Timing Analysis with Chandra

Michael Nowak, MIT-CXC (& thanks to Glenn Allen, Peter Ford, Herman Marshall ...)

See help threads at:

http://cxc.harvard.edu/ciao/documents_threads_timing.html

And keep an eye on:

http://space.mit.edu/CXC

Hey?! Isn't That RXTE's Job? Yes, but ... Chandra Can:

- Observe Crowded Fields
- Observe the 0.1 2 keV Energy Band
- Obtain 10⁻⁵ Hz 10⁻³ Hz (not since EXOSAT!)
- Faint Objects (*Single* Photons Matter!)

And It Can do msec Timing Too!

Normalizing Power as (RMS)²/Hz :

(Noise Limit * df)^{1/2} = (2/R) ^{1/2} (1 + B/R)^{1/2} (df / T)^{1/4},

i.e., RMS limit in *averaged* frequency bin, where: R = Signal Count Rate B = *Constant* Background Count Rate T = Total Observation Length *df* = Width of Frequency Bin

Example: T = 40 ksec, RMS Limit ~10% R^{-1/2} $(1 + B/R)^{1/2} (df / 1 Hz)^{1/4}$

Reducing Background Can Really Help!

Fundamental Times: ACIS, Timed Exposure

• Frame Time:

T (msec) = (41 + 0.040 * q)*m + 2.84*n + 5.2,

q = # of rows from readout

m = # of active chips

n = # of rows read

- Reality: Frame Time is *Integer* Multiple of 0.1 sec (0.2 – 10 sec) + 41.04 msec
- Caveat: Images are Transferred to Frame Store (Quasi-) Serially, so up to a 5*41.04 msec Delay Between Chips
- Event Times are *Middle* of Frame Time

Fundamental Times: ACIS, Timed Exposure

- Frames Take 41.04 msec to Transfer to Frame Store, so a Given Amount of (Uniform) Deadtime is Expected
- Charge Moved at 40 µsec/row, which Gives the Potential for Very Fast Timing of Readout Streaks
- (Sources that bright will otherwise be difficult to deal with...)

Fundamental Times: ACIS, Timed Exposure

- Event Time is *Terrestrial Time*, Referenced to: MJD = 50814.0 (January 1, 1998)
 - MJD = Julian Date 2,400,000.5 (TJD = MJD – n * 10,000. *Don't Use!*)
- XMM: Same Reference MJD
- *RXTE* Referenced to: MJD = 49353.000696574074 (January 1, 1994) (fraction is ~ 1 minute, i.e. Terrestial Time vs. UTC)
- ASCA is Referenced to: MJD = 48988.0

(January 1, 1993)

Check FITS Headers!

None of These Times are Barycentered! (Requires Orbital Ephemeris; *axbary*, *fxbary*)

- Useful Links for Understanding/Converting:
- NIST Time Glossary -

http://www.bldroc.gov/timefreq/

• Date Conversion Utility -

http://heasarc.gsfc.nasa.gov/cgi-bin/Tools/DateConv/dateconv.pl

• Look at *ahelp* file for *axbary*

Fundamental Times: ACIS, Continuous Clocking

- Rows are Read Out Every 2.85 msec
- Chips are Read in Parallel
- Time is Read Out Time, *Not* Arrival Time. (Read out delay from aim point, modulo dither, etc.) Tool (from G. Allen) to Correct This Almost Ready.
- Generalization to CC-Gratings Observation Should Be Straightforward
- 40 µsec row shift "deadtime" still applies

Fundamental Times: HRC-S

- HRC-I Wiring Problem Limits Time to ~4 msec on Average
- HRC-S Can Achieve 16 µsec Accuracy
- Faster Timing than ACIS, but more severe Telemetry Limits (184 cps), and Higher Backgrounds
- But, no Deadtime, and HRC-S is Linear Up to at Least 5 cps for a Point Source

Tools at Your Disposal:

- CIAO Tools: *lightcurve*, *axbary*, threads on filtering lightcurves, creating phase-folded spectra
- Data products can be further analyzed with *XRONOS*, *IDL*, *S-Lang*, ...
- More to be Developed. What Do Users Want?

Self-Promoting Example:

- *lightcurve* corrected with *axbary*, allowed comparison to previous observations
- Folded Spectrum Created with CIAO Tools, Following Threads
- 0 Background! df ~ 5 x 10⁻⁵ Hz , RMS > 4%
- Single photon events used to determine rapidity of eclipse, yielding limits on source size



Other Examples Have Included:

- Crab Pulsar (Easily Detected)
- RX J185635-375 (astro-ph/0204159): Strong Upper Limits to Pulse Fraction
- Important Caveat/Strong Suggestion: Bin on Integer Multiples of "Natural" Time Unit, Watch Out for Instrumental Time Scales (dither, etc.)