

URL: http://cxc.harvard.edu/ciao3.4/ptsrc2d.html Last modified: December 2006

AHELP for CIAO 3.4

ptsrc2d

Context: sherpa

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#### **Synopsis**

A 2–D file–based point–source fitting model.

## Description

PTSRC is a file-based model that may be used in fitting image data, to determine, e.g., if they are point-like or extended.

This model cannot be used as an instrument model.

FITS formats are allowed for the input model file. The current implementation requires that the image bins of the data and the PTSRC2D model file are the same.

Assume that the PSF is provided in a FITS image (file) of size  $N_x * N_y$ .  $N_x$  and  $N_y$  may be much larger than the PSF size in pixels. An optimally sized sub–image of size xsize \* ysize can speed up the computation of model amplitudes.

If xoff = yoff = 0, the sub-image is extracted from the center of the original image contained in file. The user may find that changing xoff and/or yoff can result in his or her being able to extract a smaller optimally sized sub-image if, say, the PSF is too asymmetric to be fit easily into a centered rectangle.

If norm is thawed its best-fit value for a point-like source will be approximately equal to the number of detected counts from the source.

The parameters xpos and ypos indicate where the center of the sub–image is to be placed in the data image that is to be analyzed, i.e., they should be set to the location of the source centroid. Note that the initial values of the xpos and ypos are estimated using the input dataset.

Number	Name	Description		
1	file	input FITS file name		
2	xsize	x-width of the subset region of kernel file to use in convolution		
3	ysize	y-width of the subset region of kernel file to use in convolution		
4	xoff	x-direction offset		
5	yoff	y-direction offset		
6	xpos	x-position of the centroid in data image coordinates		

#### **PTSRC Parameters**

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7 yr	pos y-position of the centroid in data image coordinates
8 no	norm normalization defined as a sum over the model image

### Example

The example below shows the initial parameter values for the PTSRC2D model. Note that the default xsize and ysize values are set to 32. After reading the data file these parameter can be adjusted to expand the image and include most of the PSF fraction. Notice after image command indicates the fraction of PSF included in the sub–image.

sherpa> ptsrc2d[p2]											
she	sherpa> show p2										
<pre>ptsrc2d[p2] (integrate: on)</pre>											
	Param	Туре	Value	Min	Max	Units					
1	file st	ring:	"none"								
2	xsize fr	rozen	32	1	1024						
3	ysize fr	rozen	32	1	1024						
4	xoff fr	rozen	0	-512	512						
5	yoff fr	rozen	0	-512	512						
б	xpos th	nawed	128.5	0.5	256.5						
7	ypos th	nawed	128.5	0.5	256.5						
8	norm th	nawed	1	0	1000						
a h a											
sne	rpa> pz.r		SIZ.IILS								
sne	rpa> snow	/ pz									
pus	rcza[pz]	(1116	egrale. on)	261-0	M = ==	The late of					
	Param	туре	value	MTU	Max	UNIUS					
1	file st	ring:	"psf2 fits"								
2	xsize fr	ozen	32	1	1024						
3	vsize fr	ozen	32	1	1024						
4	xoff fr	ozen	0	-512	512						
5	voff fr	ozen	0	-512	512						
6	xpos th	awed	128 5	0 5	256 5						
7	vpos th	awed	128.5	0.5	256.5						
8	norm th	nawed	1	0.5	1000						
she	rpa> imag	re p2	-	Ŭ	2000						
NOT	E: PSF fr	raction	n for (xsize.vs	size): FRA	AC = 0.9847	33					
sherpa> p2.xsize=64											
she	sherpa> p2.ysize=64										
she	sherpa> image p2										
Shorter Twode Fr											

#### Bugs

See the <u>Sherpa bug pages</u> online for an up-to-date listing of known bugs.

# See Also

sherpa

atten, bbody, bbodyfreq, beta1d, beta2d, box1d, box2d, bpl1d, const1d, const2d, cos, delta1d, delta2d, dered, devaucouleurs, edge, erf, erfc, farf, farf2d, fpsf, fpsf1d, frmf, gauss1d, gauss2d, gridmodel, hubble, jdpileup, linebroad, lorentz1d, lorentz2d, models, nbeta, ngauss1d, poisson, polynom1d, polynom2d, powlaw1d, ptsrc1d, rsp, rsp2d, schechter, shexp, shexp10, shlog10, shloge, sin, sqrt, stephi1d, steplo1d, tan, tpsf, tpsf1d, usermodel, xs, xsabsori, xsacisabs, xsapec, xsbapec, xsbbody, xsbbodyrad, xsbexrav, xsbexriv, xsbknpower, xsbmc, xsbremss, xsbvapec, xsc6mekl, xsc6pmekl, xsc6pvmkl, xsc6vmekl, xscemekl, xscemekl, xscemekl, xscflow, xscompbb, xscompls, xscompst,

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xscomptt, xsconstant, xscutoffpl, xscyclabs, xsdisk, xsdiskb, xsdiskline, xsdiskm, xsdisko, xsdiskpn, xsdust, xsedge, xsequil, xsexpabs, xsexpdec, xsexpfac, xsgabs, xsgaussian, xsgnei, xsgrad, xsgrbm, xshighecut, xshrefl, xslaor, xslorentz, xsmeka, xsmekal, xsmkcflow, xsnei, xsnotch, xsnpshock, xsnsa, xsnteea, xspcfabs, xspegpwrlw, xspexrav, xspexriv, xsphabs, xsplabs, xsplcabs, xsposm, xspowerlaw, xspshock, xspwab, xsraymond, xsredden, xsredge, xsrefsch, xssedov, xssmedge, xsspline, xssrcut, xssresc, xssssice, xsstep, xstbabs, xstbgrain, xstbvarabs, xsuvred, xsvapec, xsvarabs, xsvbremss, xsvequil, xsvgnei, xsvmcflow, xsvmeka, xsvmekal, xsvnei, xsvnpshock, xsvphabs, xsvpshock, xszraymond, xsvsedov, xswabs, xswndabs, xsxion, xszbbody, xszbremss, xszedge, xszgauss, xszhighect, xszpcfabs, xszphabs, xszpowerlw, xsztbabs, xszvrabs, xszvfeabs, xszvphabs, xszwabs, xszwndabs

slang

<u>usermodel</u>

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