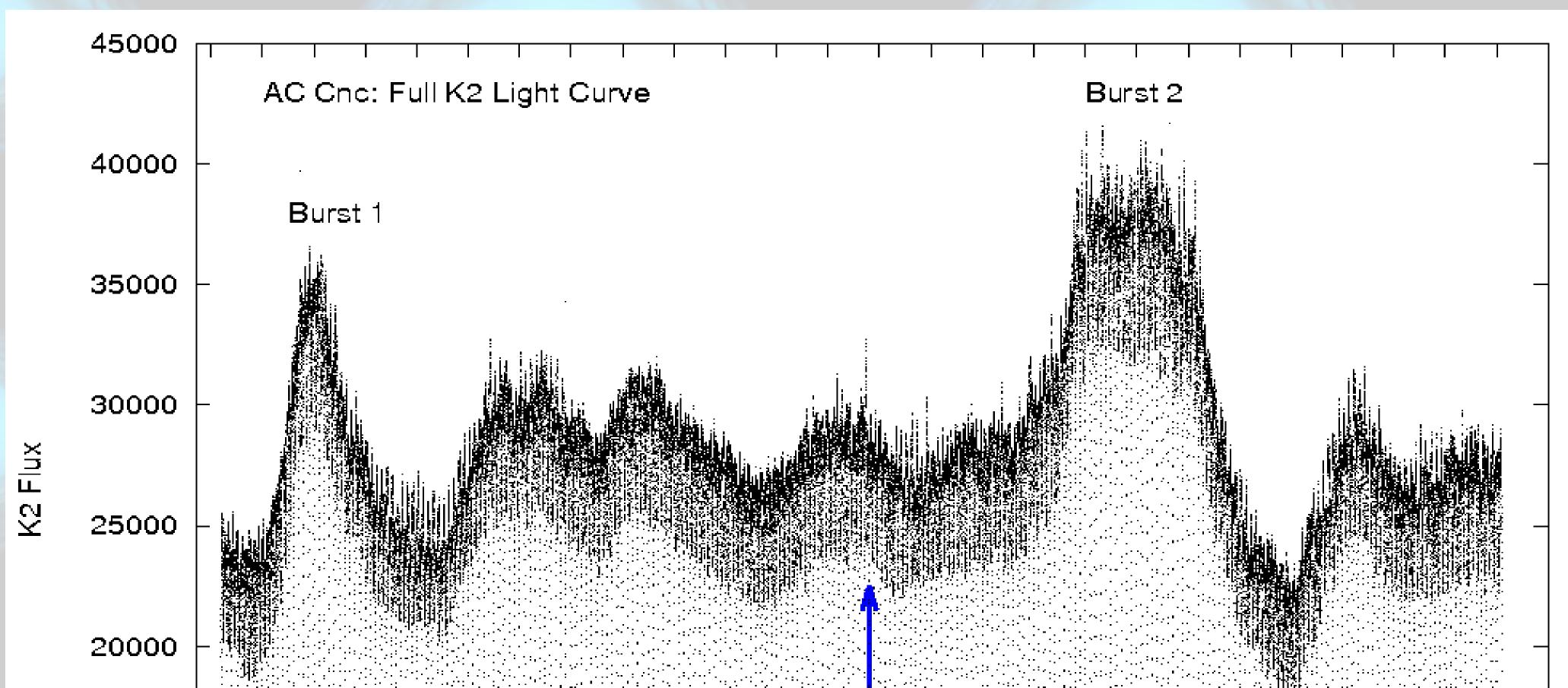
## The Cataclysmic Variable AC Cnc: Accretion Variations and Stunted Bursts Revealed by Kepler2 Eric M. Schlegel<sup>1</sup> and R. Kent Honeycutt<sup>2</sup> <sup>1</sup>Vaughan Family Professor, Physics & Astronomy, The University of Texas at San Antonio <sup>2</sup>Department of Astronomy, Indiana University



Abstract: We describe the K2 observation, using Kepler at 1-minute cadence, of the nova-like cataclysmic variable AC Cnc. The K2 data set lasted 74 days, covered 246 eclipses, and yielded 106,703 photometric measurements. This CV exhibits stunted outbursts and we describe them in relation to other stunted-bursting CVs. The light curve shows no evidence of disk brightening before either burst, suggesting that mass transfer is not the answer.

<<<< The K2 light curve of AC Cnc. There are two stunted bursts and a continuous evolution in and out of each burst.

Plot of FWHM vs amplitude for the stunted

117 bursts of AC Cnc (K2 = red; ground = green) and other CVs exhibiting stunted 15000 bursts. The value of the K2 observations for stunted bursts: the `continuous' nature of the observations reveals a constantly changing accretion disk. The idea of a 10000 Bottoms, Primary Eclipses Bottoms, Secondary Eclipses `disk state' may be a misleading viewpoint. 100 50 150 200 250 Ο Orbit Phase AC Cnc (K2) AC Cnc (grd) Ж V368 Aql 25 V405 Aur 1:10 35 UU Aqr 25 Q Cyg Ο 20 Ο 11:20 21:30 V426 Oph 15 FY Per FWHM (days) 31:40 41:50 35 51:60 RW Tri ▲ 15  $\odot$ 25 **ж**⊡ Ω 15 1000 Θ 10 71:80 81:90 61:70 × Counts ×+\_\_\_\_ 101:110 111:120 35 91:100

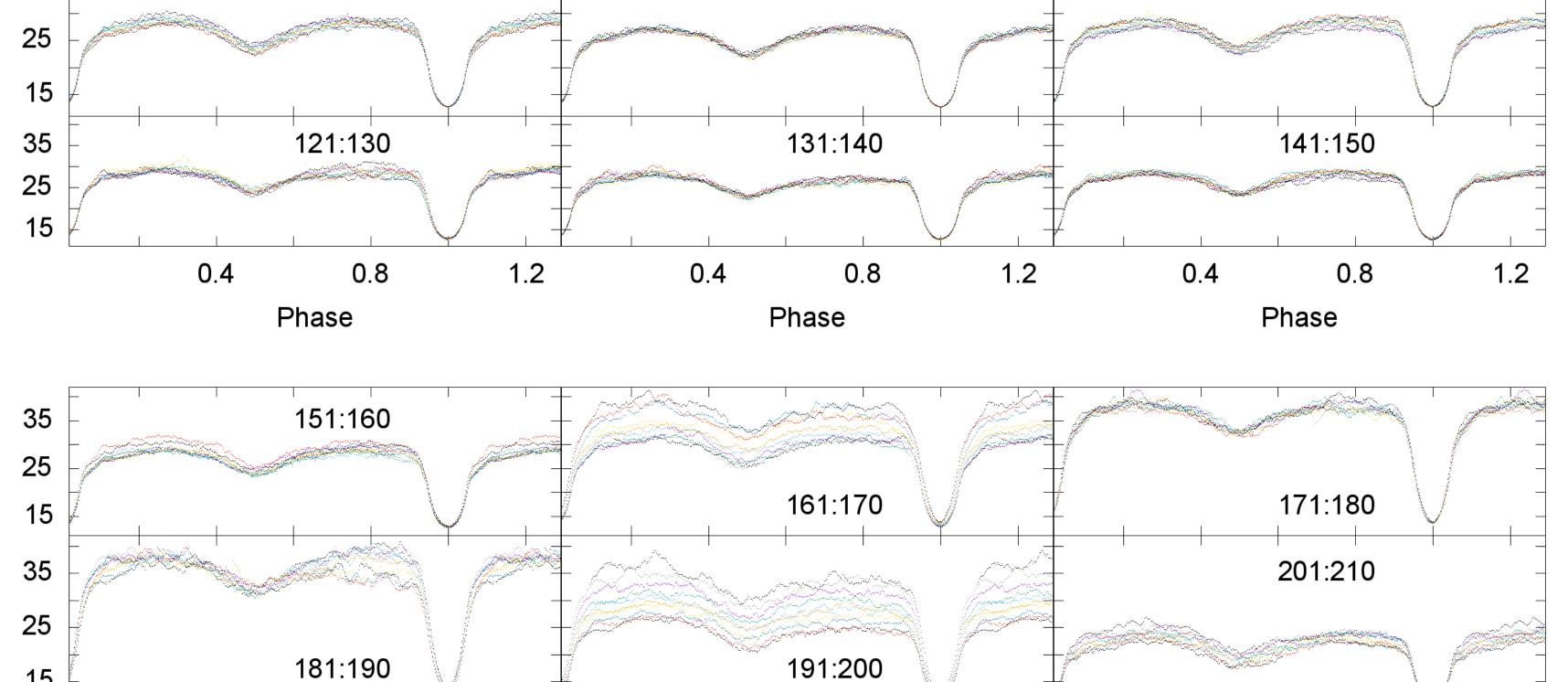
231:240

0.8

Phase

0.4

1.2



221:230

0.8

Phase

0.4

1.2

1.2

181:190

211:220

241:246

0.8

Phase

0.4

15

1000

 $\boldsymbol{\times}$ 

Counts

35

25

15



Essentially *none* of the orbits reveal evidence of a pre-eclipse bright spot. Further, an enhanced bright spot does not occur prior to either burst. This essentially requires the mass transfer to be steady, up to and through the burst.

Yet the disk clearly responds – compare orbits 1-10 with 11-20 or 151-160 versus 161-170. Furthermore, the eclipse widths do not change, so any disk `enhancements' must occur without disturbing the eclipse profile.

This is a puzzle. We are working to generate eclipse maps for all of the K2 data and use them to investigate the secondary eclipses. Stay tuned!

