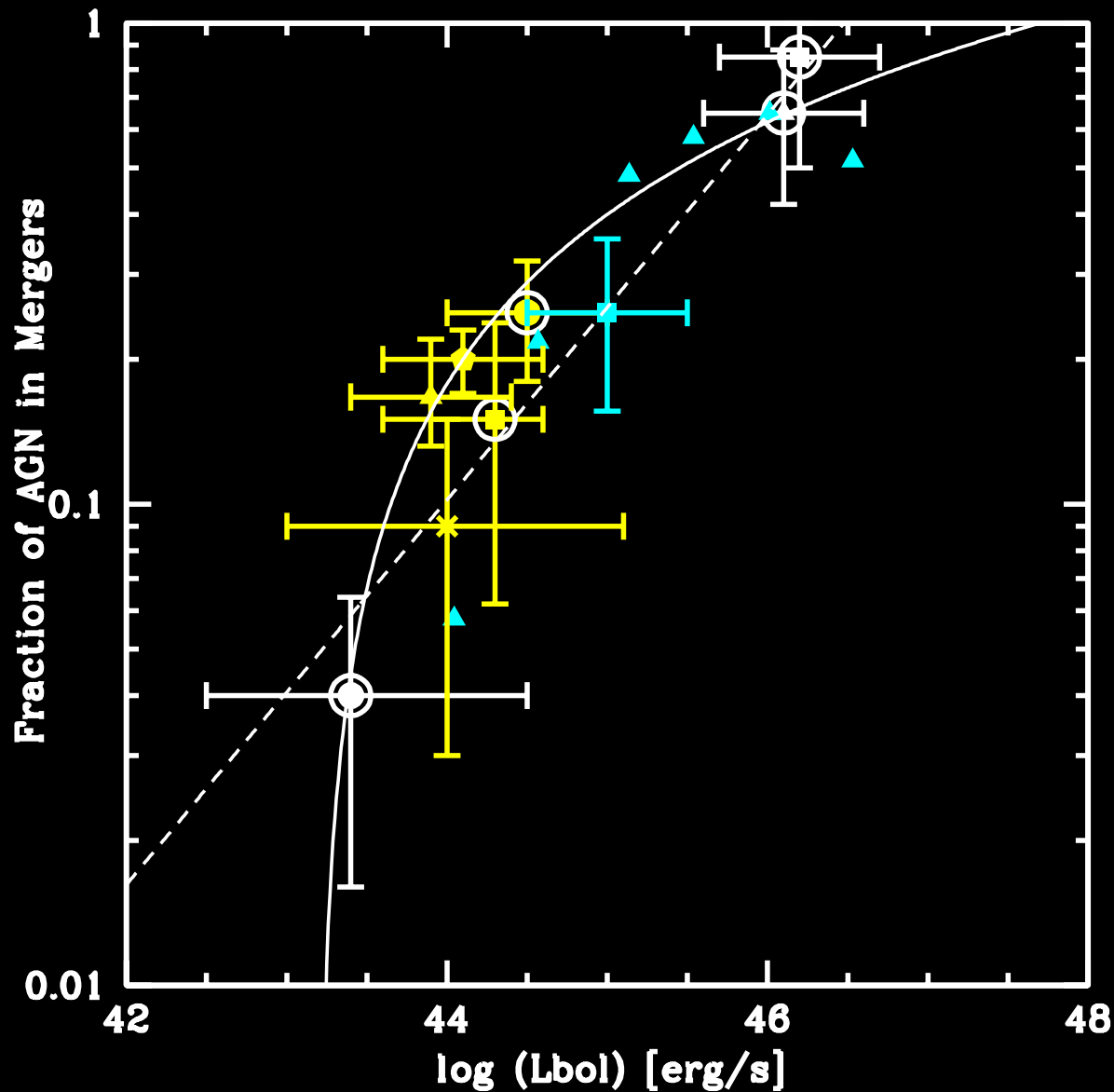
AGN: A Diverse Population



Don't Dismiss Moderate-Luminosity AGN

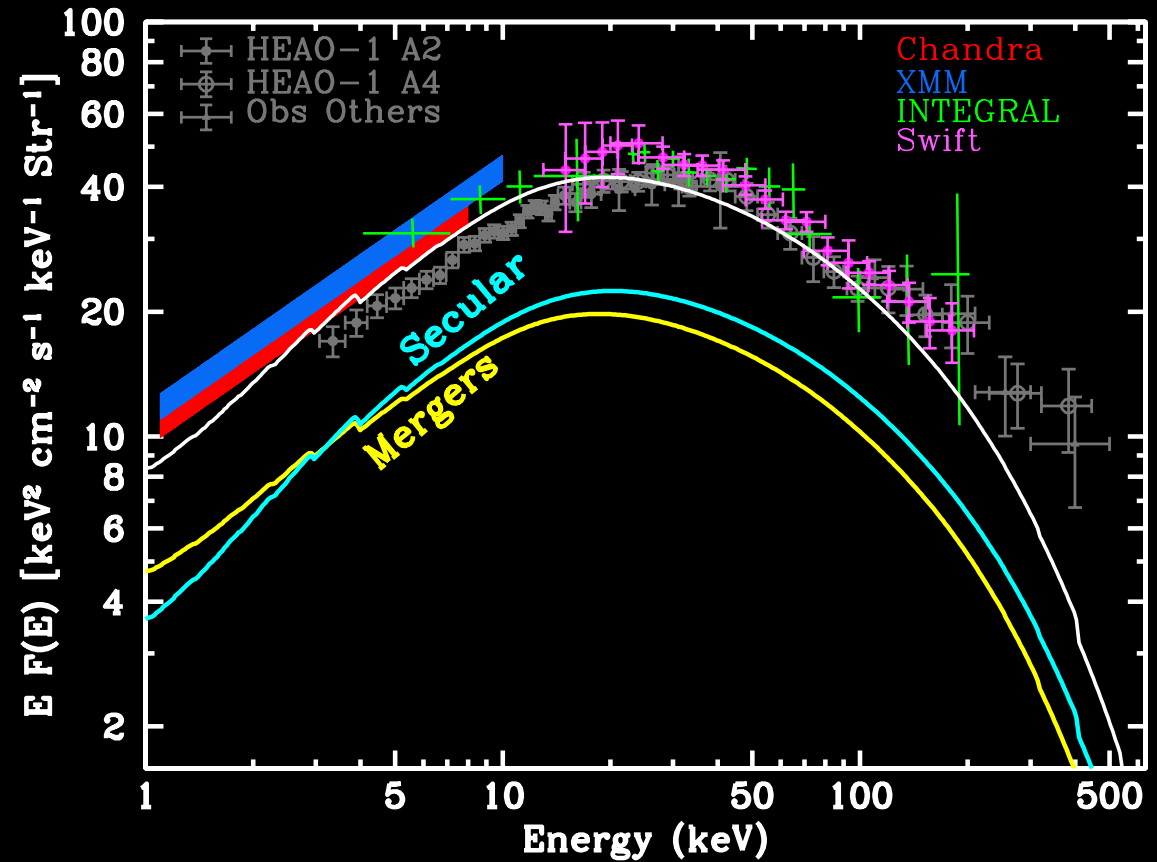
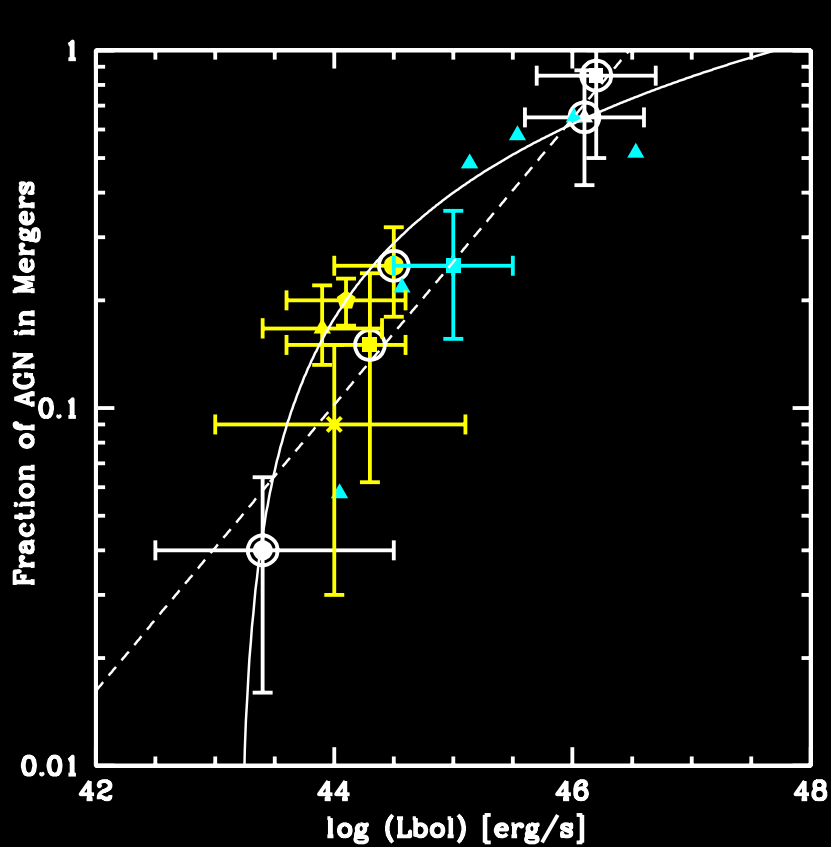


Major mergers only trigger luminous AGN



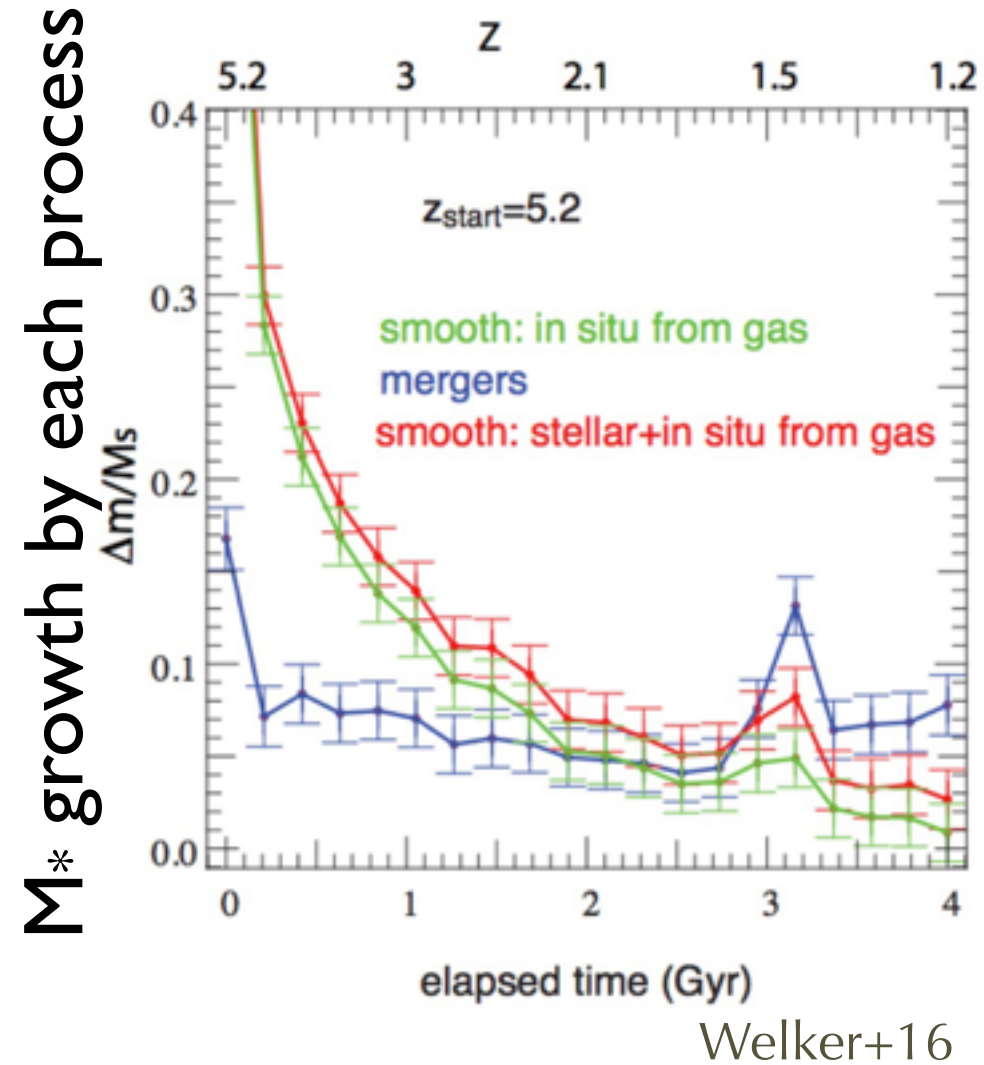
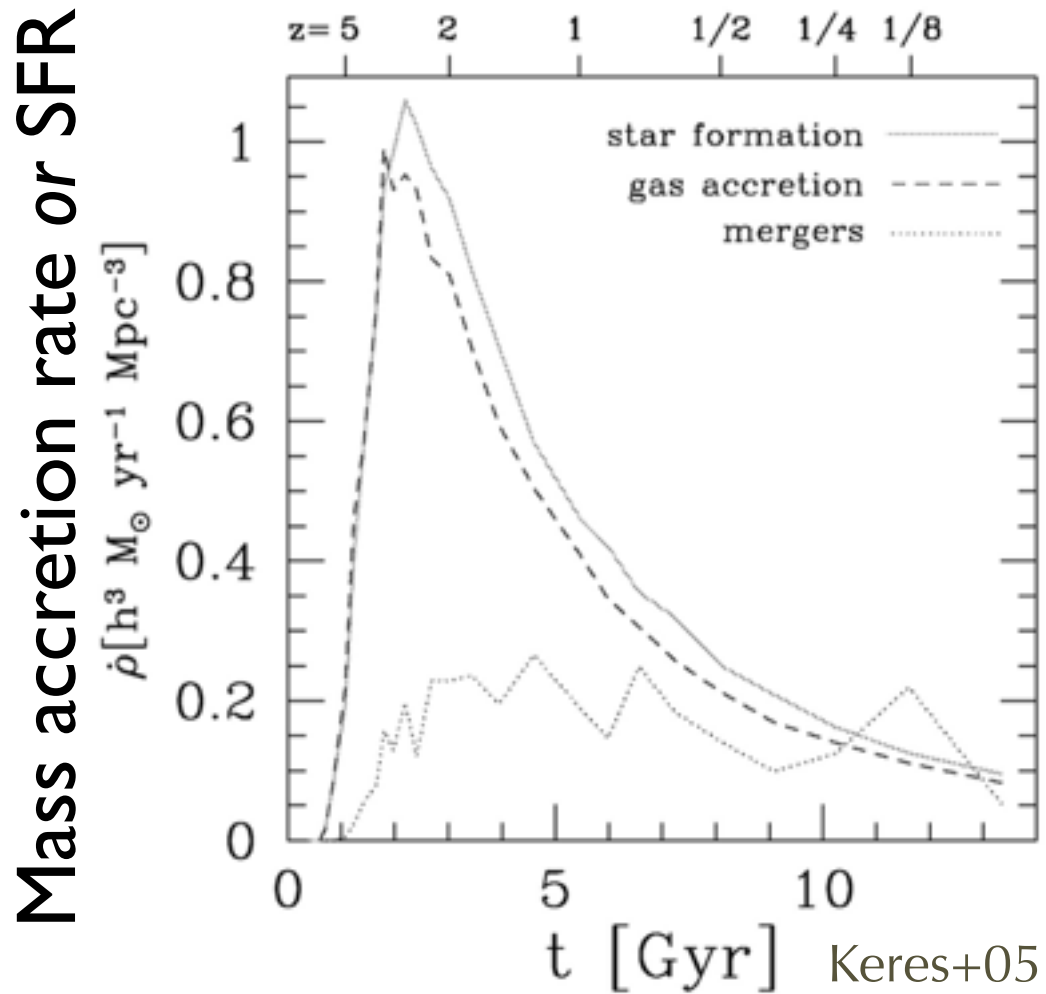
Treister+12

Major Mergers Only Trigger Luminous AGN

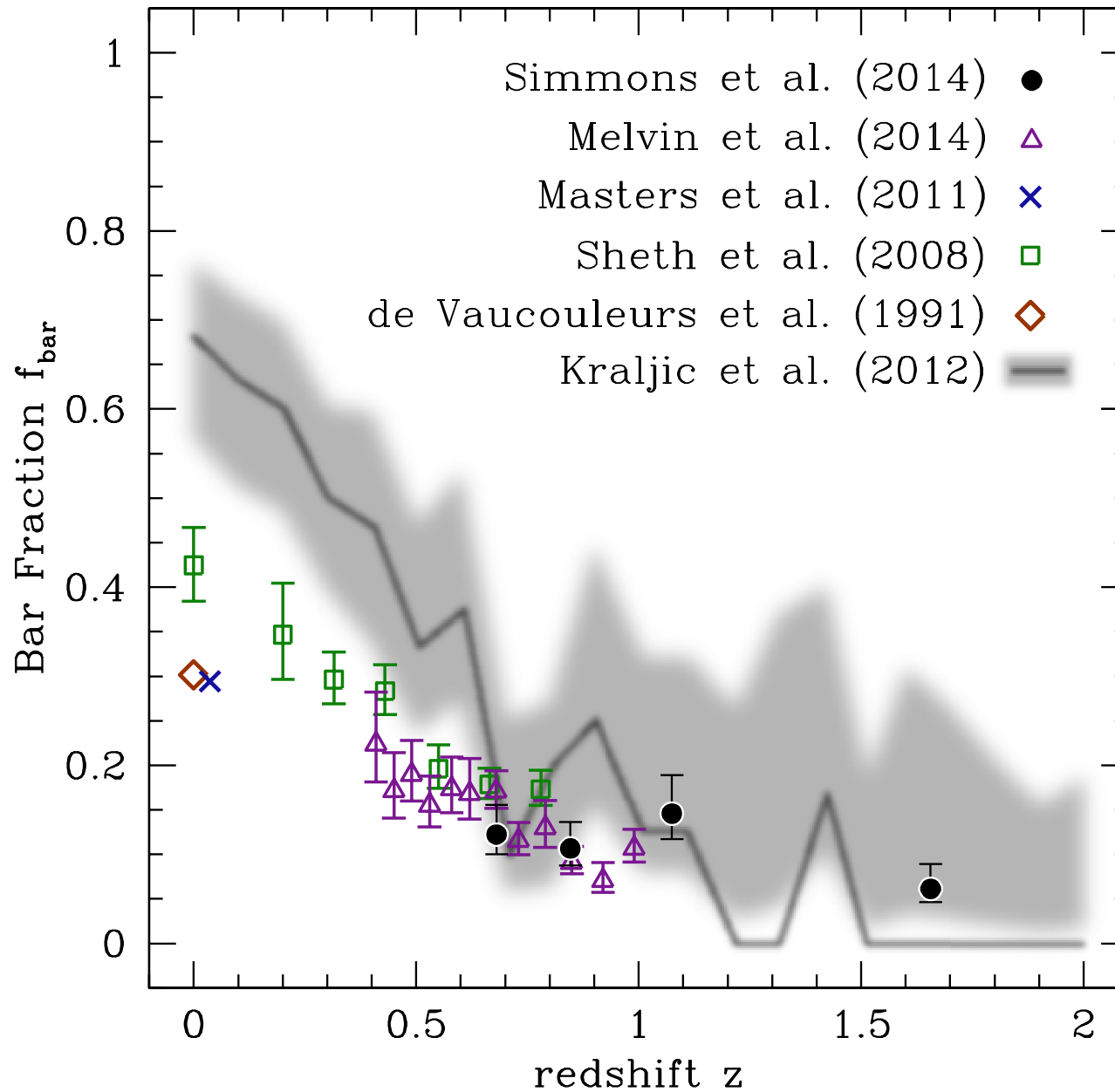


Treister+12

Merger-Free galaxy growth: cold accretion



Populations of settling disks since $z \sim 2$



Disentangling growth histories

Morphologies

Bulgeless

“Pseudo”-bulge

“Classical” bulge

Evolutionary Processes

Calm Merger-Free:
No dynamical heating

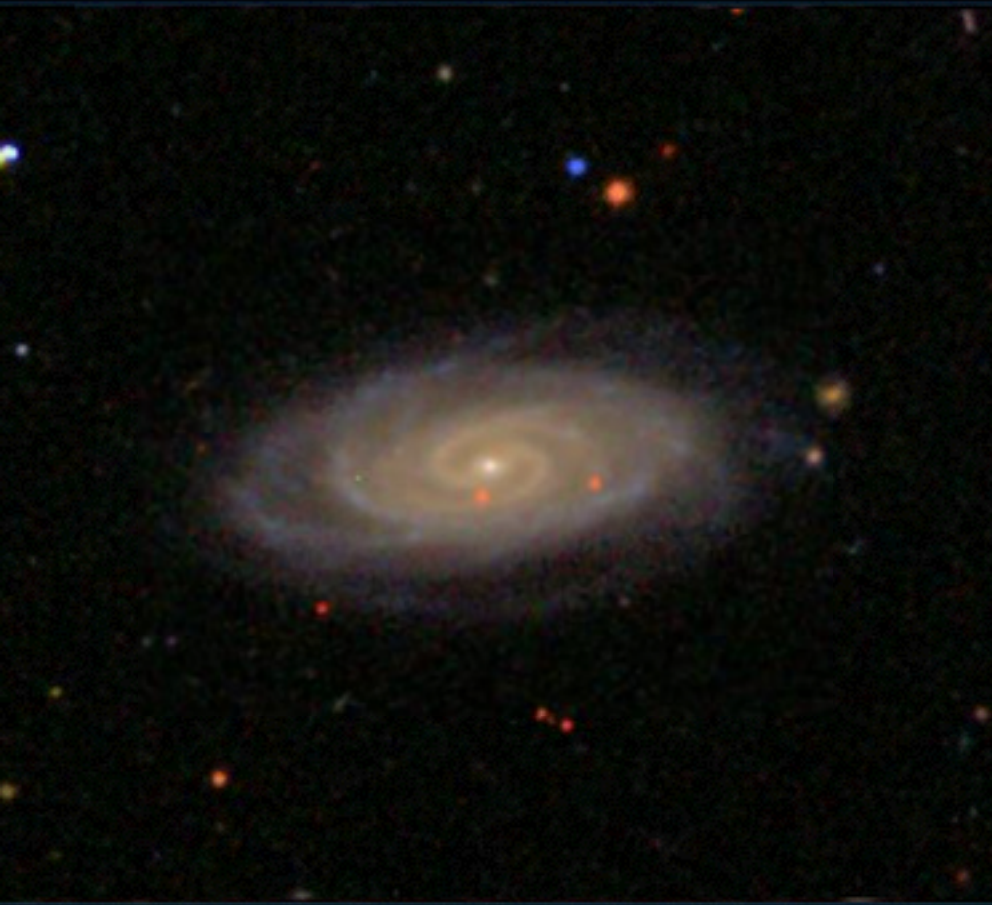


Violent Merger-Free:
Some dynamical heating



Merger-driven:
Significant dynamical heating





Classify

SDSS Favourite Invert

[Help](#) [Restart](#)

BULGE

How prominent is the central bulge, compared with the rest of the galaxy?



No bulge



Just noticeable

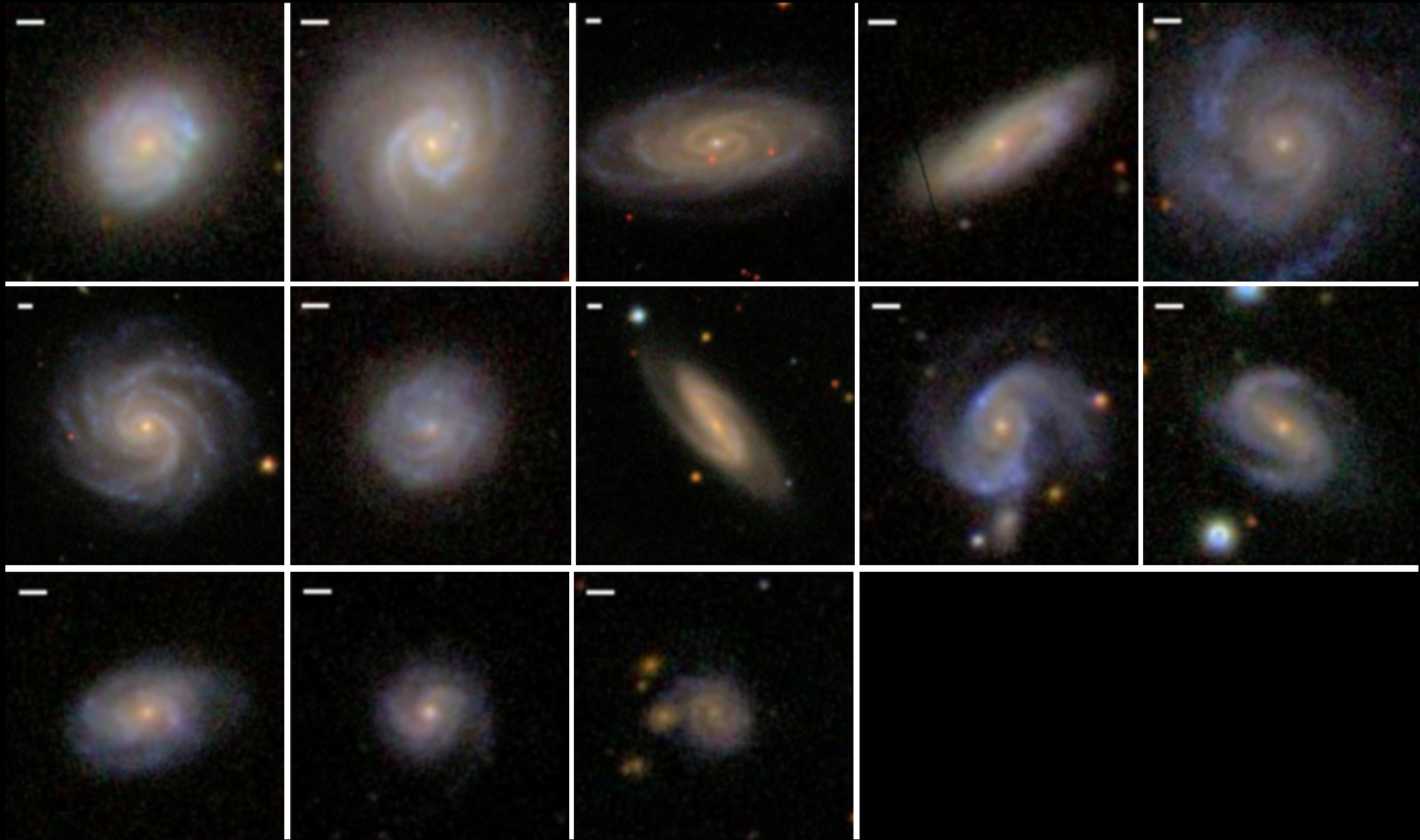


Obvious

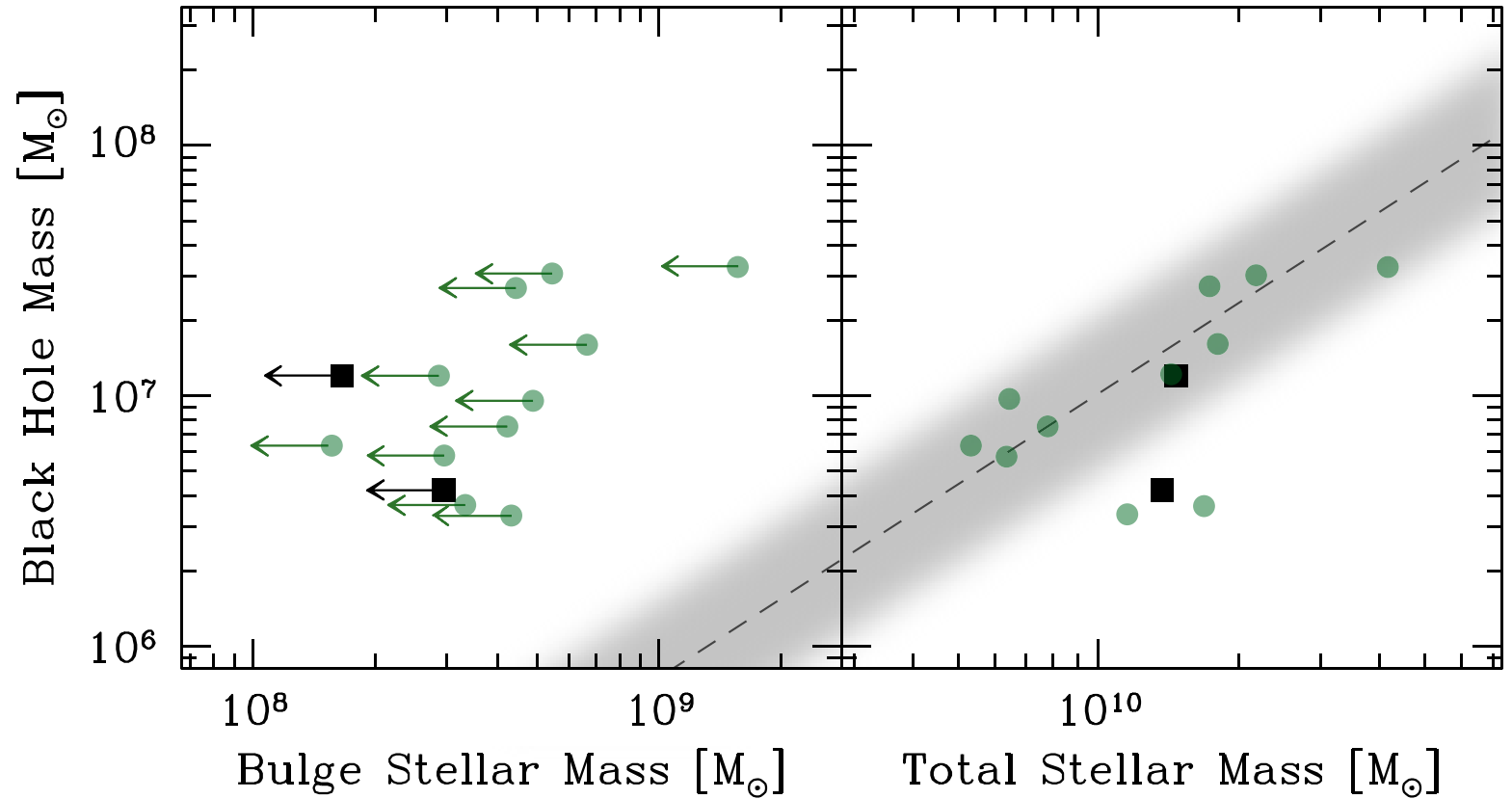
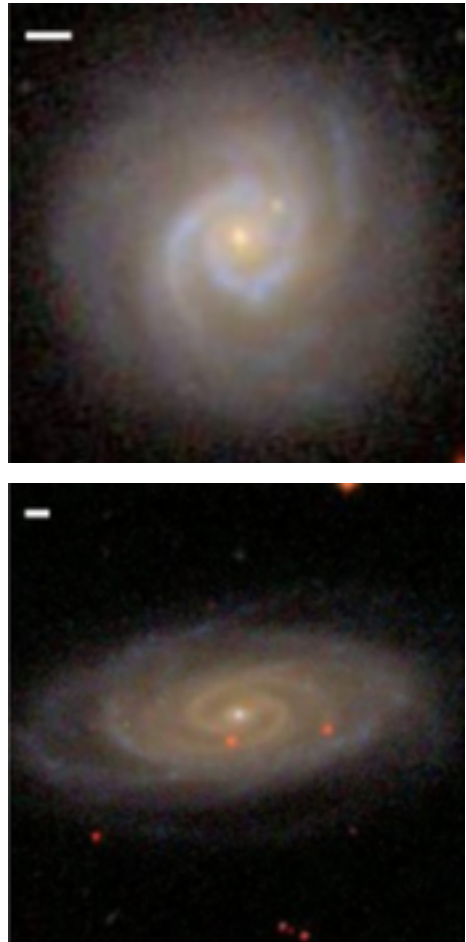


Dominant

Bulgeless Galaxies with Growing Black Holes

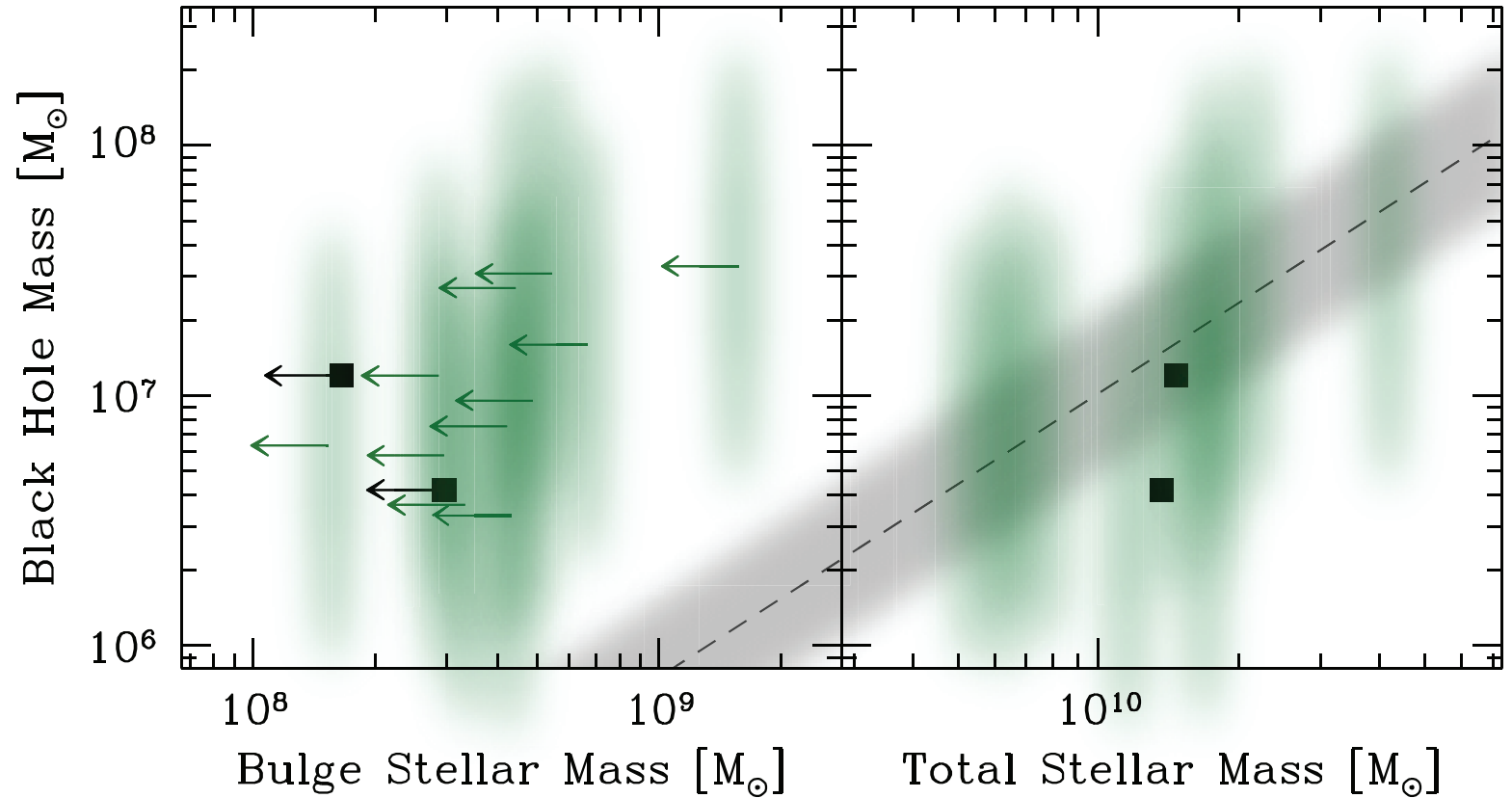
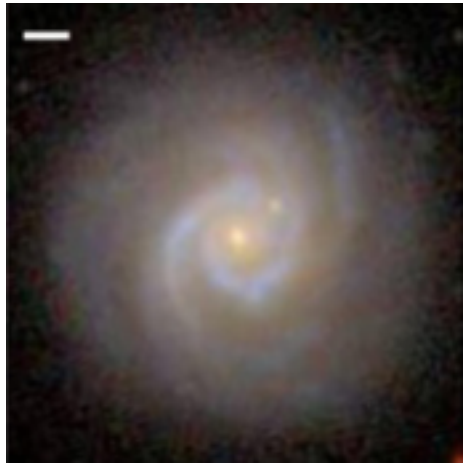


Merger-Free AGN & Host Galaxies



Simmons et al. (2013)

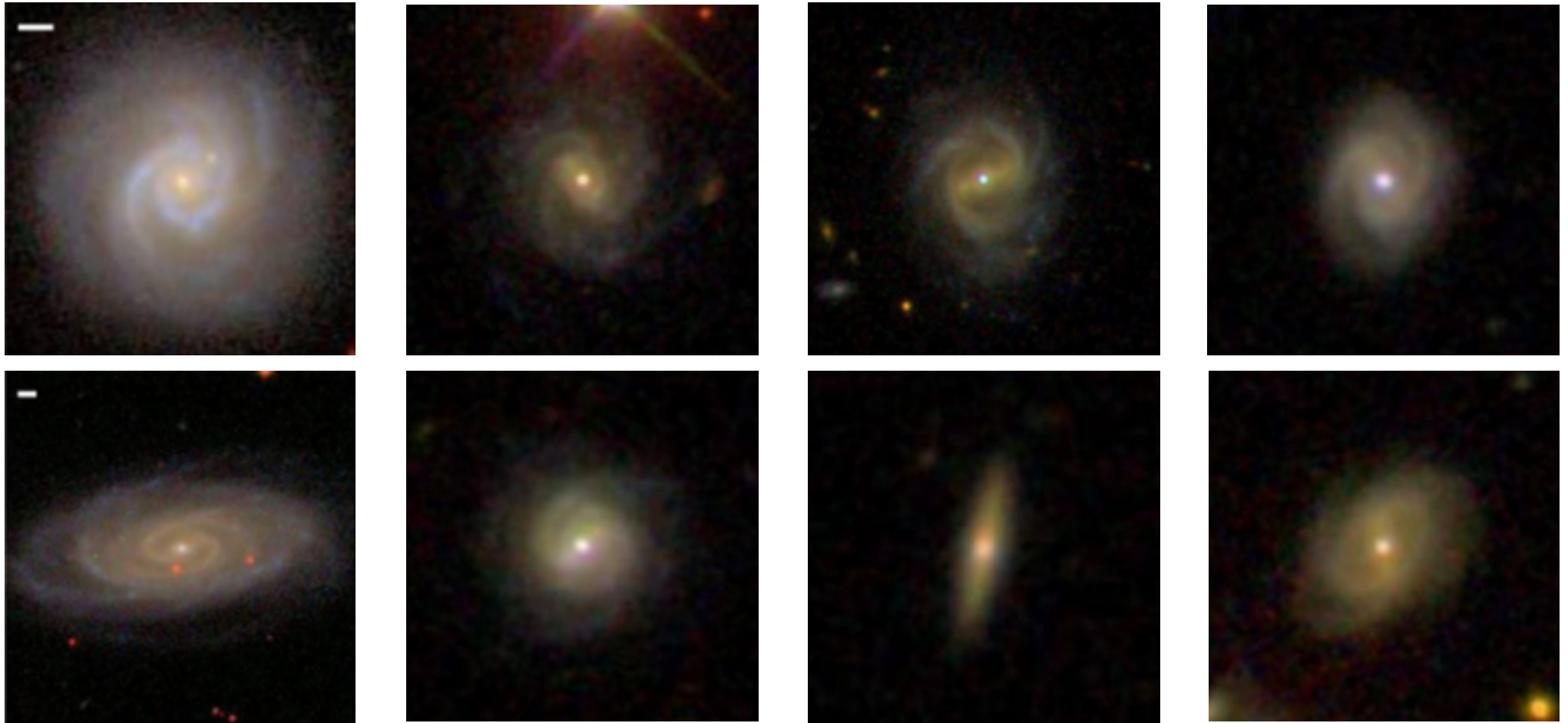
Merger-Free AGN & Host Galaxies



Simmons et al. (2013)

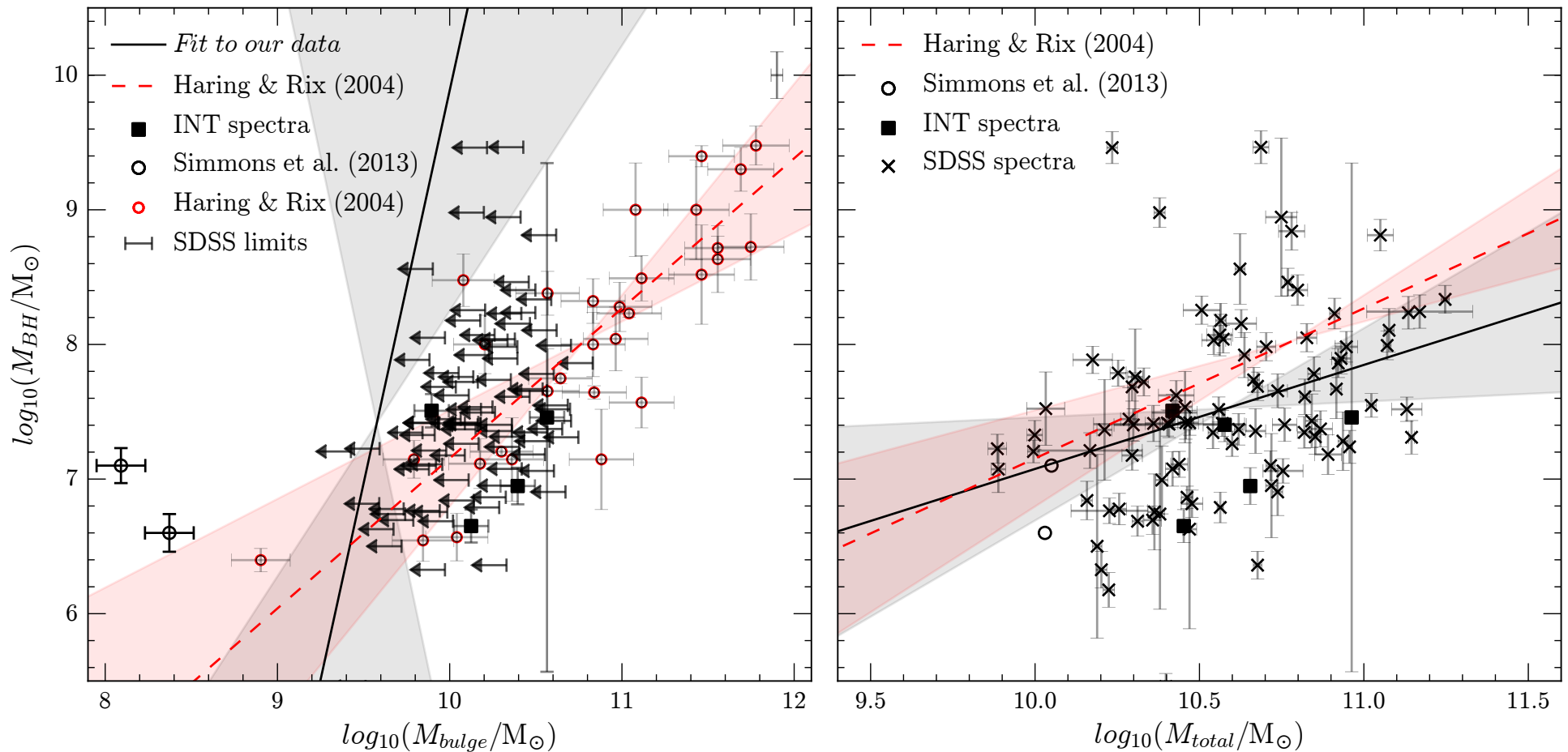
More realistic uncertainties versus the previous slide.

Merger-Free AGN & Host Galaxies



Merger-Free AGN & Host Galaxies

Secular processes can drive BH-galaxy co-evolution



Simmons+ (in prep)

Morphology tells a tale of formation history

Morphologies

Bulgeless

“Pseudo”-bulge

“Classical” bulge

Evolutionary Processes

Calm Merger-Free:
No dynamical heating



Violent Merger-Free:
Some dynamical heating

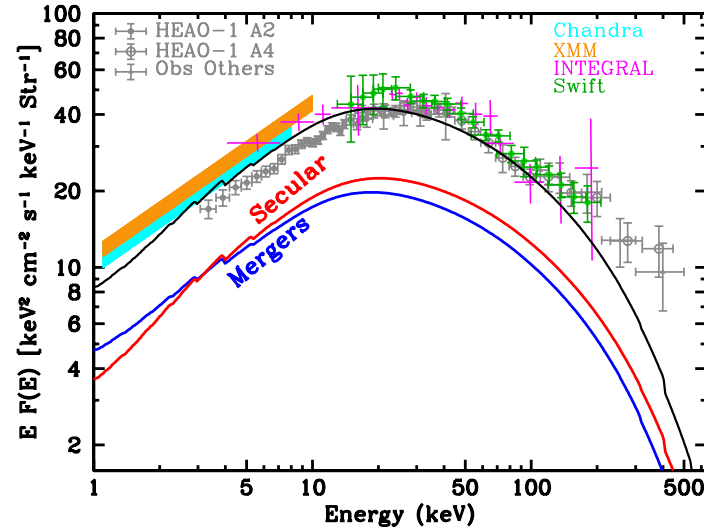
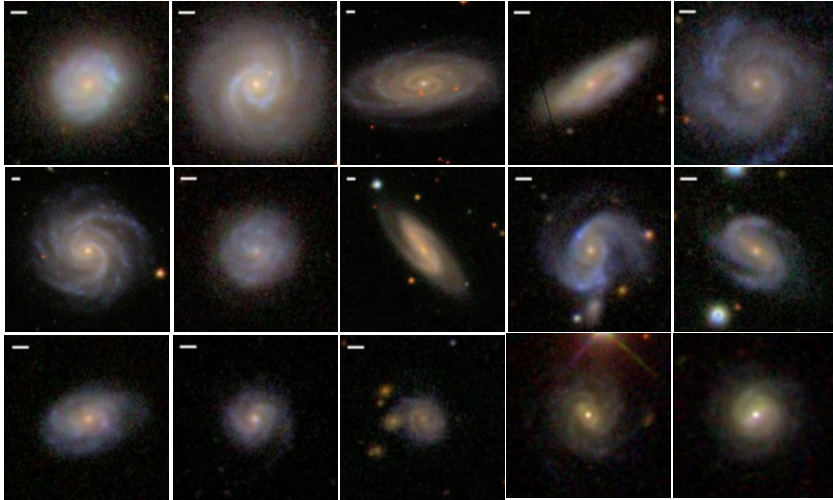


Merger-driven:
Significant dynamical heating

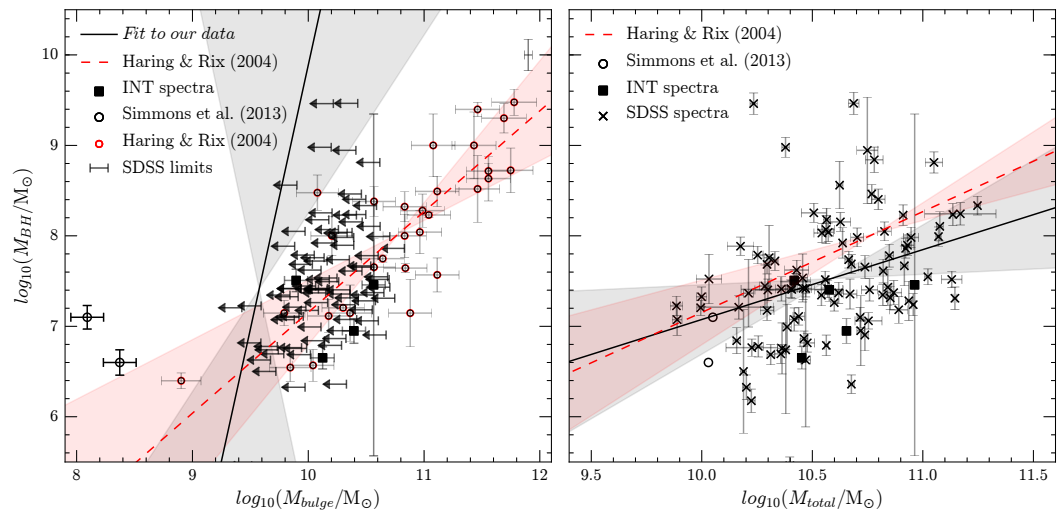


Summary

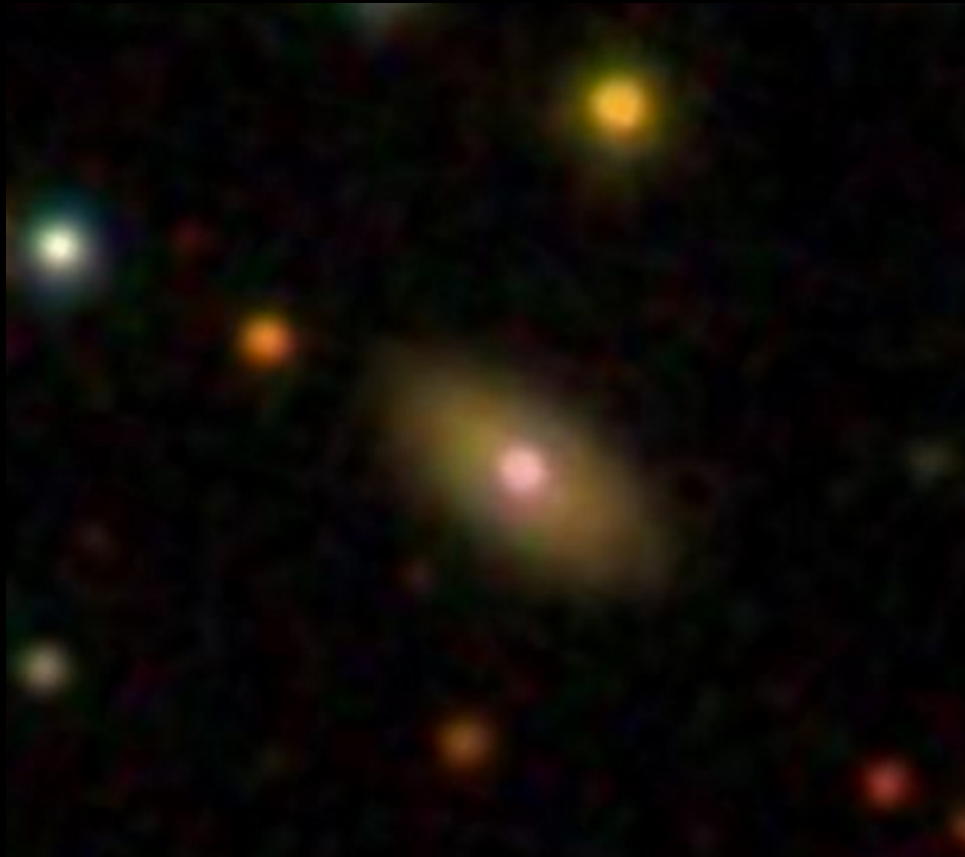
Substantial BH growth can occur in the absence of mergers.



Merger-free BHs appear to co-evolve with galaxies - co-evolution may be more fundamental than mergers.



Fresh Off The Telescope



SDSS *gri*



HST F814W

We found more disk-dominated galaxies and HST gave us lots of time: this is the first observation of the survey.