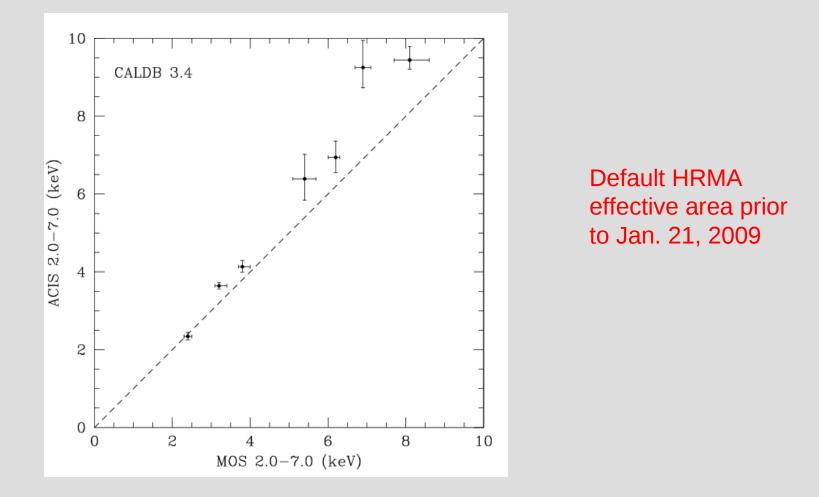
XMM-Newton/ Chandra Cross Calibration with Clusters of Galaxies

L. David & J. Nevalainen

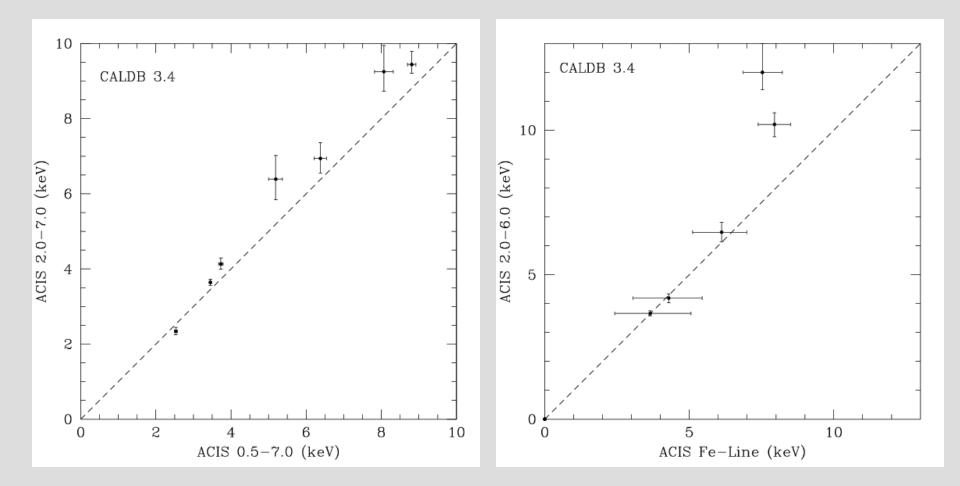
CXC Calibration Review September 21 2009

Plot from 2 annual IACHEC Meeting

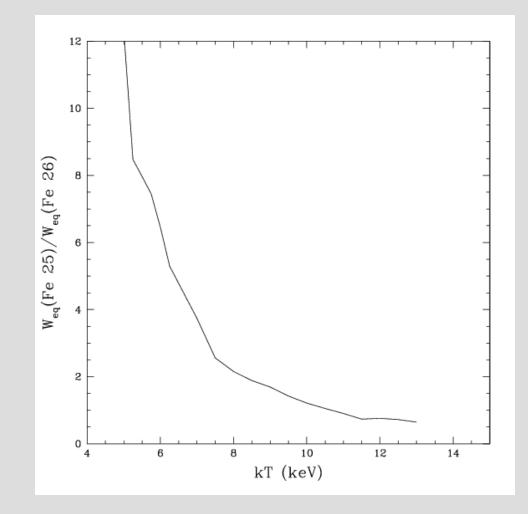
Comparison of XMM-Newton and Chandra derived Cluster Temperatures



Internal comparison of ACIS derived temperatures

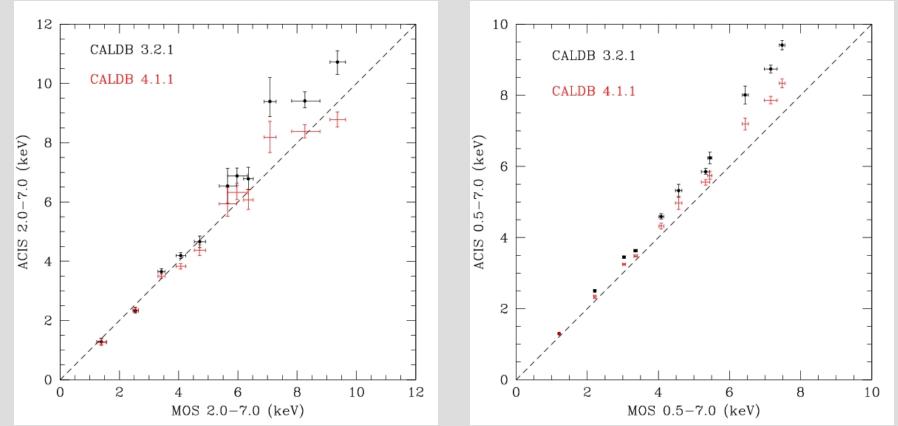


Cluster Temperature vs. Fe-line Ratio

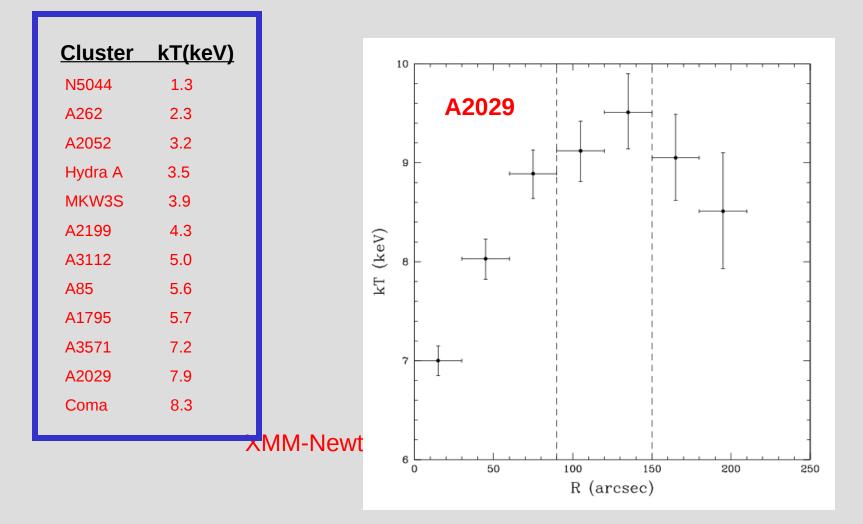


Comparison between ACIS and MOS derived temperatures in the 2-7 keV band.

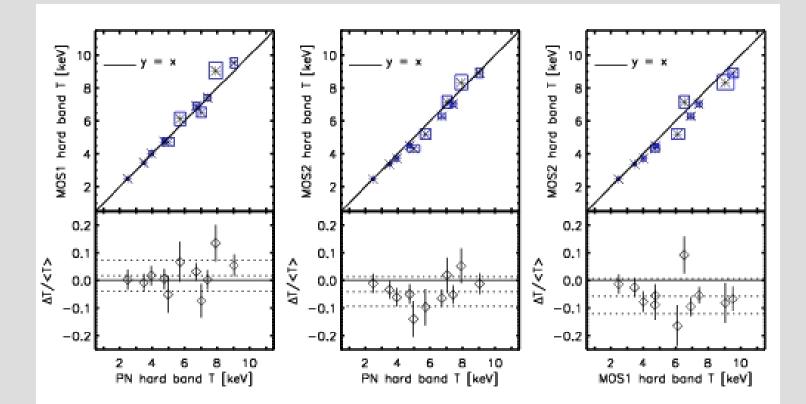
Comparison between ACIS and MOS derived temperatures in the 0.5-7 keV band

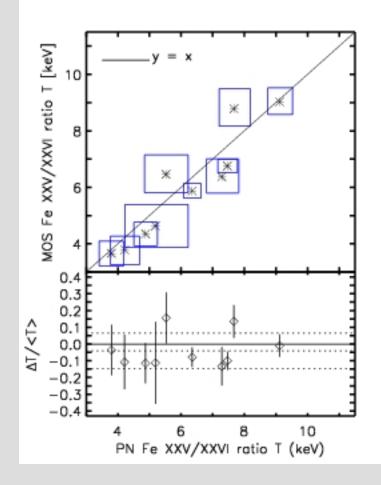


Chandra/XMM-Newton Cluster Cross-Calibration using SASS v8.1 and CALDB 4.1.1

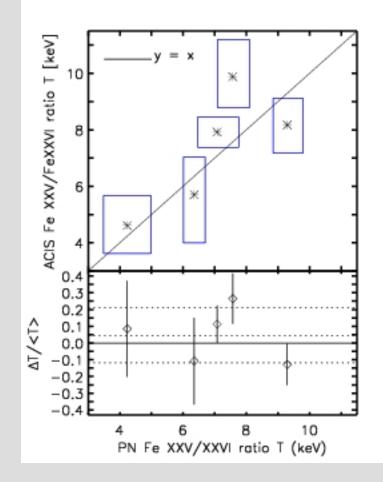


XMM-Newton Internal Cross-Calibration



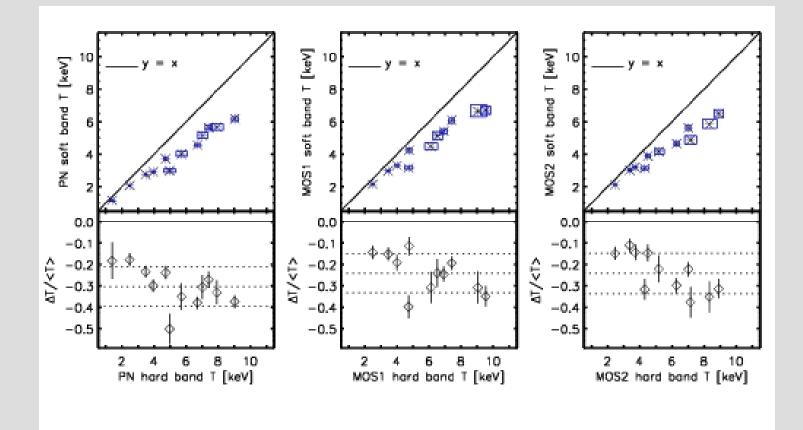


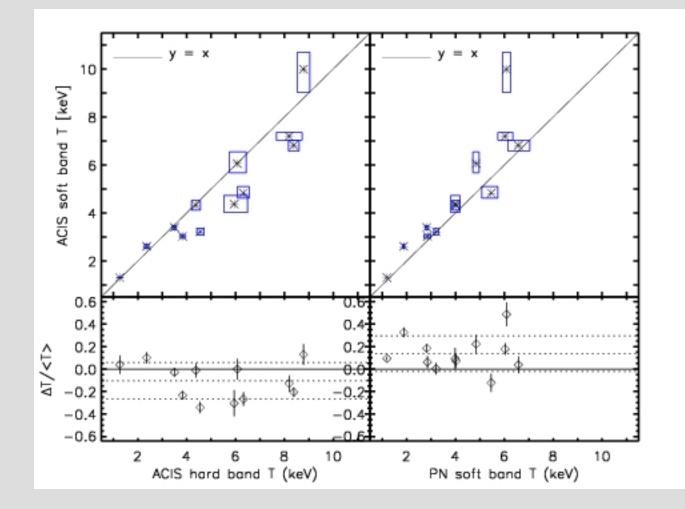
Comparison between MOS and PN derived temperatures from the Fe-line ratio.



Comparison between ACIS and PN derived temperatures from the Fe-line ratio.

XMM-Newton Internal Comparison (Soft vs. Hard Band Temperatures)



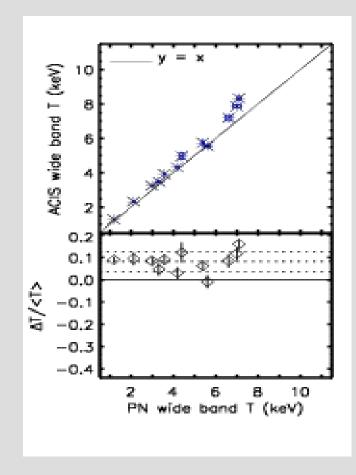


E0102 Cross-Calibration



Similar trend as derived from Cluster temperatures in the 0.5-2.0 keV energy band

Chandra/XMM broad energy band comparison



Summary

Hard energy band (2.0-7.0 keV)

- MOS1 temperatures are approximately 5% greater than MOS2 temperatures
- MOS1 and PN temperatures are essentially consistent
- ACIS, MOS1 and PN temperatures are essentially consistent

Soft (0.5-2.0 keV) vs. Hard energy bands

ACIS - soft band temperatures are 10% less than hard band temperatures MOS1 and MOS2 - soft band temperatures are 20% less than hard band temperatures PN - soft band temperatures are 30% less than hard band temperatures

Broad energy band

ACIS temperatures are 5% greater than MOS temperatures MOS temperatures are 5% greater then PN temperatures