The Remarkable X-ray Jet in the Quasar 4C 20.24

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PKS 1055+201=4C 20.24

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Outline

1. Context of an X-ray Jet Survey

- Flat spectrum radio sources
- Extended radio jet longer than 2"
- Predict detectable X-ray flux in 5ks

- 2. What is Normal About 4C20.24?
- **3. What is Remarkable About 4C20.24?**

Outline

- 1. Context of an X-ray Jet Survey
- 2. What is Normal About 4C20.24?
 - X-ray jet correlates with radio jet
 - X-rays modelled as IC/CMB
 - Magnetic field strength B about 10 μ G,
 - Doppler factor δ about 6
- 3. What is Remarkable About 4C20.24?

Outline

- 1. Context of an X-ray Jet Survey
- 2. What is Normal About 4C20.24?
- 3. What is Remarkable About 4C20.24?
 - Extended X-ray emission tube symmetric around radio/X-ray jet
 - Similar extended X-ray emission tube around unseen counter-jet
 - Jets appear to be "swept" back from quasar
 - Test case of seeing both the X-ray jet and its interaction with external medium?

The Jet Sample

- Flat Spectrum Quasars. Two Samples: $S_{5GHz} > 1Jy^a$ or $S_{2.7GHz} > 0.34 Jy^b$
- Radio Maps with< 2"resolution have jets >2" with detection expected by analogy to PKS 0637-752.



PKS 1055+201=4C 20.24 z=1.11

0.5–7 keV

21''= 173 kpc

1.4 GHz











Broad X-ray Tube Surrounds Jets







System is very symmetric North to South

PKS 1055+201

4GH2 VLA-B

28"=230kpc

0

0

0

Both Jets are swept to West



System is very symmetric North to South

Jet spectrum is softer than Tube spectrum



Extended X-ray "Tube"



Extended X-ray "Tube"

Thermal Interpretation $L_x=2.5 \ 10^{44} \ ergs \ s^{-1}$ kT = 1.34 keV $n_e = 0.0054 \text{ cm}^{-3}$ $t_{gas} \approx 4 \ 10^9 \ years$ $P_{gas} \approx 1.2 \ 10^{-11} dyne \ cm^{-2}$ $P_{gas} \approx P_{iet}$ $U_{gas} \approx 1.5 \ 10^{61} \ ergs$ $M_{gas} \approx 2.10^{12} M_{\odot}$

Predict: Fe Line at 3.2 keV Cooler spectrum away from jet Non-Thermal Interpretation $L_x=5.4 \ 10^{44} \ ergs \ s^{-1}$

Model: Electrons diffuse out of jet, into low magnetic field region, and are not in bulk relativistic motion. $n_e \approx 3 \ 10^{-8} \ cm^{-3} \approx n_e$ in jet So $1/\delta^2 \approx 3\%$ of electrons diffuse out $\tau_e \approx 10^8$ yrs, against CMB Radius of 65 kpc gives $v_d \approx 300$ km/s

Predict:

Low frequency radio emission Steeper spectrum away from jet

Summary

- **1. Detailed IC/CMB structure of a Mpc scale Jet**
 - Magnetic fields $\approx 10 \ \mu Gauss$
 - Doppler and Lorentz factors ≈ 6
 - Angle to line of sight $\leq 9^{\circ}$
 - Kinetic Flux 2×10^{45} ergs s⁻¹
 - Need a $\delta \approx 1$ model for terminal hotspot X-rays
- 2. Extended X-ray emitting "tube" surrounds jet
 - $L_x \approx 3.4 \ 10^{44} \ ergs \ s^{-1}$
 - Gas Heated by Jet?
 - Entrained material, part of jet structure?
 - Electrons diffusing out of jet?
- **3. Direct Evidence of an unseen counter jet**