

URL: http://cxc.harvard.edu/ciao3.4/xscevmkl.html

Last modified: December 2006

AHELP for CIAO 3.4

xscevmkl

Context: sherpa

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Synopsis

Multi-temperature vmeka. XSpec model.

Description

A multi-temperature plasma emission model built from the mekal code. Emission measures follow a power law in temperature (i.e. emission measure from temperature T is proportional to (T/Tmax)^alpha). The abundances are relative to the Solar abundances set by the xspecabundan command. The switch parameter determines whether the mekal code will be run to calculate the model spectrum for each temperature, or whether the model spectrum will be interpolated from a pre-calculated table; the former is slower but more accurate. The reference for this model is Singh et al. (1996, ApJ, 456, 766).

xscevmkl Parameters

Number	Name	Description
1	alpha	index for power law emissivity function
2	Tmax	maximum temperature
3	nН	nH (cm^-3) from mekal
4-17	(element)	Abundances for He, C, N, O, Ne, Na, Mg, Al, Si, S, Ar, Ca, Fe, Ni with respect to Solar
18	redshift	redshift, z
19	switch	0 = calculate, 1 = interpolate
20	norm	normalization

This information is taken from the XSpec User's Guide. Version 11.3.1 of the XSpec models is supplied with CIAO 3.2.

Bugs

For a list of known bugs and issues with the XSPEC models, please visit the XSPEC bugs page.

See Also

sherpa

atten, bbody, bbodyfreq, beta1d, beta2d, box1d, box2d, bpl1d, const1d, const2d, cos, delta1d, delta2d, dered, devaucouleurs, edge, erf, erfc, farf, farf2d, fpsf, fpsf1d, frmf, gauss1d, gauss2d, gridmodel, hubble, jdpileup, linebroad, lorentz1d, lorentz2d, models, nbeta, ngauss1d, poisson, polynom1d, polynom2d, powlaw1d, ptsrc1d, ptsrc2d, rsp, rsp2d, schechter, shexp, shexp10, shlog10, shloge, sin,

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sqrt, stephild, steplold, tan, tpsf, tpsfld, usermodel, xs, xsabsori, xsacisabs, xsapec, xsbapec, xsbbody, xsbbodyrad, xsbexrav, xsbexriv, xsbknpower, xsbmc, xsbremss, xsbvapec, xsc6mekl, xsc6pmekl, xsc6pmekl,

slang

usermodel

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