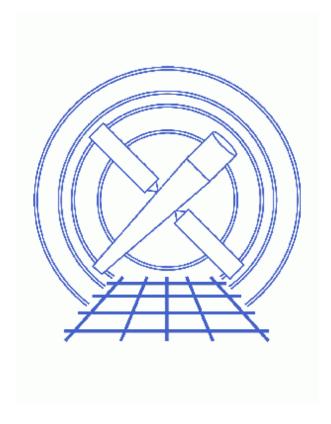
Calculating Statistics of Images



CIAO 3.4 Science Threads

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Last modified: 1 Dec 2006

Calculating Statistics of Images

CIAO 3.4 Science Threads

Overview

Last Update: 1 Dec 2006 – reviewed for CIAO 3.4: ChIPS version

Synopsis:

Although in many cases it is easier (and faster) to use the <u>dmstat</u> tool, the <u>sstats.sl</u> script described in this thread allows you to ignore regions of the image by using a mask file.

Purpose:

Use the <u>S-Lang</u> interpreter in <u>ChIPS</u> and <u>Sherpa</u> to calculate statistics of images.

Read this thread if:

you are interested in calculating the statistics of an ACIS or HRC image.

Related Links:

- Estimating Source Counts thread: related functionality of the <u>dmextract</u> tool.
- <u>dmopt</u> and <u>dmstat</u> ahelp pages: an alternative method of marking image regions as "uninteresting" setting such pixels to the Null value for that image type is discussed. Note that the sstats.sl script will *not* handle such cases.

Proceed to the <u>HTML</u> or hardcopy (PDF: <u>A4 | letter</u>) version of the thread.

Getting Started

Sample ObsID used: 1838 (ACIS-S, G21.5-09)

It is assumed that you have completed the <u>Single Chip Exposure Map</u> thread.

This thread uses the sstats.sl script. The most recent version of sstats.sl is v0.4 (4 Oct 2001):

unix% grep Id \$ASCDS_CONTRIB/share/slsh/local-packages/sstats.sl % \$Id: sstats.sl,v 0.4 2001/10/04 18:45:12 dburke Exp \$

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Note that \$ASCDS_CONTRIB/share/slsh/local-packages/ is the default path in the standard CIAO scripts installation; see the <u>Scripts page</u> for more information. *Please check that you are using the most recent version before continuing*. If you do not have the script installed or need to update to a newer version, please refer to the <u>Scripts page</u>.

Create an image of the region of interest

For this thread, we want to look at the average signal level in an annulus surrounding the source center (this is for the 1838_img_s3.fits file created in the <u>Single Chip Exposure Map</u> thread). We saved the selected region to stat.reg in CIAO format; its contents are:

```
unix% cat stat.reg
# Region file format: CIAO version 1.0
annulus(4070.5,4248.5,130,218)
```

Further details are available for <u>interactive region definition</u> using ds9.

We create an image of the region using this spatial filter. For the mask region – which should contain pixel values greater than 0 only for those pixels within the filter – we use a filtered copy of the exposure map (we can not use img_stat.fits as the mask image, since some of the pixels within the annulus contain no counts).

```
unix% <a href="mailto:dmcopy">dmcopy</a> "1838_img_s3.fits[sky=region(stat.reg)]" img_stat.fits unix% dmcopy "expmap_1.7kev_7.fits[sky=region(stat.reg)]" emap_stat.fits
```

We show a mosaic of the original image (log scaling) together with the filtered image (sqrt scaling) and the filtered exposure map (linear scaling).

Statistics of the selected region

We show results from using the <u>dmstat</u> tool and the sstats.sl S-Lang script:

A. dmstat

```
unix% punlearn dmstat
unix% dmstat img_stat.fits centroid=no
EVENTS_IMAGE
                         @: ( 4057.6402575 4030.9302575 )
   min: 0
                                  ( 4089.6402575 4082.9302575 )
   max:
            0.027786151623
  mean:
            0.16761614994
 sigma:
             2673
   sum:
  good:
             96199
  null:
              94770
```

The mean and standard deviation values represent the signal level within the annulus. This could equivalently be found from the original image:

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```
max: 2 @: (4089.6402575 4082.9302575)
mean: 0.027786151623
sigma: 0.16761614994
sum: 2673
good: 96199
null: 94770
```

B. sstats.sl

```
unix% chips --slscript sstats.sl

Welcome to ChIPS, version CIAO 3.4
Copyright (C) 1999-2003, Smithsonian Astrophysical Observatory
chips> sstats( "img_stat.fits" )
Statistics of img_stat.fits.
Mean = 1.3997e-02 RMS = 1.1977e-01 Total = 2.6730e+03
Min = 0.0000e+00 Max = 2.0000e+00 Num pixels = 190969
```

The statistics have been calculated over the whole image. To restrict the calculation to just those pixels within the annulus we need a mask: an image the same size as img_stat.fits in which those pixels which should be considered for the calculation have values > 0. In this case we use the filtered exposure map (emap_stat.fits):

```
chips> sstats( "img_stat.fits", "emap_stat.fits" )
Statistics of img_stat.fits using (emap_stat.fits > 0) as a mask.
Mean = 2.7784e-02 RMS = 1.6761e-01 Total = 2.6730e+03
Min = 0.0000e+00 Max = 2.0000e+00 Num pixels = 96205
```

It is also possible to pass the routine "virtual files", so it is not necessary to create the images beforehand:

```
chips> sstats( "1838_img_s3.fits[sky=region(stat.reg)]", "expmap_1.7kev_7.fits[sky=region(stat.reg)
Statistics of 1838_img_s3.fits[sky=region(stat.reg)] using (expmap_1.7kev_7.fits[sky=region(stat.reg)
Mean = 2.7784e-02 RMS = 1.6761e-01 Total = 2.6730e+03
Min = 0.0000e+00 Max = 2.0000e+00 Num pixels = 96205
chips> exit
```

The reason for the small difference in the RMS values reported by dmstat and sstats() is because dmstat calculates

```
rms = sqrt( (mean(x*x) - mean(x)*mean(x)) / N )
whereas sstats() uses
```

```
rms = sqrt( (mean(x*x) - mean(x)*mean(x)) / (N-1) )
```

where x is a list of N numbers and mean(x) refers to the mean value of the list x.

Further examples

Above we show an example of looking at the signal level within the region in different energy bands, using the virtual file syntax to filter and bin the event list "on the fly". In this example, we do not apply any spatial filtering to the event list; instead we create a filtered, full-size image of the exposure map using the <u>DM [opt full]</u> syntax.

B. sstats.sl 5

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```
unix% dmcopy "expmap_1.7kev_7.fits[sky=region(stat.reg)][opt full]" emap_stat_full.fits
unix% more chip_s3.reg
# Region file format: CIAO version 1.0
rotbox(4182.860000,4393.150000,1024,1024,93.192576)
unix% chips
Welcome to ChIPS, version CIAO 3.4
Copyright (C) 1999-2003, Smithsonian Astrophysical Observatory
chips> () = evalfile( "sstats.sl" )
chips> sstats( "acisf01838N001_evt2.fits[energy=500:2000,sky=region(chip_s3.reg)][bin sky=1]", "emap_stat_
Statistics of acisf01838N001_evt2.fits[energy=500:2000,sky=region(chip_s3.reg)][bin sky=1] using (emap_sta
Mean = 1.0561e-02 RMS = 1.0273e-01 Total
Min = 0.0000e+00 Max = 2.0000e+00 Num pixels = 96205
chips> sstats( "acisf01838N001_evt2.fits[energy=2000:8000,sky=region(chip_s3.reg)][bin sky=1]", "emap_stat
Statistics of acisf01838N001_evt2.fits[energy=2000:8000,sky=region(chip_s3.reg)][bin sky=1] using (emap_st
as a mask.
Mean = 1.4583e-02 RMS = 1.2100e-01 Total
                                                 = 1.4030e+03
Min = 0.0000e+00 Max = 2.0000e+00 Num pixels = 96205
```

History

14 Dec 2004 updated for CIAO 3.2: path for script

20 Dec 2005 reviewed for CIAO 3.3: no changes

01 Dec 2006 reviewed for CIAO 3.4: ChIPS version

URL: http://cxc.harvard.edu/ciao/threads/image_stats/
Last modified: 1 Dec 2006

6 History

Image 1: Annulus in which we wish to measure the signal level

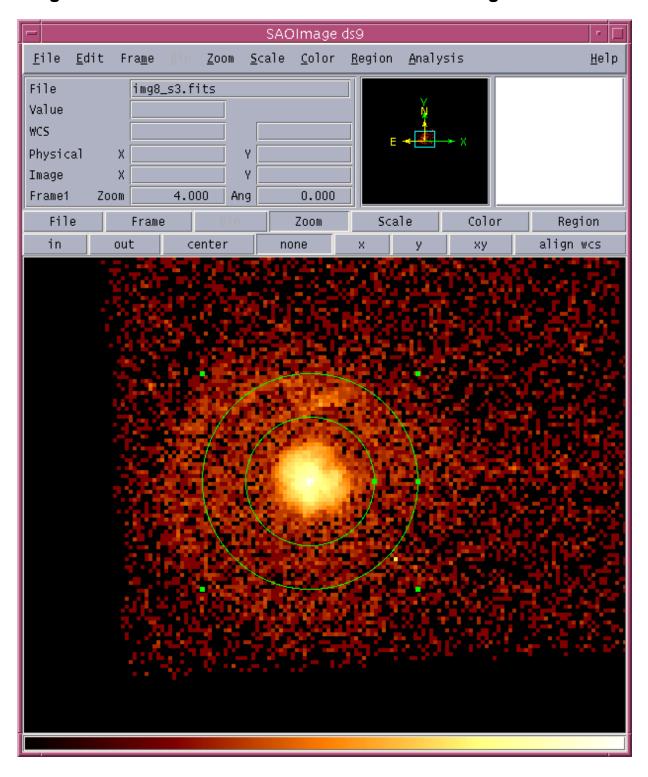
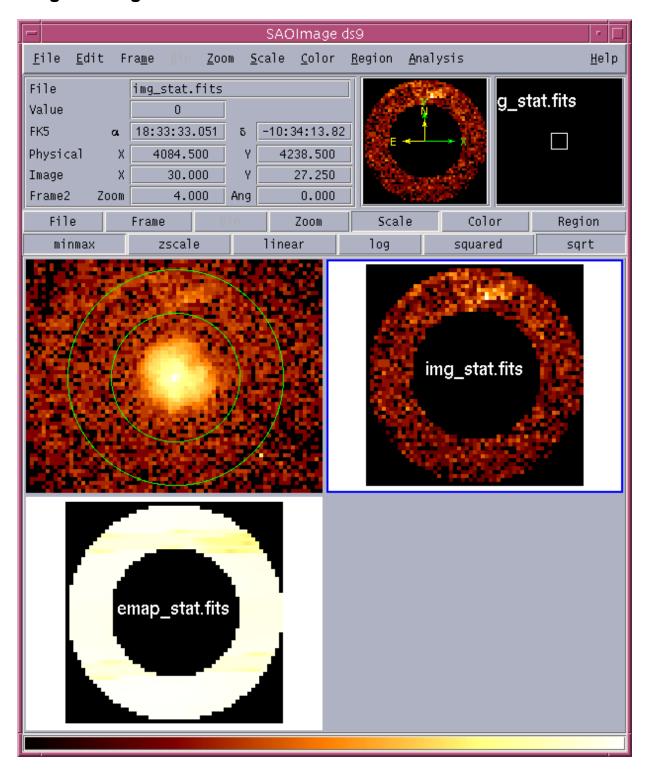


Image 2: Image mosaic



The top-left image is the original image (log scaling), the top-right image is the filtered image (sqrt scaling), and the bottom-left image shows the filtered exposure map (linear scaling).