

The XMM-Newton View of the $z < 0.5$ Warm-Hot Intergalactic Medium

F. Nicastro (OAR-INAF)

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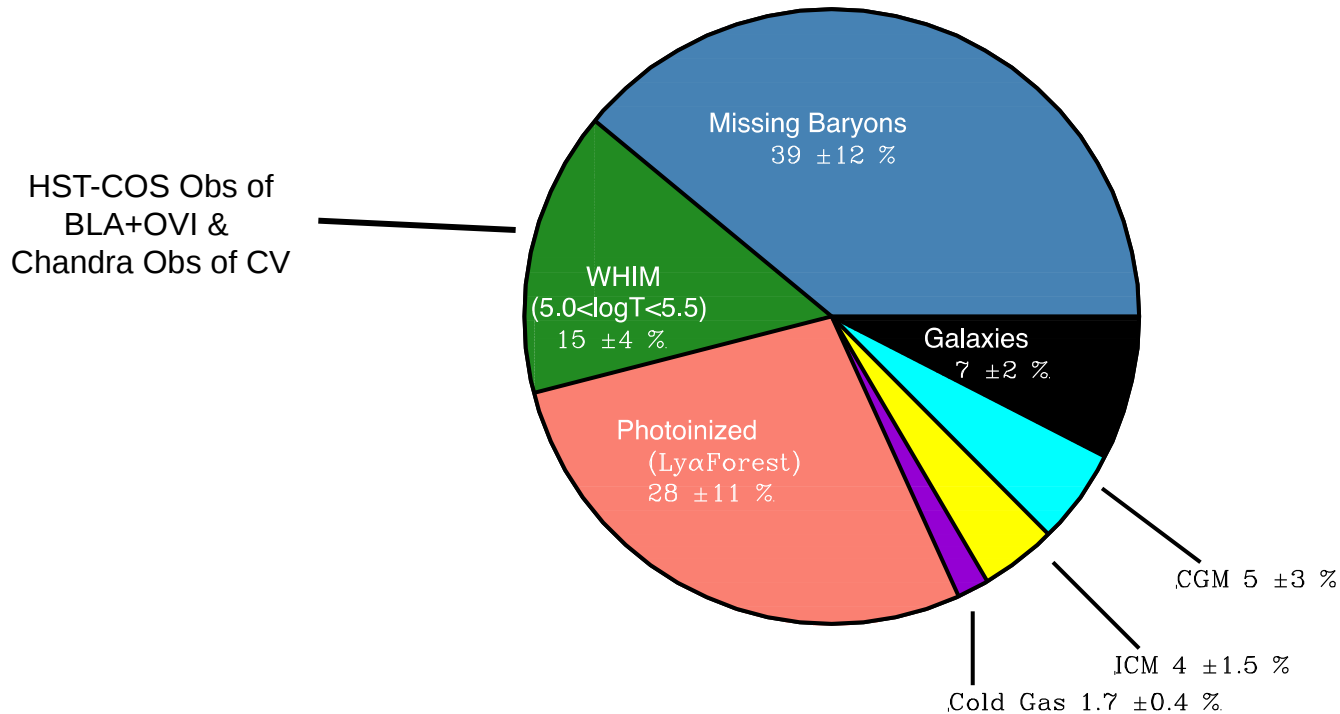
Outline

- The Missing Baryon Problem
- Preliminary Results from the XMM-Newton VLP on 1ES 1553+113
- From XMM/Chandra to Athena/Lynx

The Missing Baryons Problem

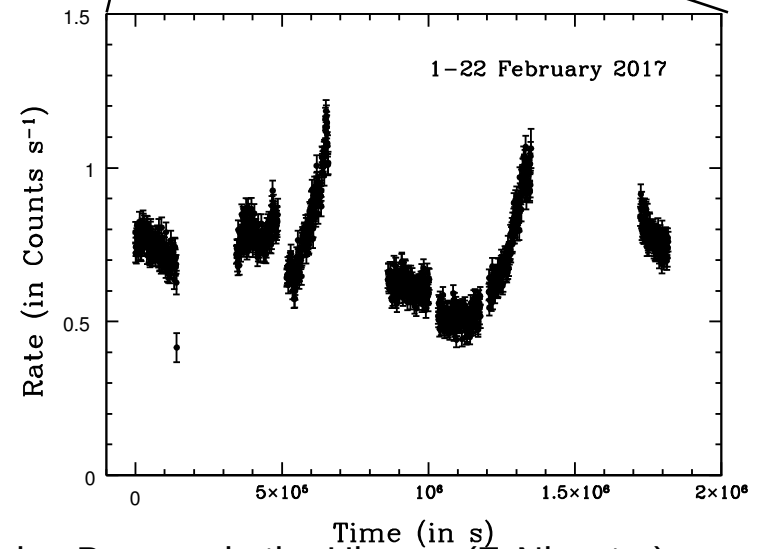
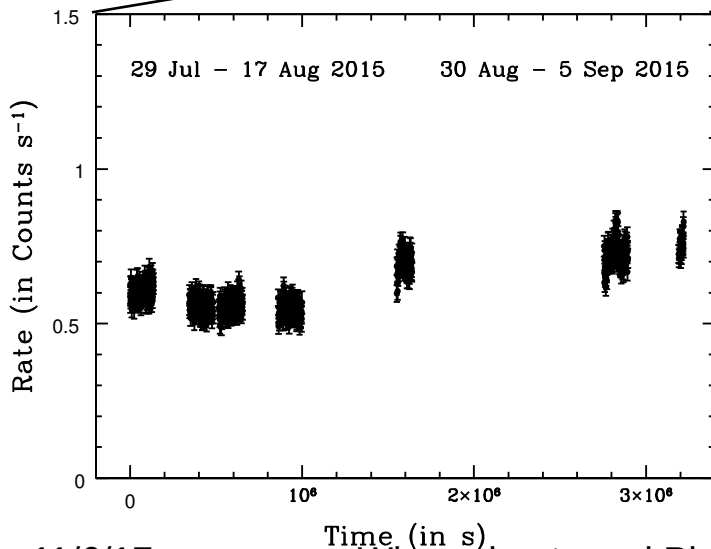
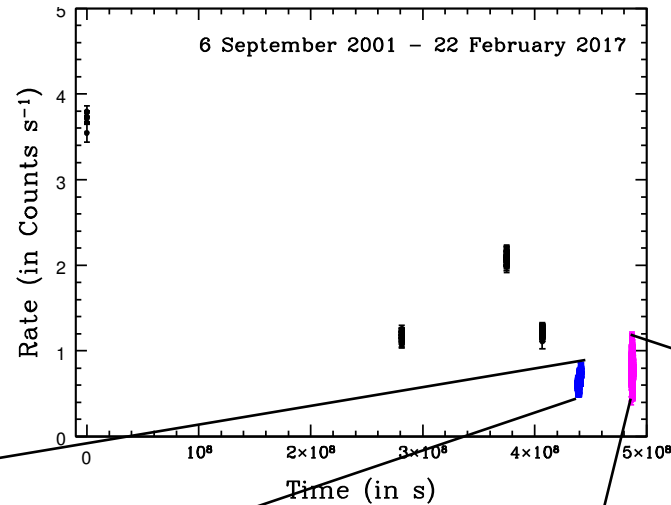
Nicastro+16

$$\Omega_b^{\text{WMAP}} h^{-2} = 0.0226 h^{-2} = 0.0456 \sim 5\%$$



~ 30-50% (or more) of Baryons Still Missing at z~0

The XMM-Newton VLP on 1ES 1553+113



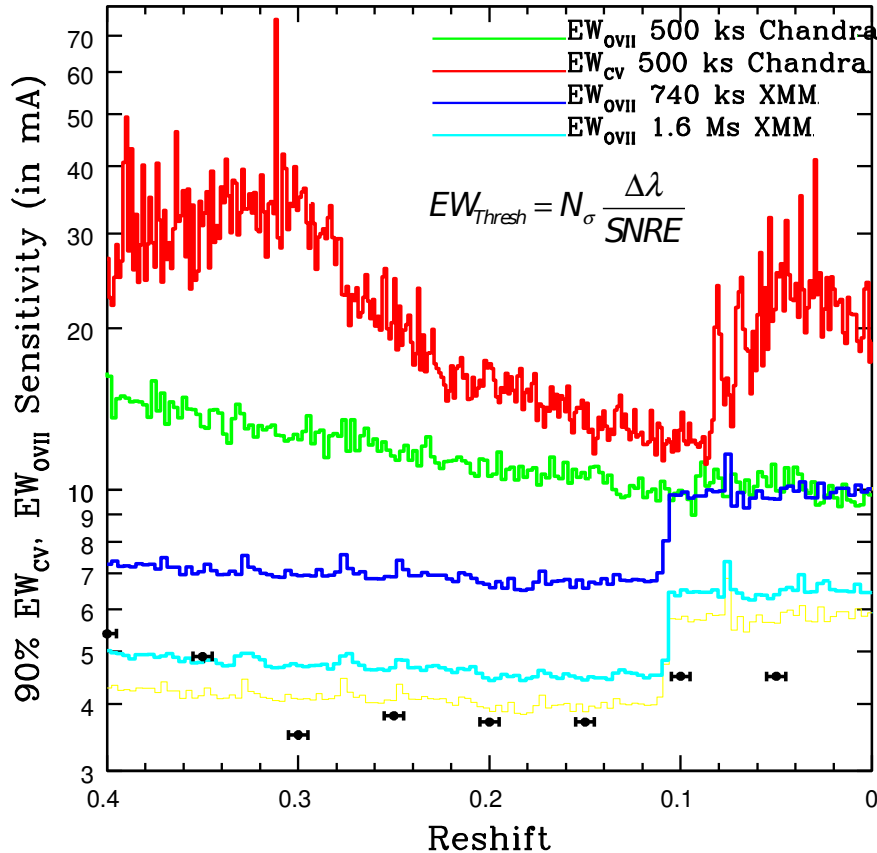
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Whereabouts and Physics of the Roaming Baryons in the Universe (F. Nicastro)

The Warm-Hot (OVII) IGM

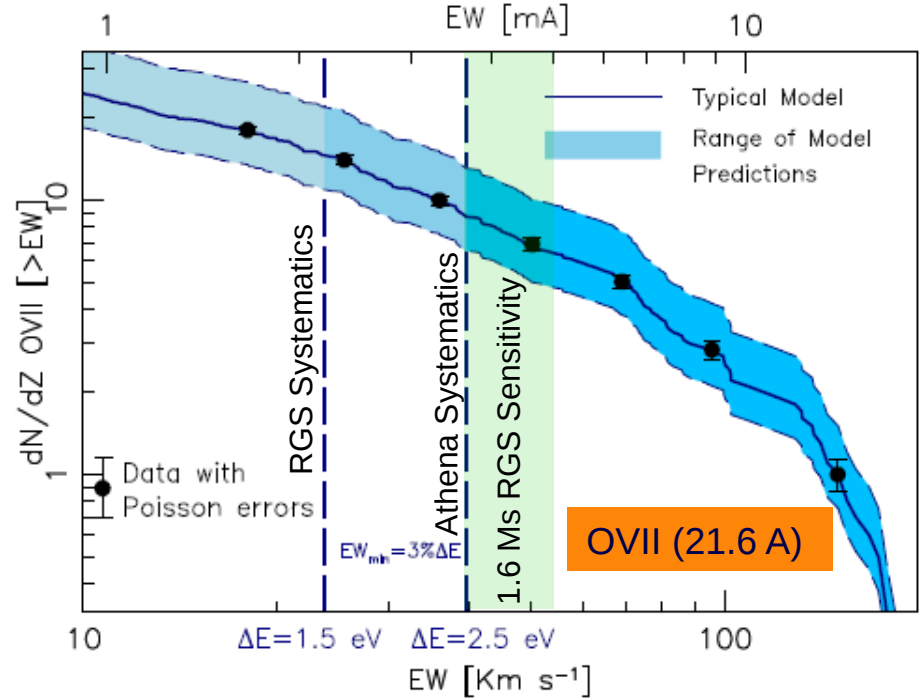
XMM-Newton RGS Spectrum of 1ES 1553+113

X-Ray Spectra of 1ES 1553+113
($z \approx 0.49 \pm 0.04$)



1.6 Ms RGS: EW > 4-5 mA @ > 90%
i.e. ~1200 cts per R.E.

Athena WHIM White Paper (Kaastra+13)



Expected:

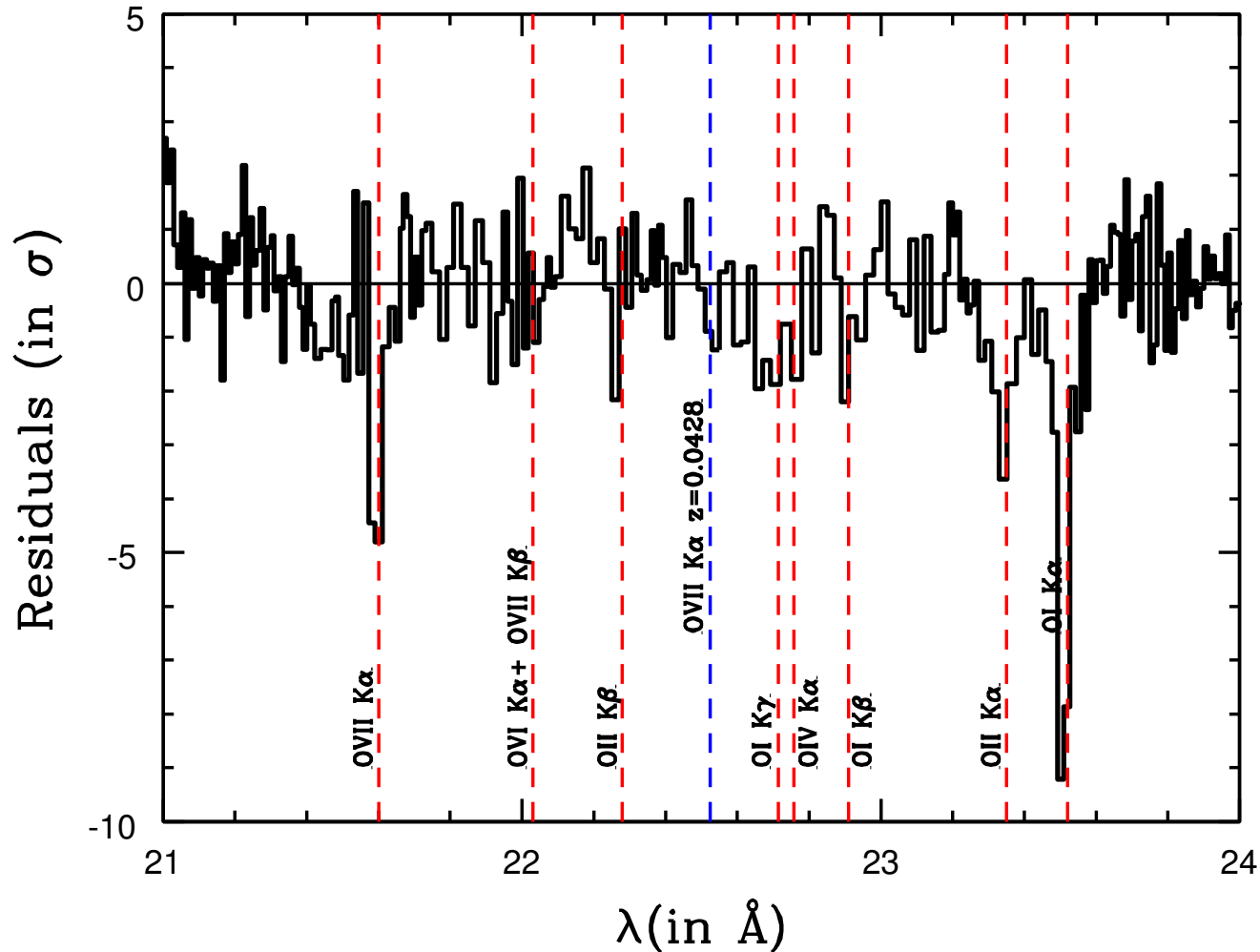
3.1-6.2 down to 90% RGS sensitivity &
up to $z < 0.53$

$P(>0) = [1 - P(3.1-6.2 | 0)] = 99.56-99.98\%$

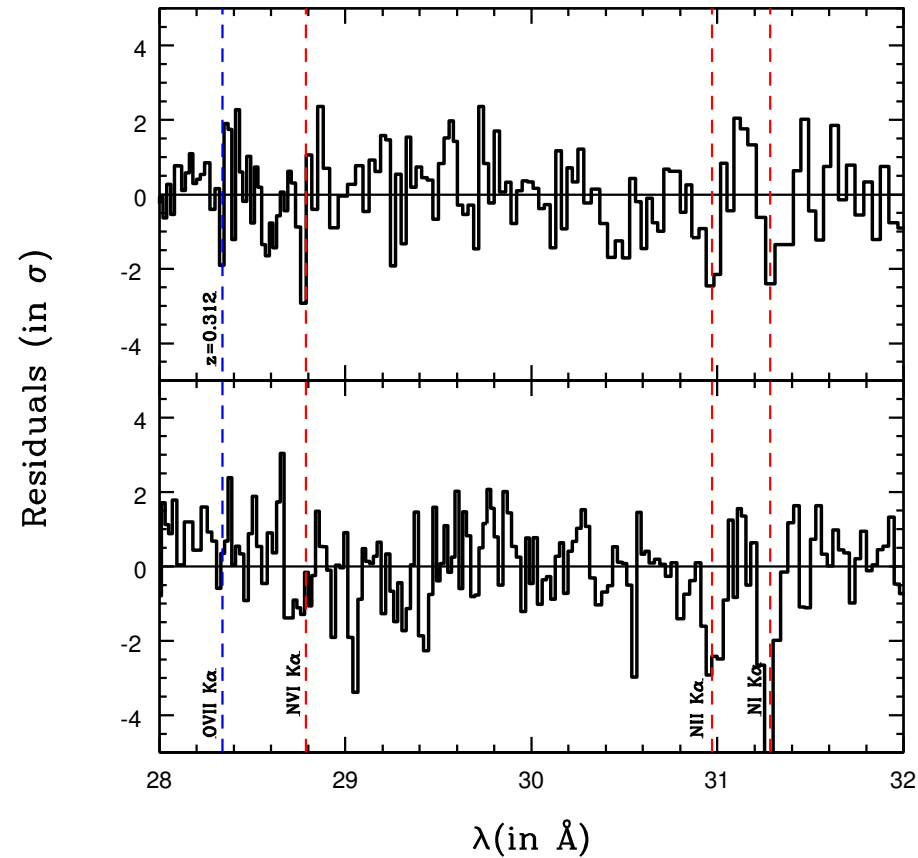
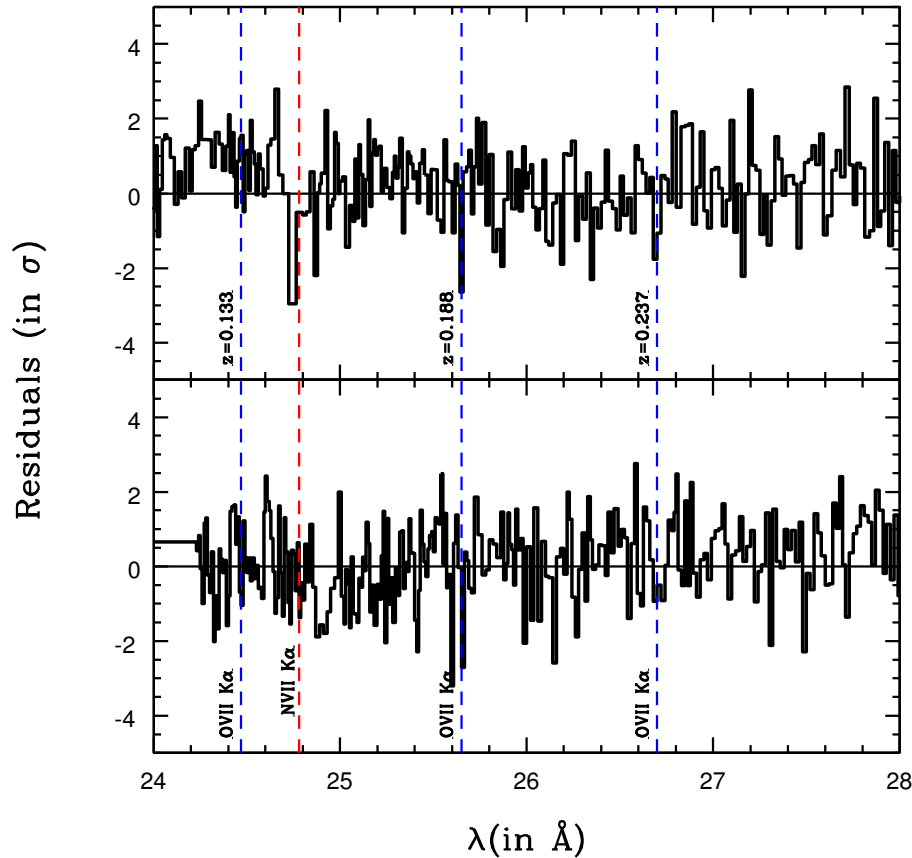
□ 2.85-3.75 σ

Physics of the Roaming Baryons in the
Universe (F. Nicastro)

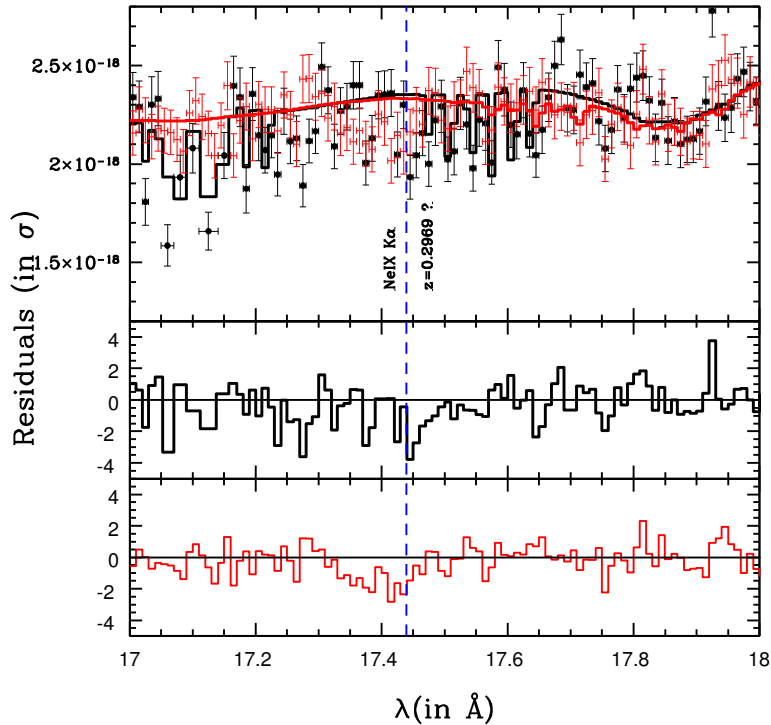
Galactic Oxygen Absorption



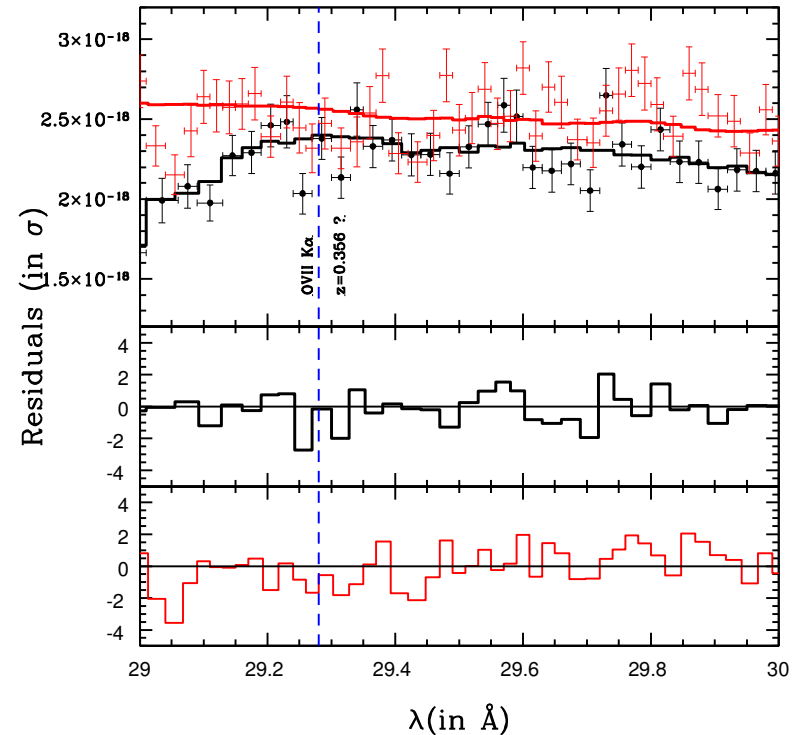
Galactic Nitrogen Absorption



Serendipitous Abs. Lines at $\lambda = 12-38 \text{ \AA}$ (in clean effective area regions)

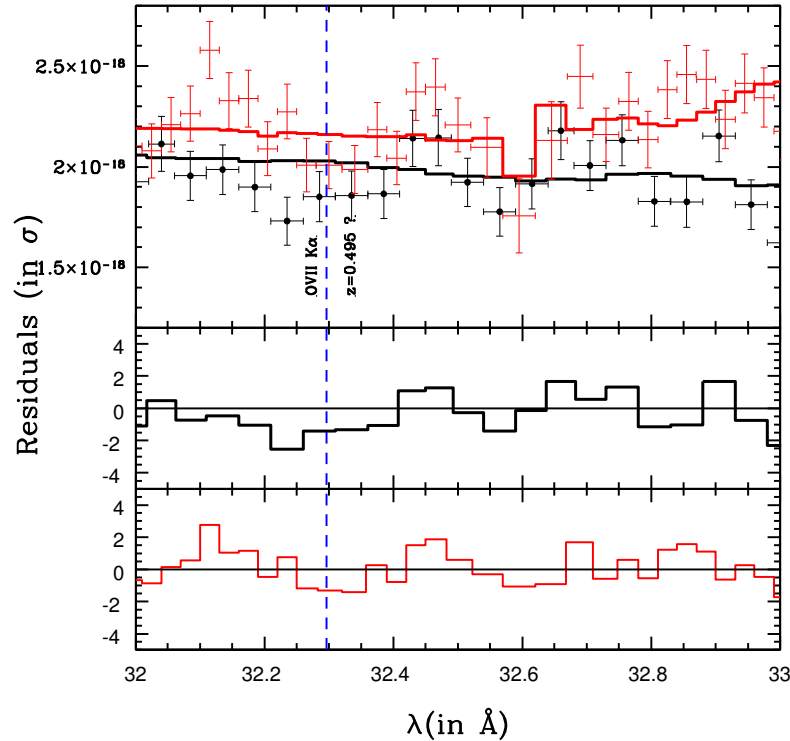


$\lambda=17.439(+0.006,-0.010) \text{ \AA}$
 $EW=12.0\pm 1.7 \text{ m\AA}$ ($\square 7\sigma$)
 If NeIX K α $\square z=0.2969$
 If NeX K α $\square z=0.4371$
 If FeXVII L $\square z=0.1614$

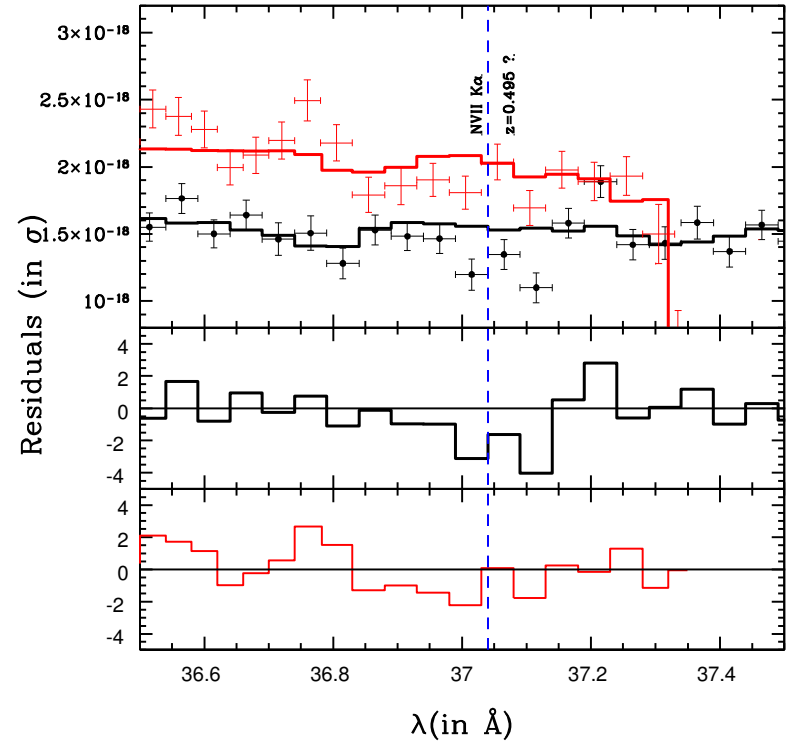


$\lambda=29.28\pm 0.02 \text{ \AA}$
 $EW=9.0(+3.9,-2.0) \text{ m\AA}$ ($\square 4.5\sigma$)
 If OVI K α $\square z=0.356$
 If NVII K α $\square z=0.182$
 If NVI K α $\square z=0.017$

Serendipitous Abs. Lines at $\lambda = 12-38 \text{ \AA}$ (in clean effective are regions)

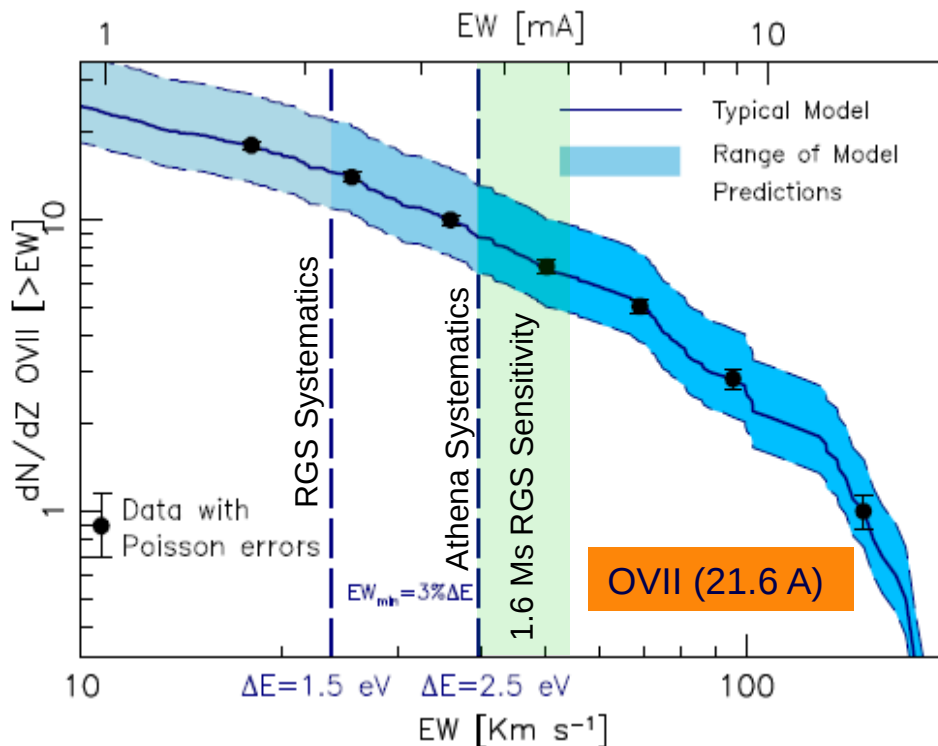


$\lambda = 32.296(+0.009, -0.032) \text{ \AA}$
 $EW = 11.8 \pm 3.7 \text{ m\AA} (\square 3.2\sigma)$
 If $OVII \text{ K}\alpha \square z = 0.495$
 If $NVI \text{ K}\alpha \square z = 0.122$
 If $NVII \text{ K}\alpha \square z = 0.303$
 If $NI \text{ K}\alpha \square z = 0.032$



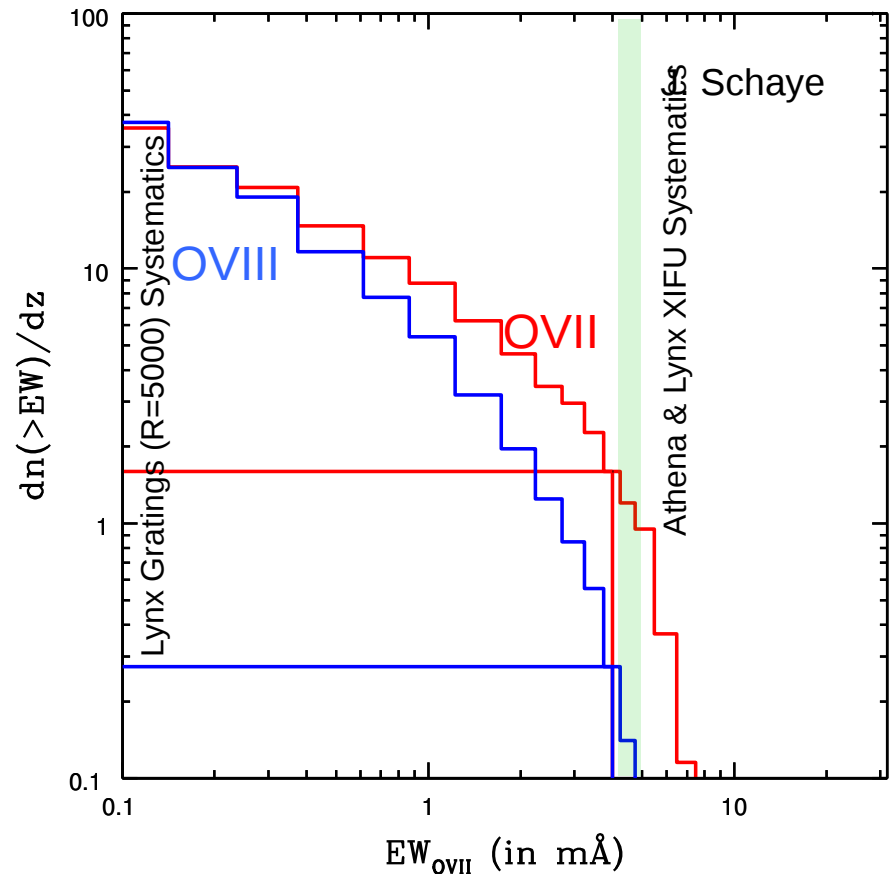
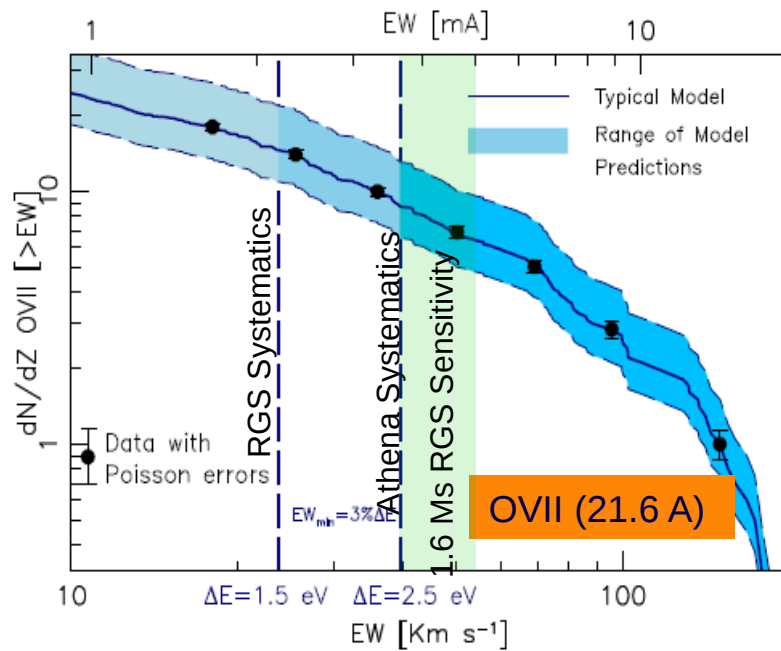
$\lambda = 37.04(+0.01, -0.03) \text{ \AA}$
 $EW = 21.0 \pm 4.6 \text{ m\AA} (\square 4.6\sigma)$
 If $NVII \text{ K}\alpha \square z = 0.495$
 If $NVI \text{ K}\alpha \square z = 0.287$
 If $NI \text{ K}\alpha \square z = 0.184$

The Future: from XMM/Chandra to Athena/Lynx



- 1-3 $\log T > 5.5$ IGM systems (1-2 in $\text{O VII } K\alpha$) have possibly been seen thanks to the deepest XMM observation of a blazar (that hits the predicted Athena systematics)
- This is (still) barely consistent with the most conservative HDS predictions, and certainly challenges optimistic predictions
- But a single line of sight! Cosmic Variance needs to be addressed: going deeper would not pay
- Shallower ($EW \geq 8 \text{ mÅ}$ at 90%) observations of a number (5-10) properly selected high- z targets are needed and would greatly benefit feasibility studies for Athena.

Which Predictions are Correct?



Our Line of Sight seems to Suggest Eagle Simulations give right predictions